The background of the slide is a vibrant cosmic scene. It features a dark, deep blue space filled with numerous small, bright stars of various colors, including blue, green, and yellow. Large, ethereal nebulae in shades of purple, orange, and red are scattered across the field, creating a sense of depth and vastness. The overall effect is that of a rich, multi-colored star field or a distant galaxy cluster.

Cosmic-ray powered FIR
from H₂ snowflakes

Mark Walker & Artem Tuntsov
(Manly Astrophysics)

Why consider solid H₂ dust?

1. Origin: cold, dense gas

- Proposed by Pfenninger & Combes (1994)
 - Close to H₂ sublimation curve
 - Would form H₂ “snowflakes”
 - Inject snowflakes into diffuse ISM
- Growing evidence for presence of dark gas
 - Gamma-ray (Greiner + 2005)
 - Microwave (Planck Collaboration 2011)

Why consider solid H₂ dust?

2. Survival

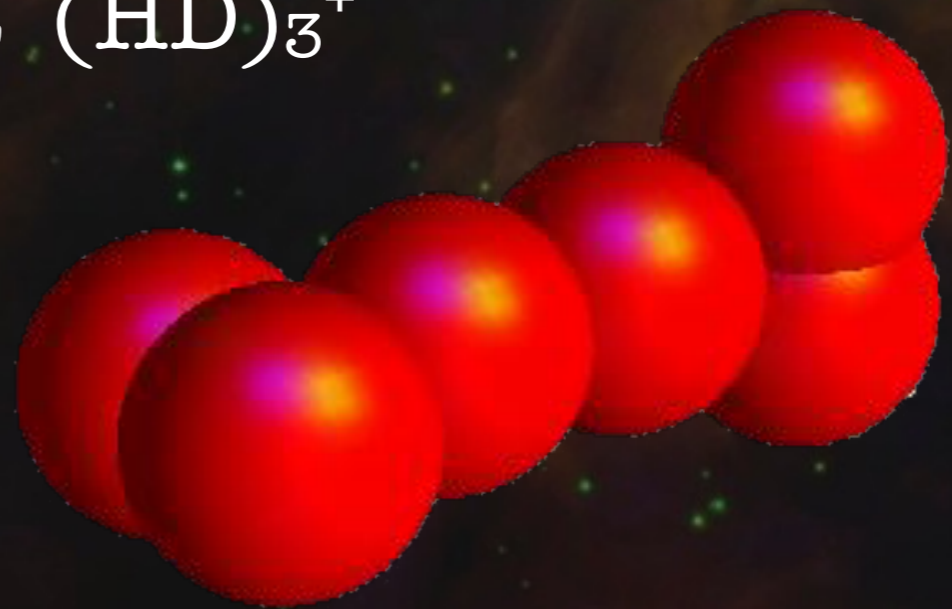
- 1968: Wickramasinghe + proposed H₂ dust
- 1969: Rapid sublimation in diffuse ISM conditions (Greenberg & de Jong; Field)
- Subsequently ignored
- 2013: Charging lowers sublimation rate (MW)
- Surface electric field → electrostatic binding
 - Strong Field → Large Effect
 - Rate lower by $\sim 10^{-85}$ @ T = 5 K
 - Snowflakes survive in diffuse ISM



Why consider solid H₂ dust?

3. Spectral Features

- Solid H₂ itself almost featureless in optical-IR
- But impurities contribute
- Ionisation chemistry differs from gas phase
 - “New” molecule : H₆⁺ (Kumada + 2005)
 - And isotopic variant (HD)₃⁺

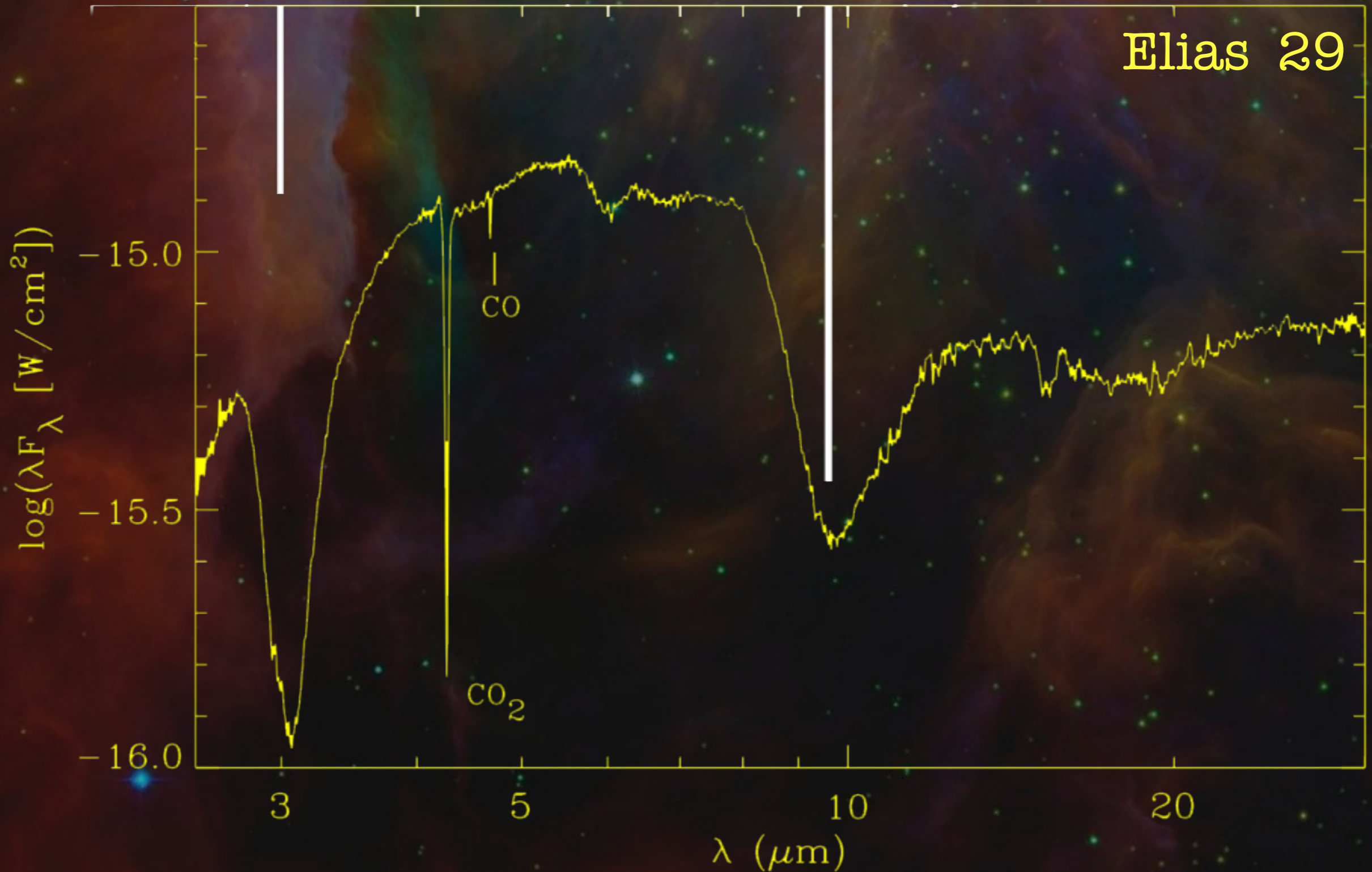


Mid-IR Absorption

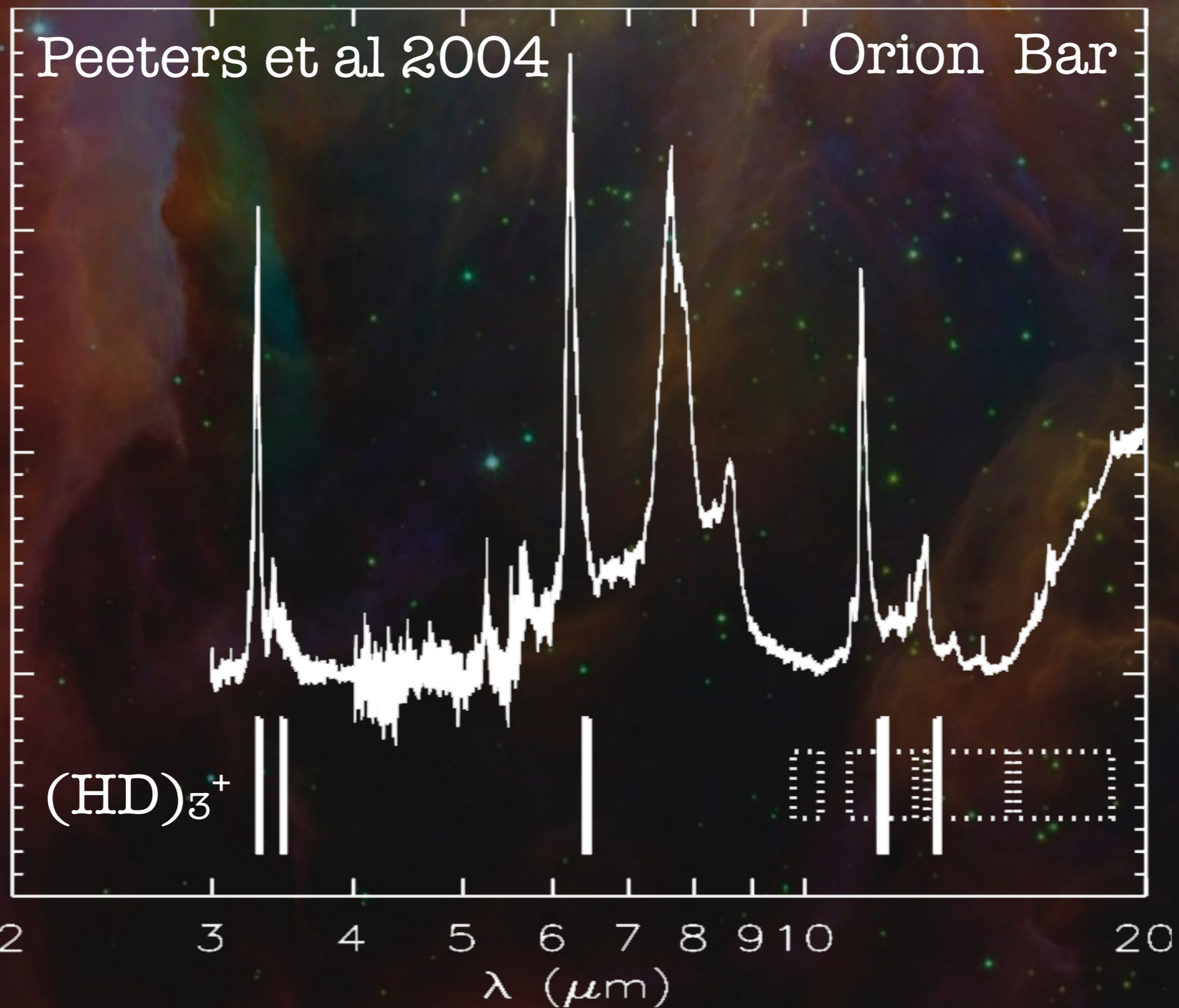
H₆⁺ (Lin + 2011)

Boogert et al 2000

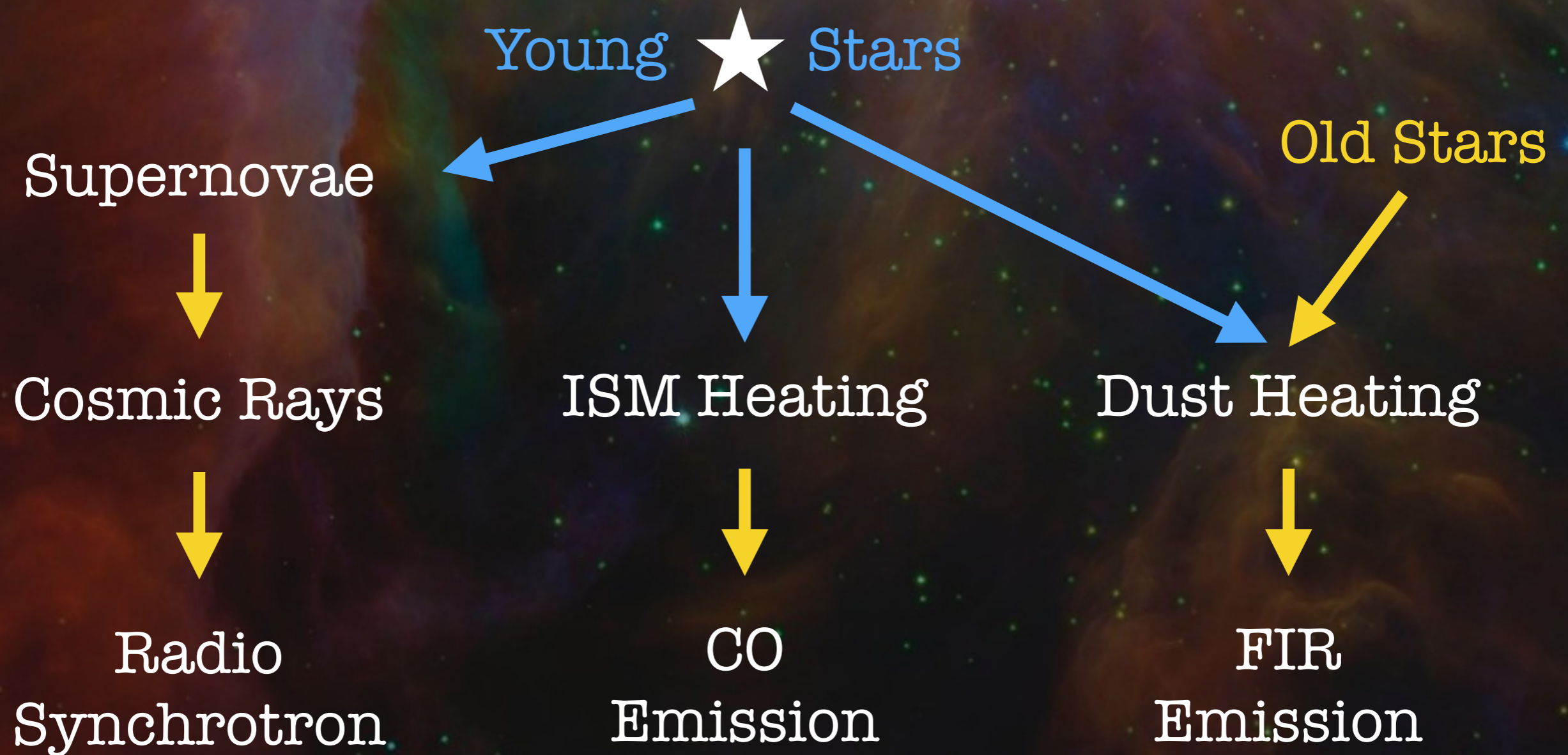
Elias 29



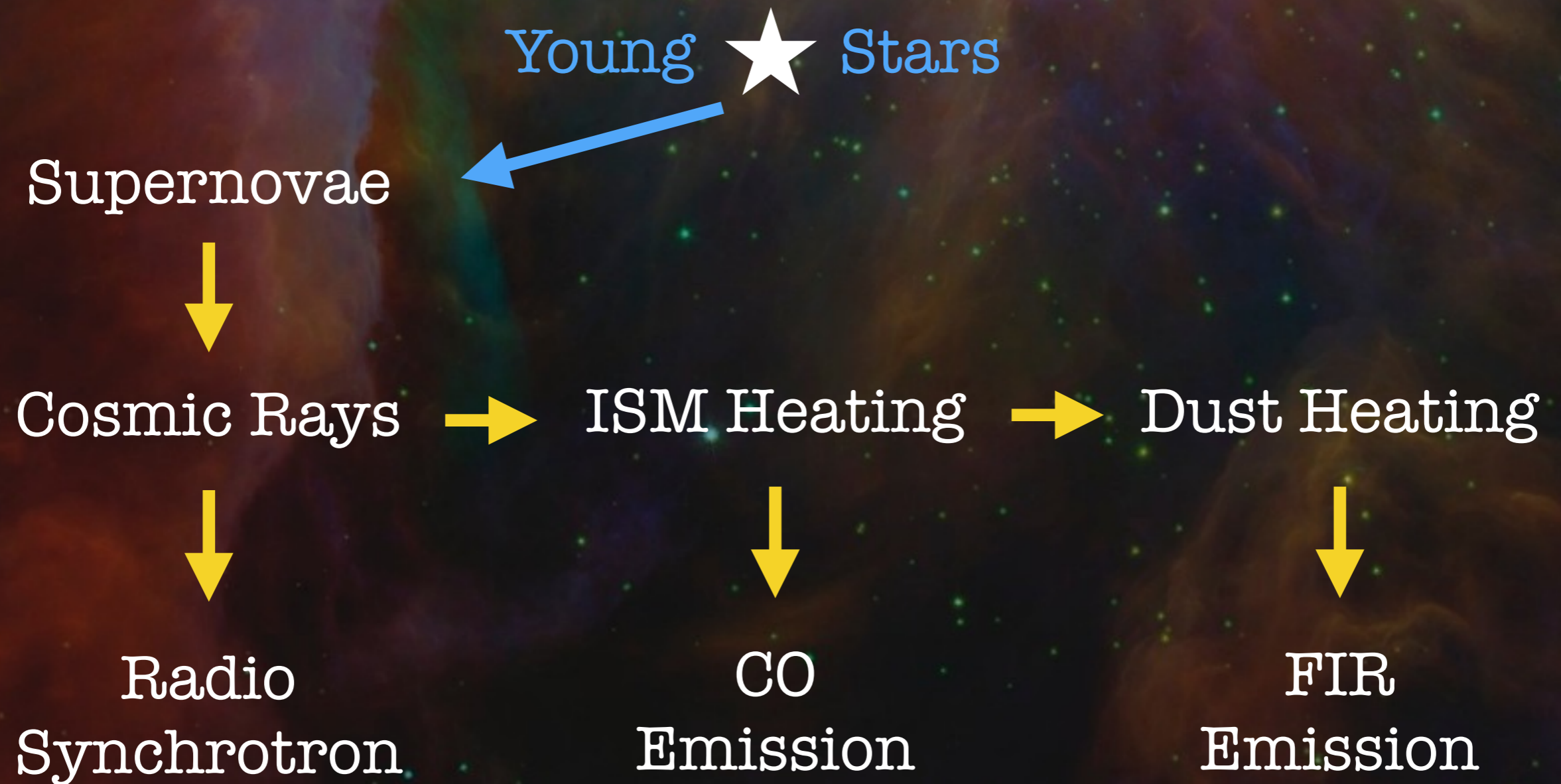
Mid-IR Emission



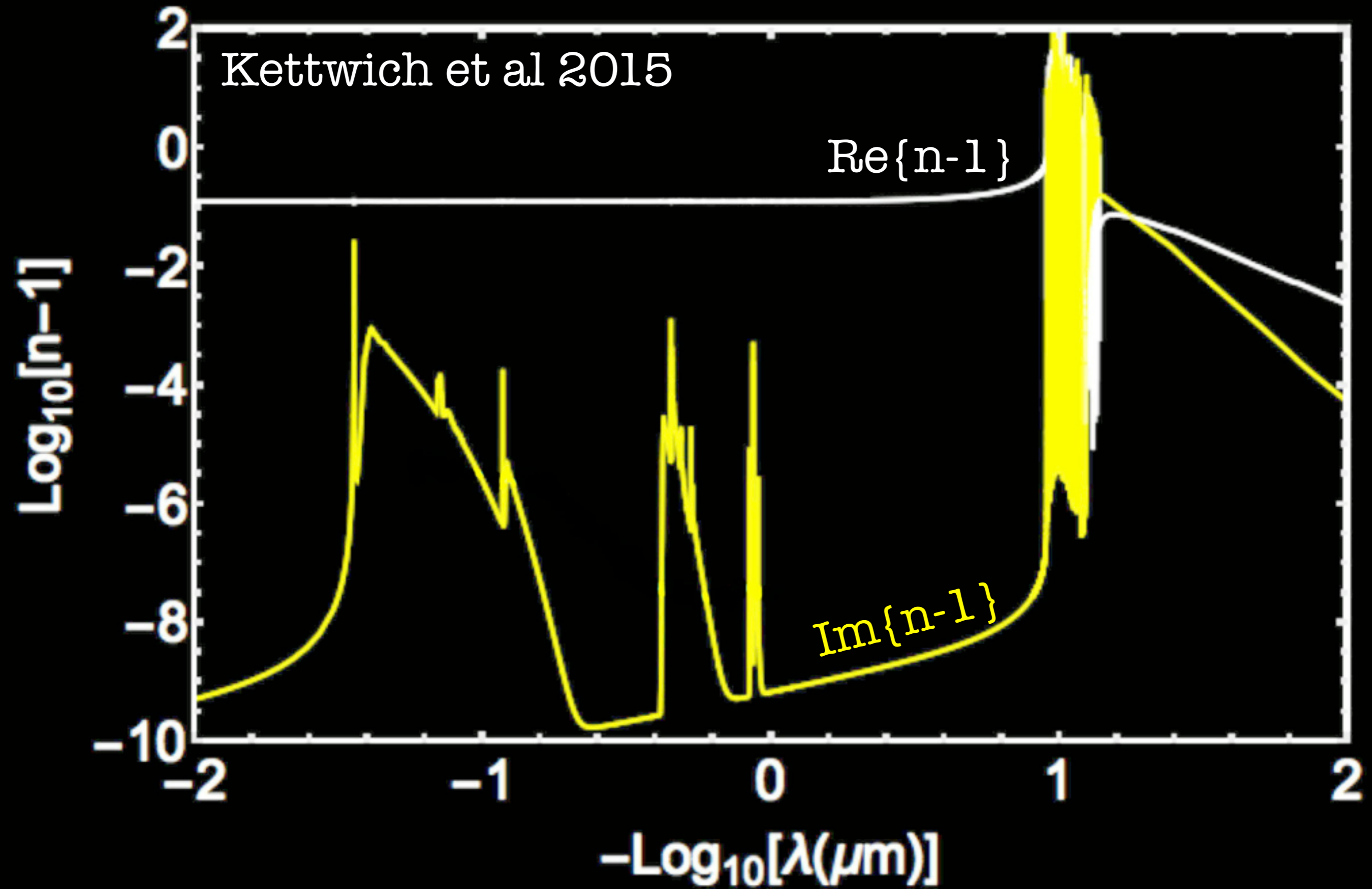
FIR & Radio: current picture



FIR & Radio: desirable picture ?



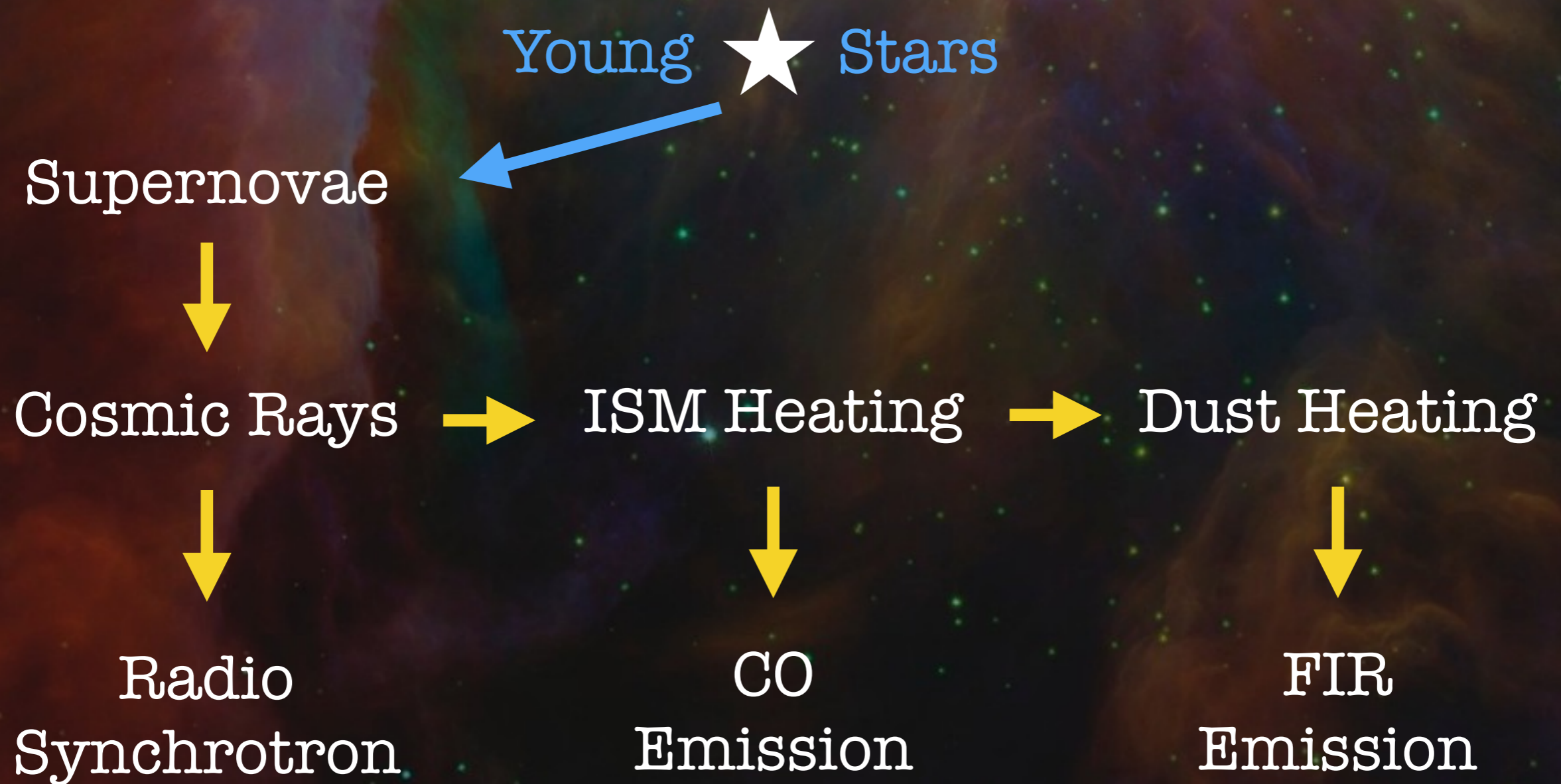
Optical constants of solid H₂



FIR emission from H₂ snowflakes

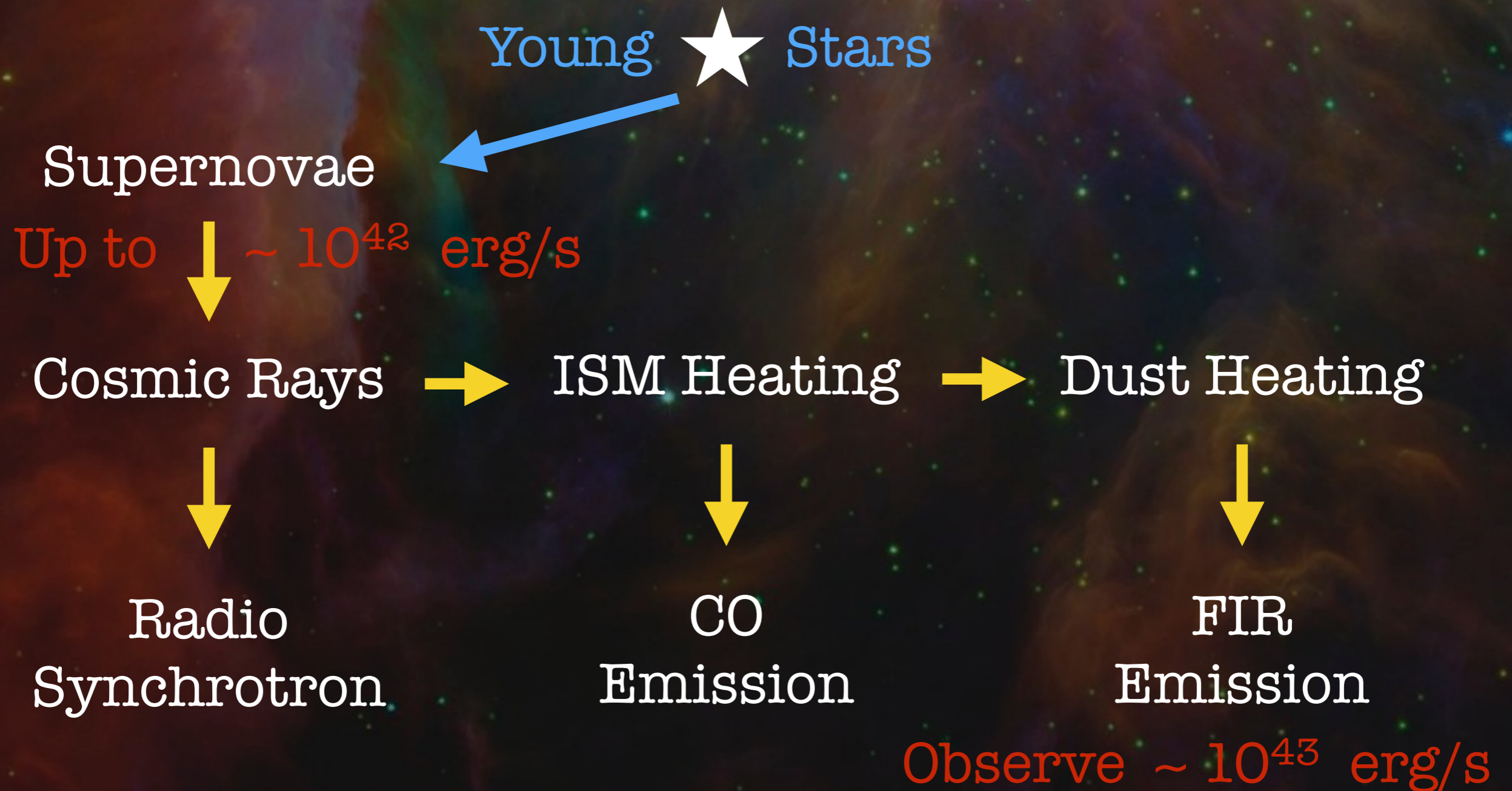
- FIR emission primarily from surface-state electrons
- But not much optical absorption (?)
- Strong heating of surface e⁻ by thermal ions in ISM
 - Distant Coulomb collisions dominate
 - Enough to explain observed Galactic FIR power
- But what heats the ions?
 - Unsolved problem (heating of WIM)
 - Cosmic-rays a possibility
 - Heating dominated by low energy particles
 - Low-energy spectrum poorly understood

FIR & Radio: snowflake picture



FIR & Radio: snowflake picture

Problem with Milky Way power budget



Summary

- Charged hydrogen snow is an alternative to silicate+graphite dust models:
 - Most abundant element, robust grains, many mid-IR bands coincident with H_6^+
 - H_2 snowflakes only scatter starlight
 - FIR emission not powered by UV/Optical
 - FIR emission from surface-state electrons
 - Heated by thermal ions in ISM
 - Ionised gas heated by cosmic-rays (?)
 - Both Radio and FIR driven by cosmic-rays
 - Good correlation likely
- Main problem is energetics: In our Galaxy SNe mechanical power < 10% FIR Power