



Status of VLBI in Italy

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Outline

1. Instruments, from yesterday to now
2. Quick science overview
3. I-VLBI, Eating VLBI, status and perspectives

I. Instruments

History

- Radio astronomy in Italy began exactly 50 years ago with the Northern Cross interferometer, built in Medicina by the University of Bologna



- Interest for VLBI grew in the 1970-80's
 - IRA was a founding member of EVN in 1980
 - accompanied by plans for construction of VLBI dishes for astronomy and geodesy



Milestones

- **1983 opening of 32m in Medicina (Mc)**
- **1988 opening of 32m in Noto (Nt)**
 - 2006 Mc connected with optical fiber
 - 2012 Nt connected with optical fiber
- 2013 Mc-Nt fringes found with DiFX correlator in Bologna
- **2013 opening of 64m in Sardinia (Sr)**
- 2014 Sr-Mc and Sr-Nt fringes found with DiFX correlator in Bologna
- 2014 Sr participates to EVN sessions



Medicina

- available receivers: 1.4, 1.6, 2.3, 5.0, 6.7, 8.4, 22 (dual feed) GHz
- frequency agility (~immediate for same focus receivers, few minutes for transition between different foci)
- 10 Gbps connection to Bologna
- Mark5-C recorder, FILA 10G board, DBBC transition completed
 - also Mark5-A+analog backend “legacy” configuration



Noto

- available receivers: 1.4-1.6, 2.3, 5.0, 6.7, 8.4, 22, **43** GHz
- active surface improving efficiency above 22 GHz (up to 86 GHz?)
- frequency agility being implemented these days
- 1 Gbps connection now, 10 Gbps in future
- Mark5-B recorder, FILA 10G board, DBBC transition completed



Sardinia

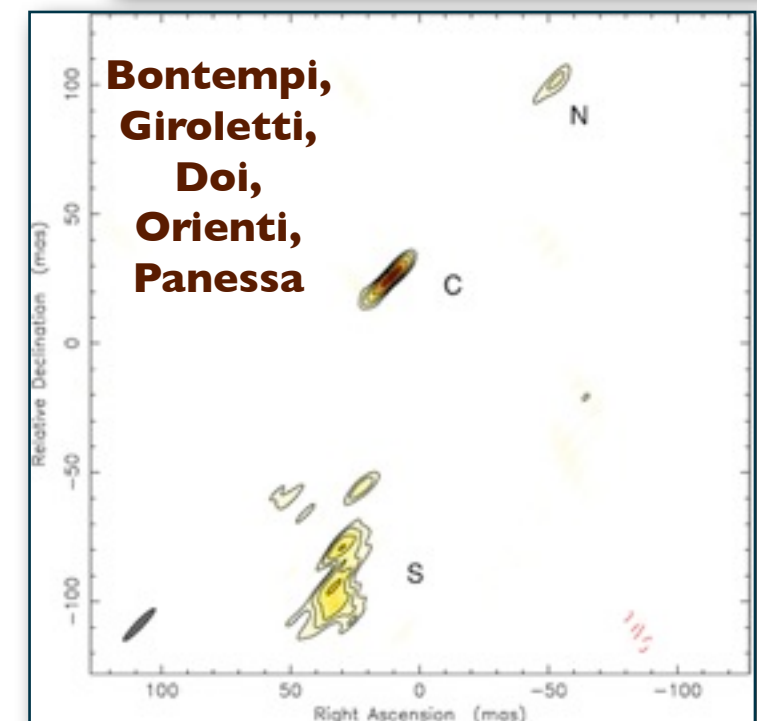
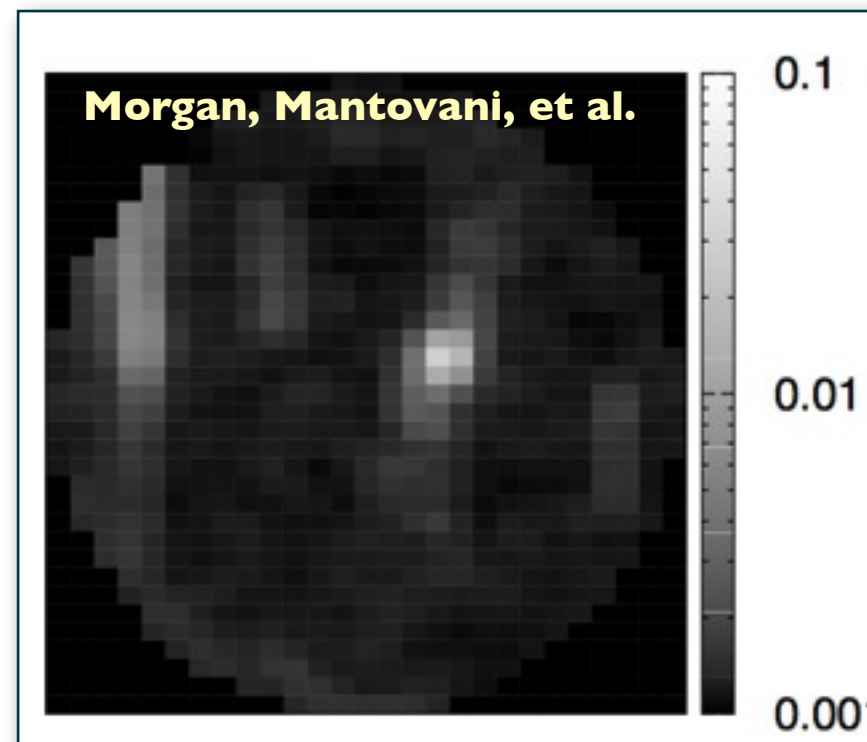
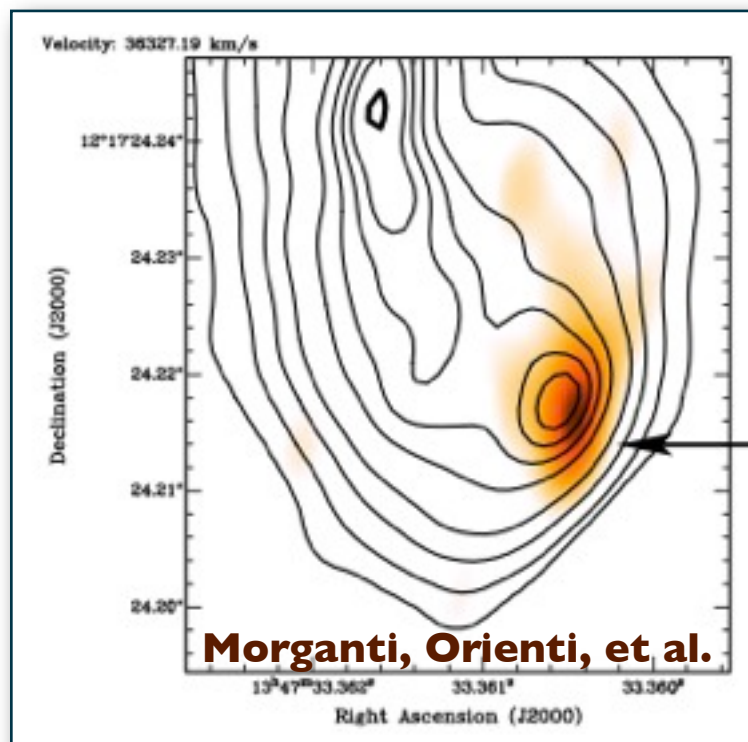
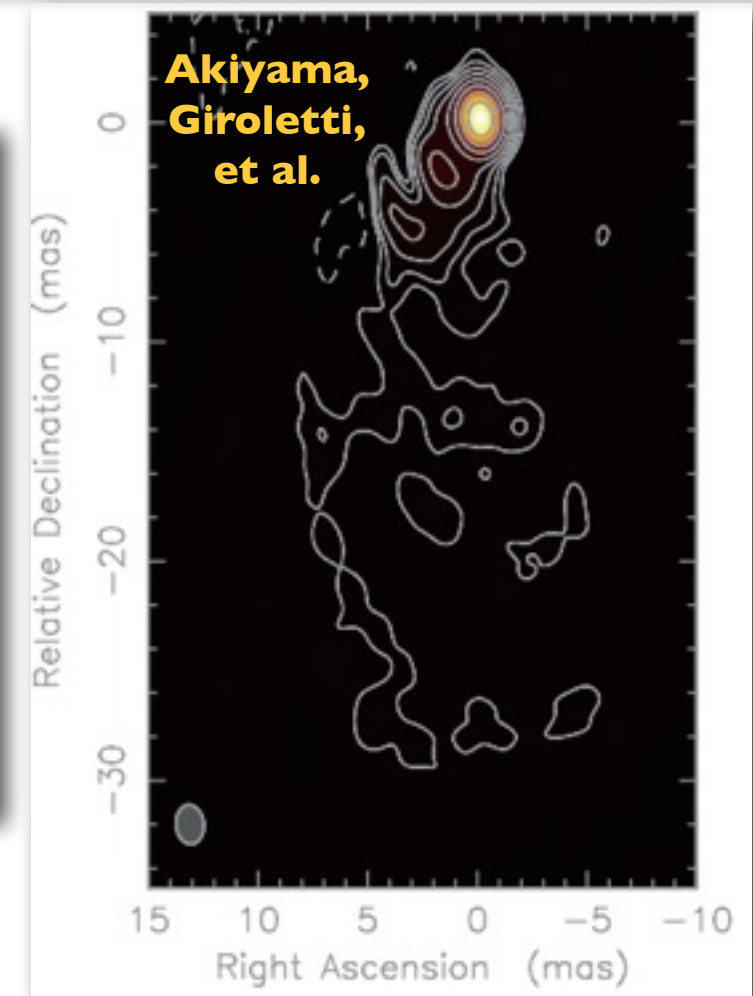
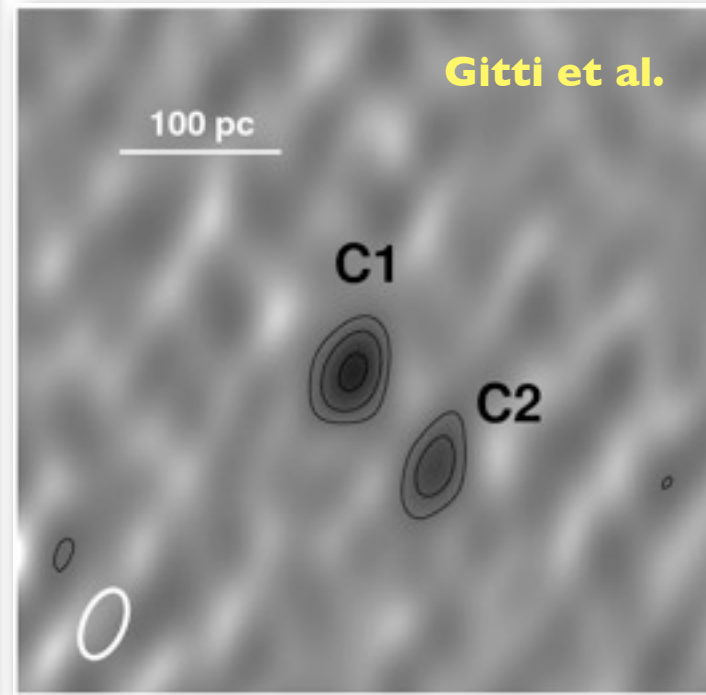
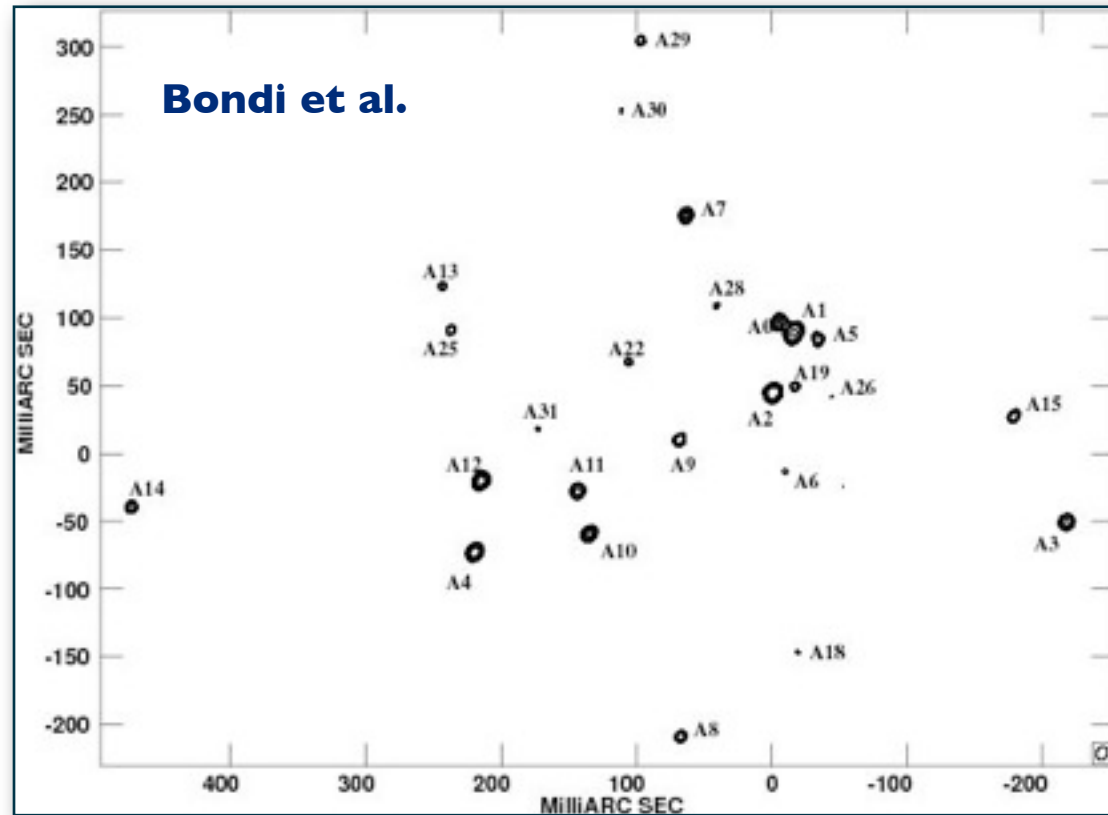
- 64 meter station
- 0.3/1.4, 6.7, 22 GHz (7-beam) receivers available
- 43 GHz receiver planned
- designed for up to 13 total receivers (three focal positions), up to 100 GHz (active surface)
- DBBC backend, FILA 10G, Mark5-C recorder
 - Mark5-B recorder being installed
- optical fiber link to be funded by regional government



2. Science overview

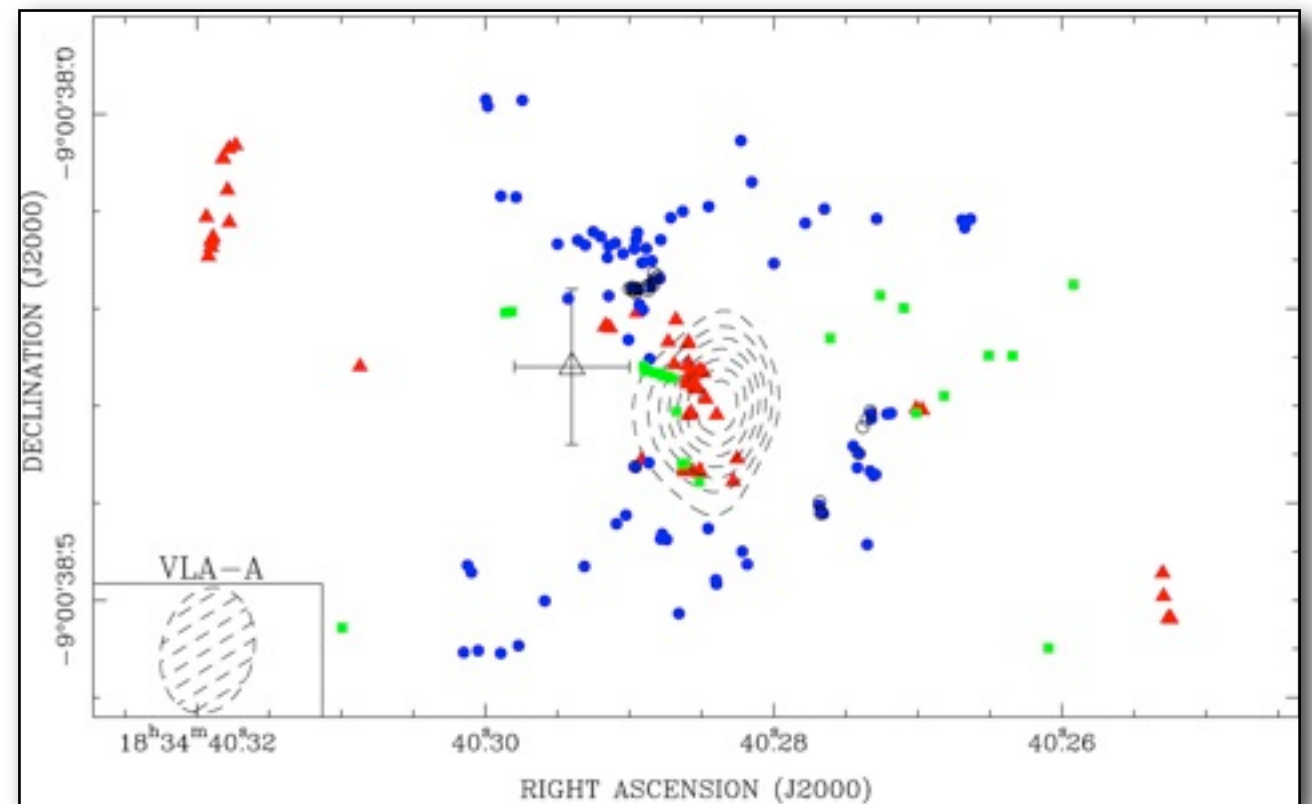
Science - at IRA

- ~10 scientific staff involved in VLBI science, mostly in Bologna
 - plus post-docs, PhD students
- blazars, relativistic jets, connection to high energy emission (*Fermi*)
- search for supermassive binary black holes
- young radio sources, CSS/GPS/HFP
- low luminosity AGN, Seyfert galaxies
- neutral hydrogen
- occasionally, galactic sources (eg novae, supernovae)
- geodesy



Science - at INAF

- Arcetri, Cagliari groups
- Stellar maser sources, giant star formation
- Collaboration with local people on blazar, LLAGNs: Rome, Milan, Bologna observatory, ...

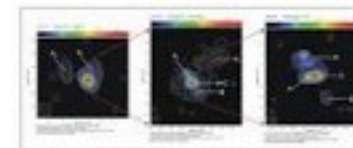


Science - International

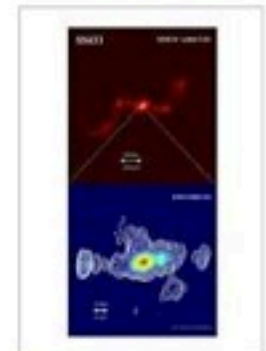
- As part of the European VLBI Network, broad range of science topics
 - AGNs
 - stellar and mega masers
 - astrometry
 - stellar evolution
 - compact objects
- Geodesy, through IVS



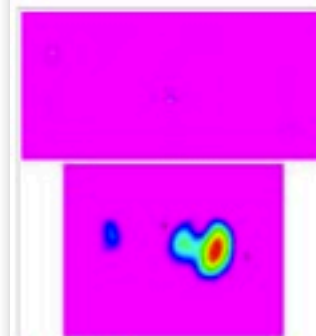
evlbi.org
AGN in Arp 299-A
1122 × 2147 - 1133k - jpg



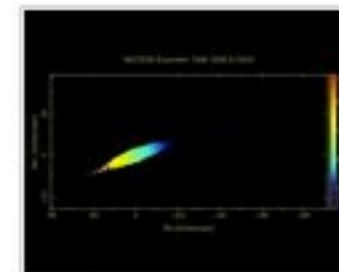
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radio structure in BALQSO
4781 × 1914 - 816k - jpg



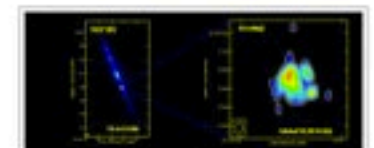
evlbi.org
VLBI Images of SS433
898 × 1298 - 79k - jpg



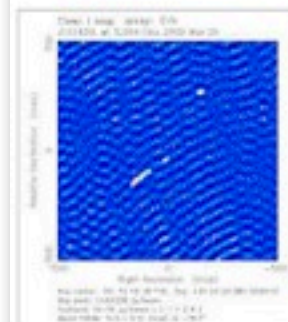
evlbi.org
Collimated Jet in Seyfert 1
741 × 797 - 45k - png



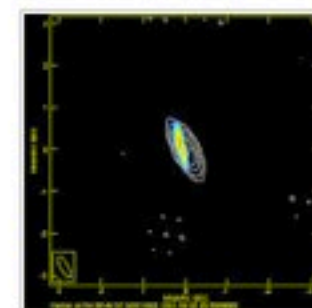
evlbi.org
seen edge-on in NGC7538
850 × 680 - 9k - gif



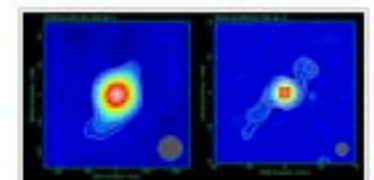
evlbi.org
A distorted radio shell in the
1461 × 613 - 47k - gif



evlbi.org
An AGN in the heart of a
533 × 606 - 141k - gif



evlbi.org
reveal relativistic jet in
533 × 526 - 10k - png



evlbi.org
in the microquasar LS 5039
1400 × 700 - 72k - gif

Activities

- **EVN:** three 3-weeks sessions, plus up to 140 hours of out-of-session observations
- **e-VLBI:** ten 24-hrs sessions per year
- **Space VLBI:** support to Radioastron KSP and GO ...not forgetting VSOP!
- **IVS**
- **single dish**
- **Italian VLBI:** so far mostly for tests - see Matteo Stagni's talk tomorrow
- **coming soon: Eating VLBI!**

Access

- EVN: three calls for proposals (Feb, Jun, Oct 1st each year)
- single dish, Italian VLBI: two calls for proposals (Apr, Oct 1st each year)
- need to provide own disks - or transfer electronically
- targets of opportunity: any time

3. I-VLBI

Italian VLBI status table

		Mc	Nt	Sr
RXs	1.4-1.6	✓	✓	✓
	5	✓	✓	✗
	6.7	✓	✓	✓
	8	✓	✓	✗
	22	✓	✓	✓
	43	✗	✓	✗
backend		DBBC	DBBC	DBBC
recorder		Mark5-C	Mark5-B	Mark5-B&C
e-VLBI		10 Gbps	1 Gbps	✗

Italian VLBI - more facts

- Baseline lengths
 - Nt-Sr 562 km
 - Sr-Mc 580 km
 - Mc-Nt 878 km
 - max angular resolution @22 GHz: 3.0 mas
- Correlator status
 - software correlator (DiFX) installed and operated in Bologna

I-VLBI status

- Two station fringes ~routinely obtained with DiFX correlator installed in Bologna
- Three station fringes not possible yet - only one failed attempt
 - limited team size and experience
 - SRT commissioning and validation
 - analog-digital transition
 - technical failures
- many of the above points shall be cleared in the future
 - more about this in Matteo Stagni's talk tomorrow

I-VLBI plans (a personal view)

- In the short term (end of 2014)
 - obtain first three-station fringes
- mid-term (6-9 months)
 - reliably carry out three-station observations
- long-term (end of 2015?)
 - operate I-VLBI as national facility

Eating VLBI

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 - and keep in mind 43 GHz receiver project for SRT

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3. 43 GHz Noto+VERA(+KVN) could be attempted
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4. real time VLBI is possible for Mc, Nt, VERA
5. maybe lower frequency observations could also be of interest?