

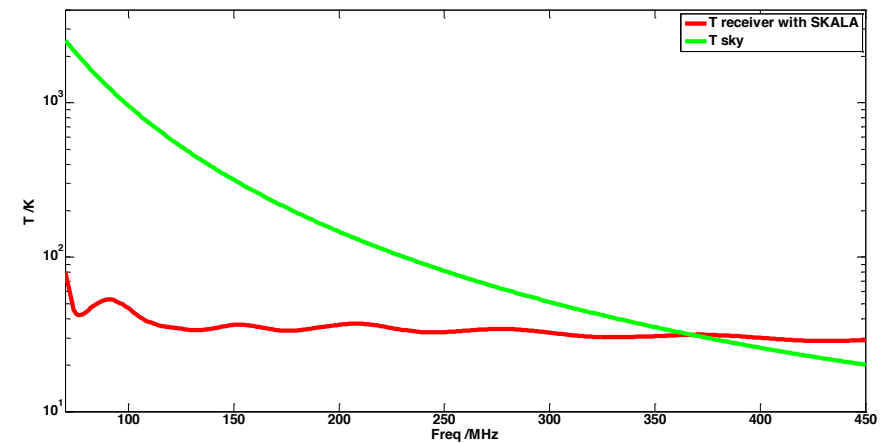
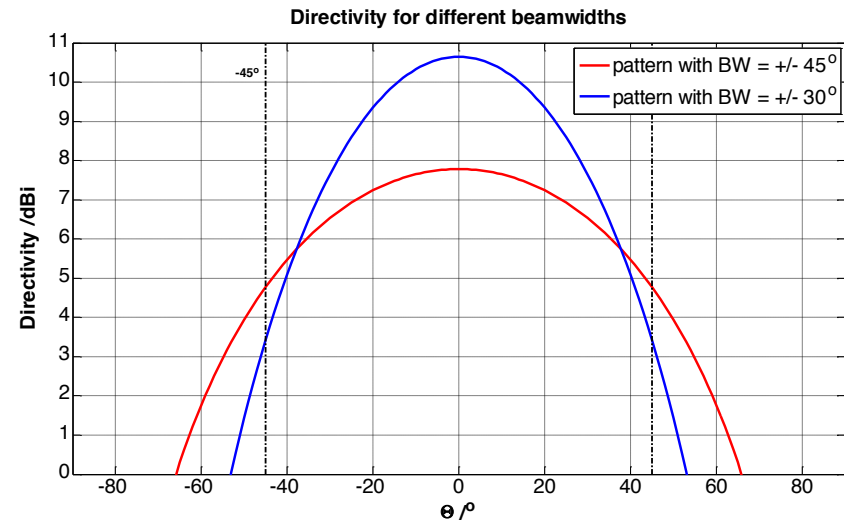
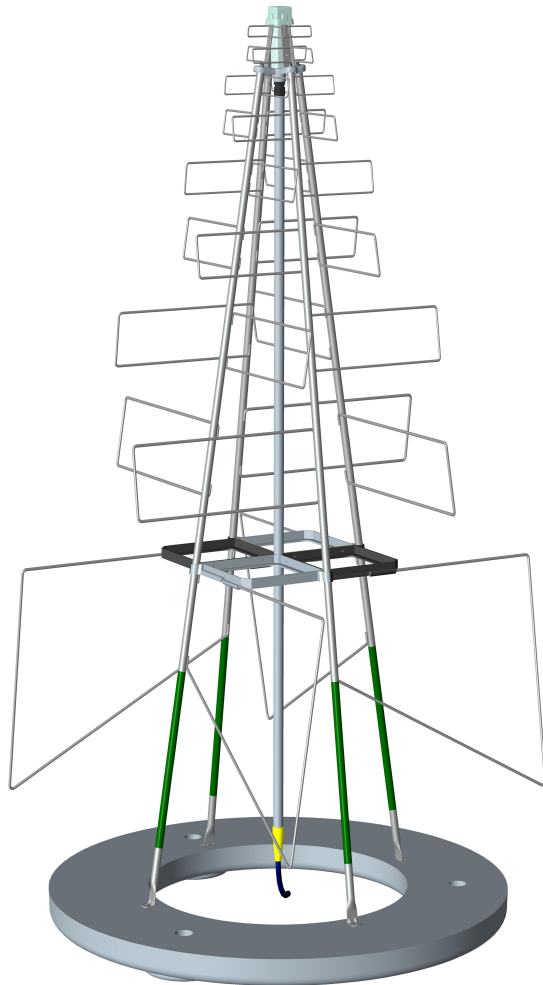
Antenna developments towards SKA1

Eloy de Lera Acedo (University of Cambridge)
email: eloy@mrao.cam.ac.uk

Overview

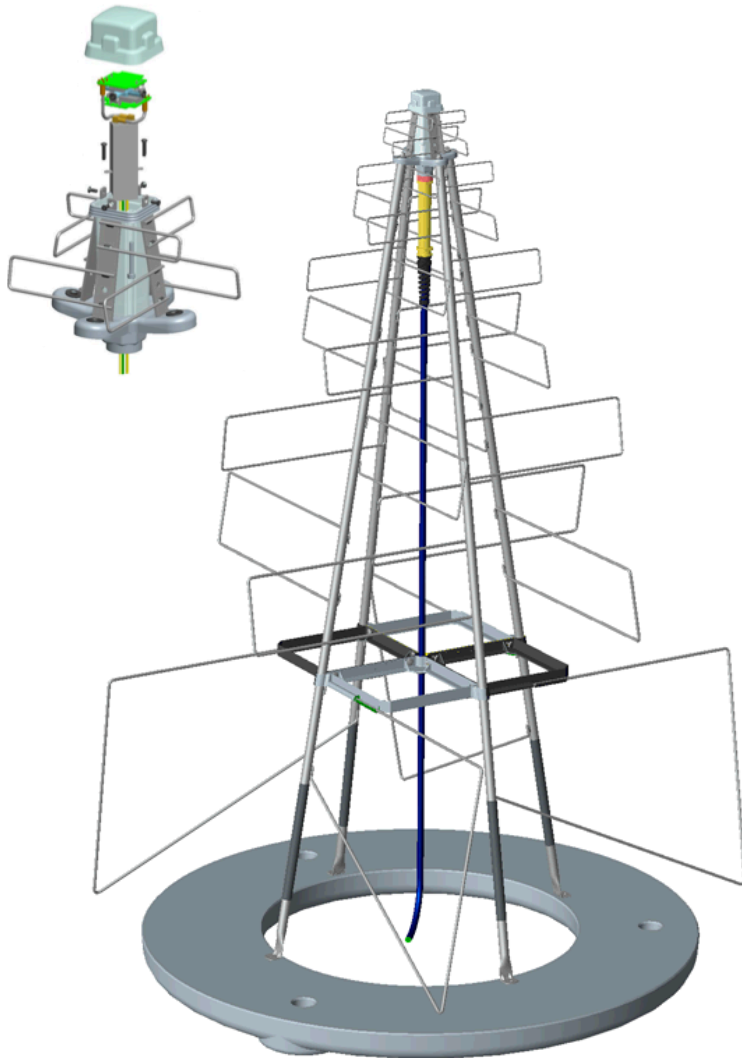
- SKALA2; AAVS1 antennas
- Frequency ripple in SKALA2
 - Current version
 - Optimization against science requirements
 - Update on the solution (SKALA3)
- Conclusions

Recap: Array element design

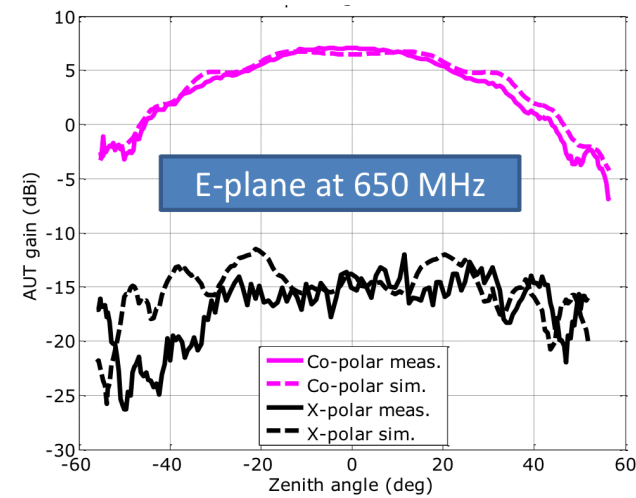
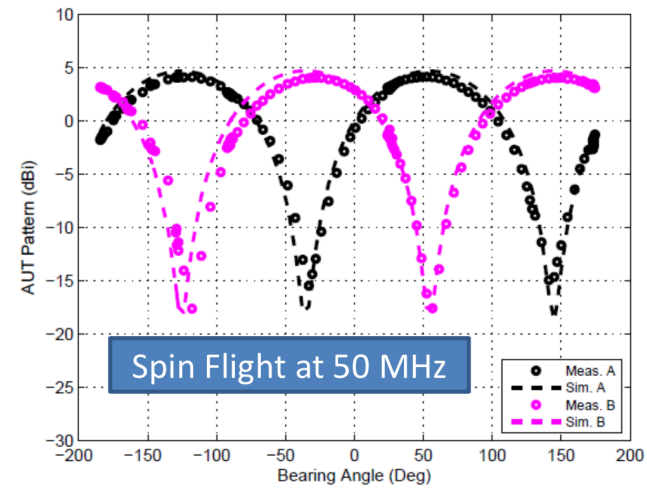


***SKALA1: EXPA 2015, de Lera et al.**

SKALA2

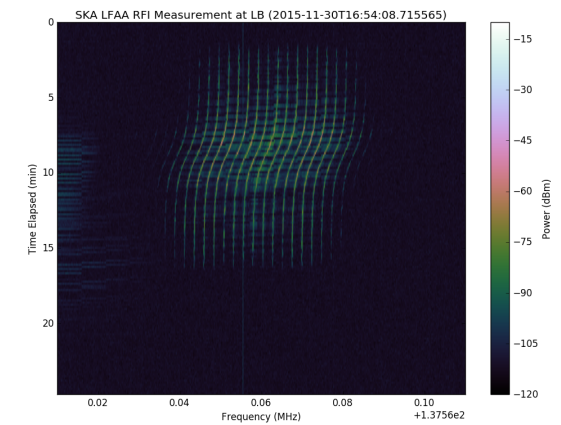
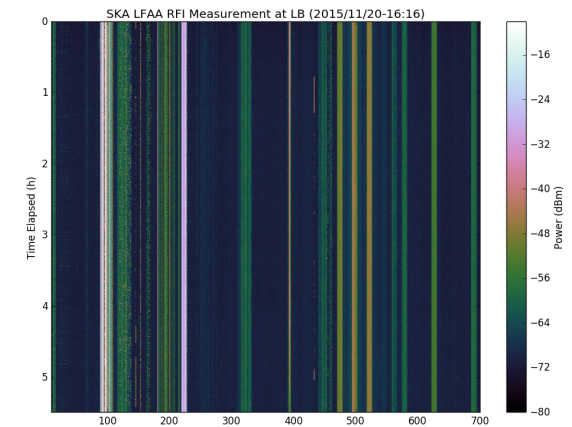


SKALA2 + LNA latest tests



SKALA2 + LNA latest tests

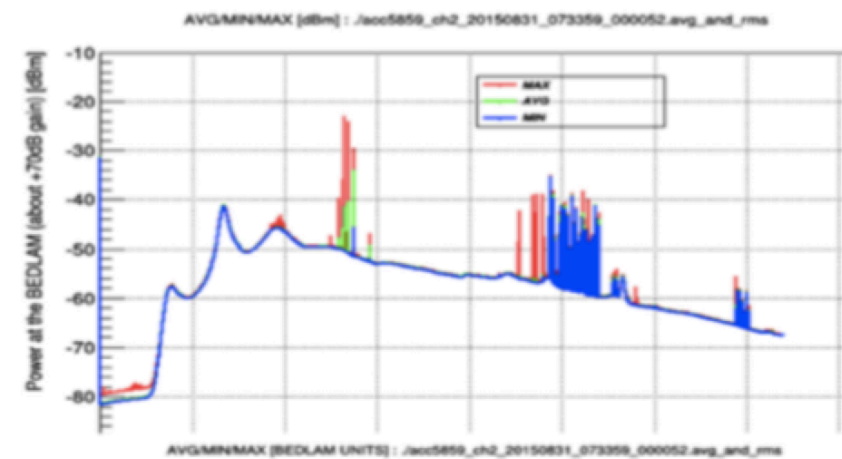
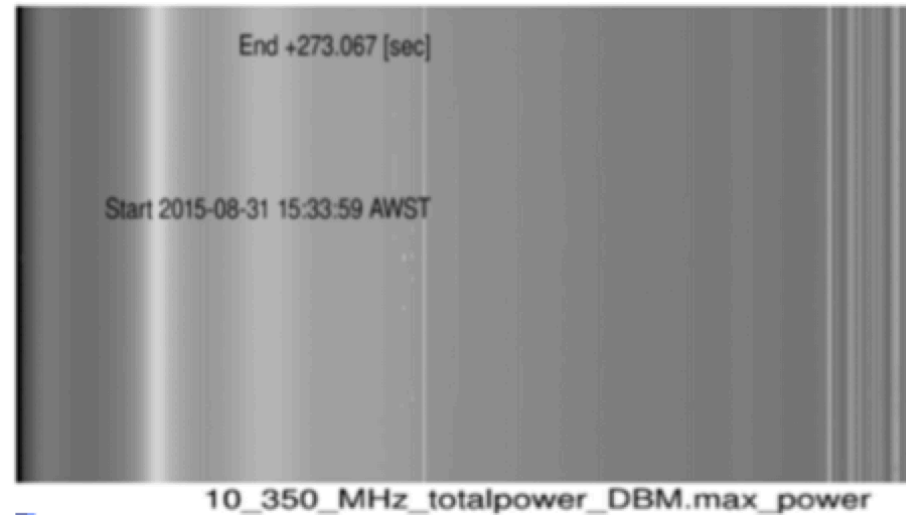
Pre-AAVS1 @ Cambridge



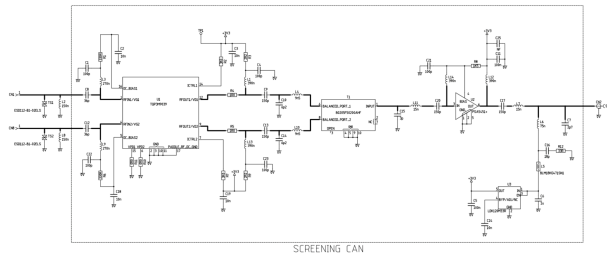
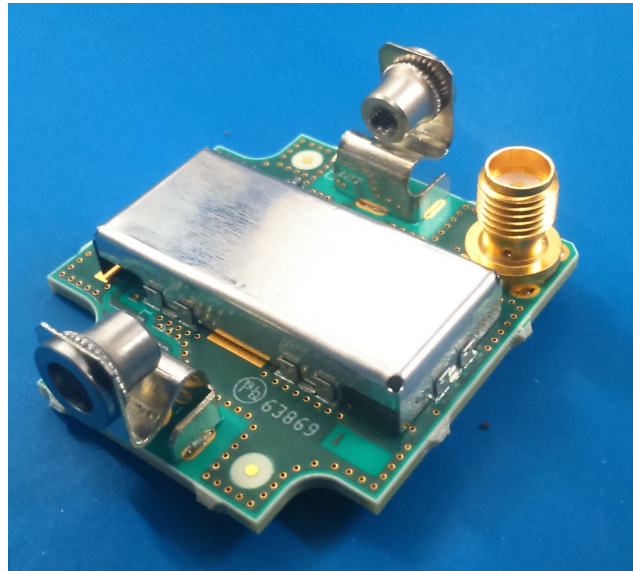
SKALA2 @ MRO



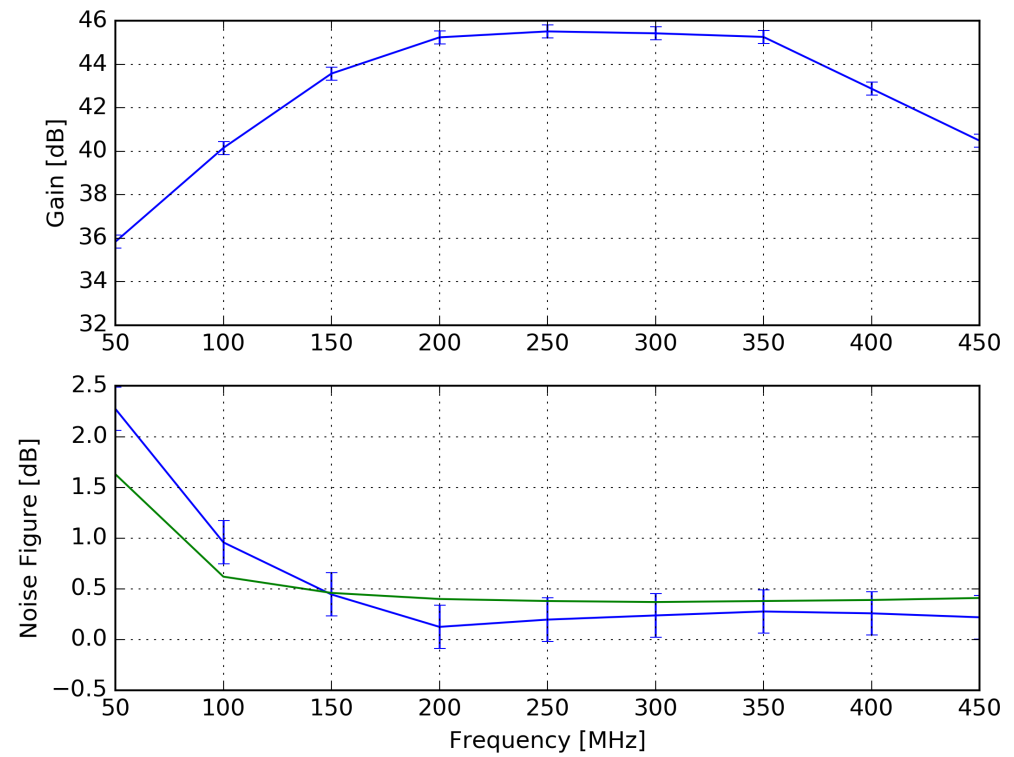
*Credit: Randall Wayth, ICRAR



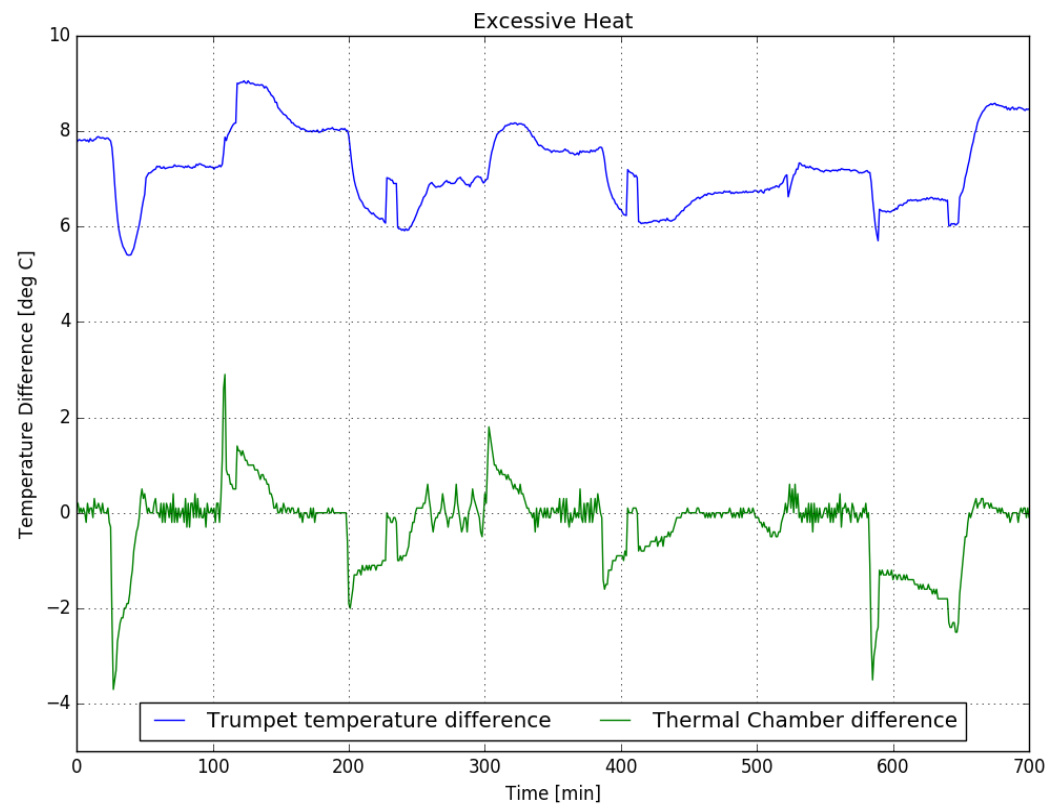
SKALA2 + LNA latest tests



NF Measurements SKALNA #076



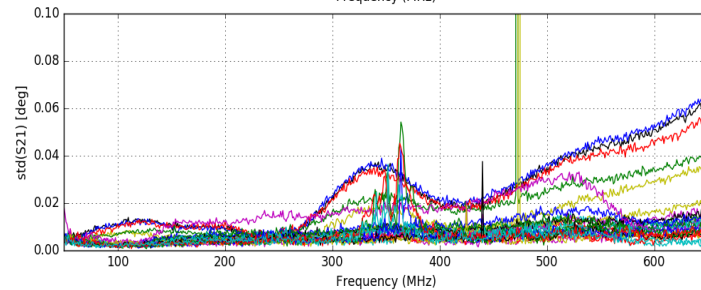
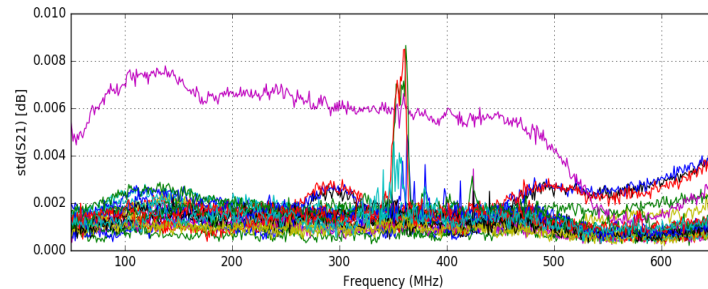
SKALA2 + LNA latest tests



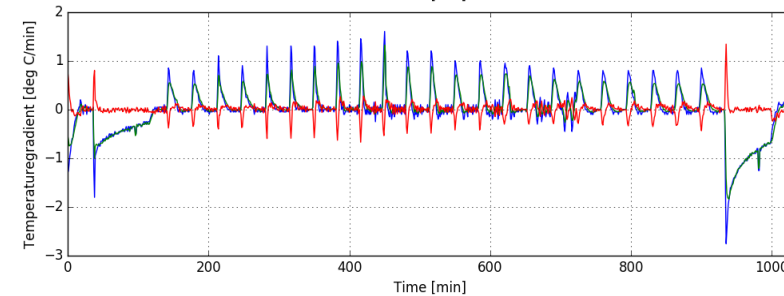
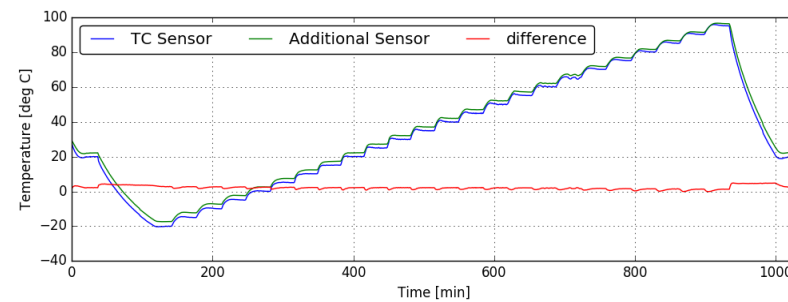
SKALA2 + LNA latest tests



Stability - stat_temp_sn076_2016-04-01 13:21:59.153208.hdf5



Temperature Profile - stat_temp_temperatures_2016-04-01 13:21:59.153208.txt



Lowest L3 requirement:
0.17 dB and 1.2°

Frequency ripple in SKALA2

LNA

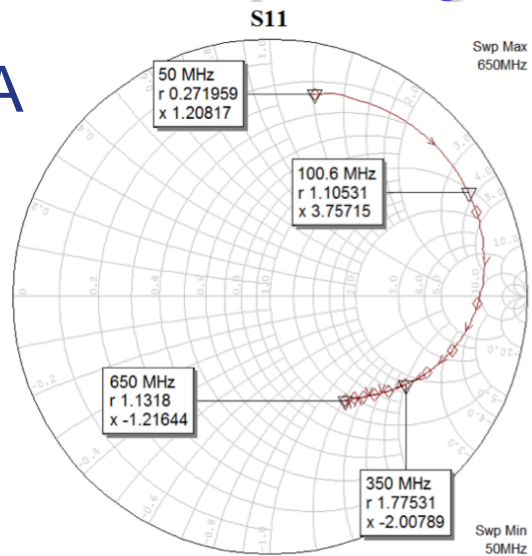
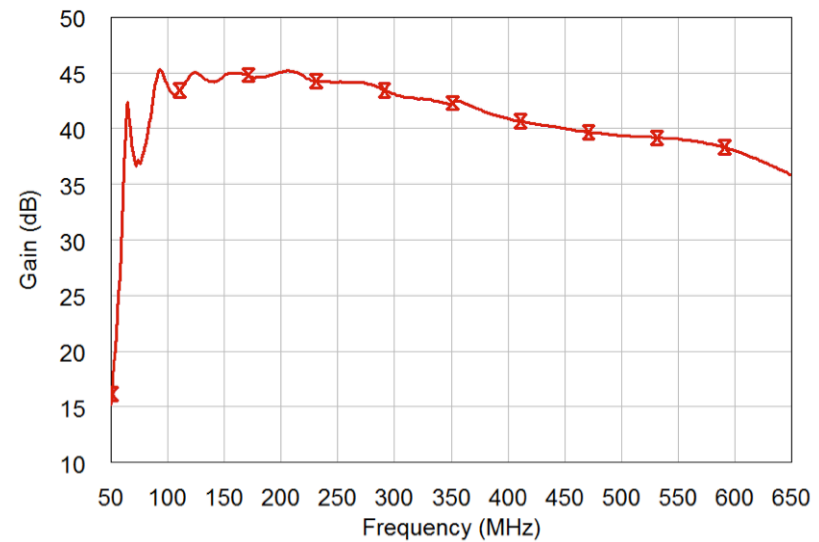


Fig. 5. LNA S_{11} in isolation including input network



Antenna

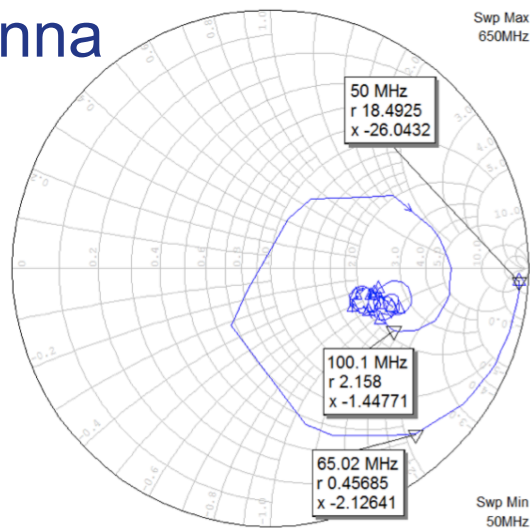
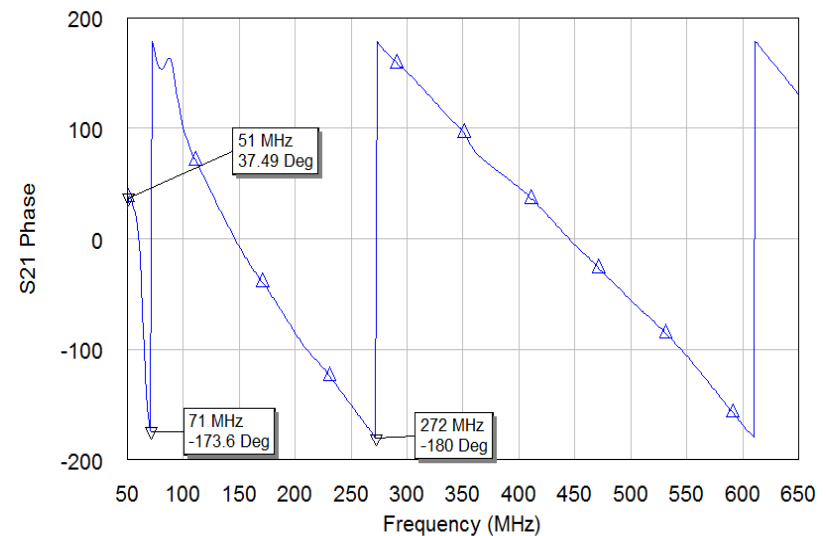
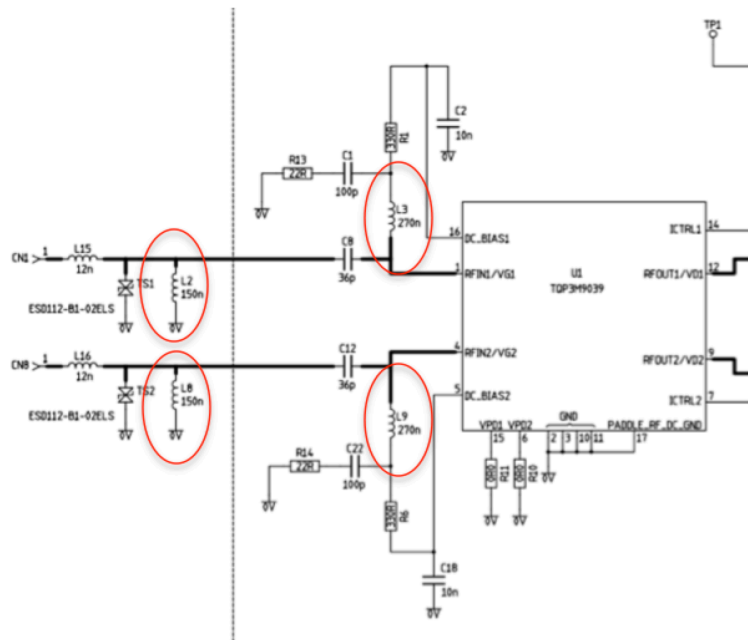


Fig. 6. Original antenna impedance as measured at Lord's Bridge

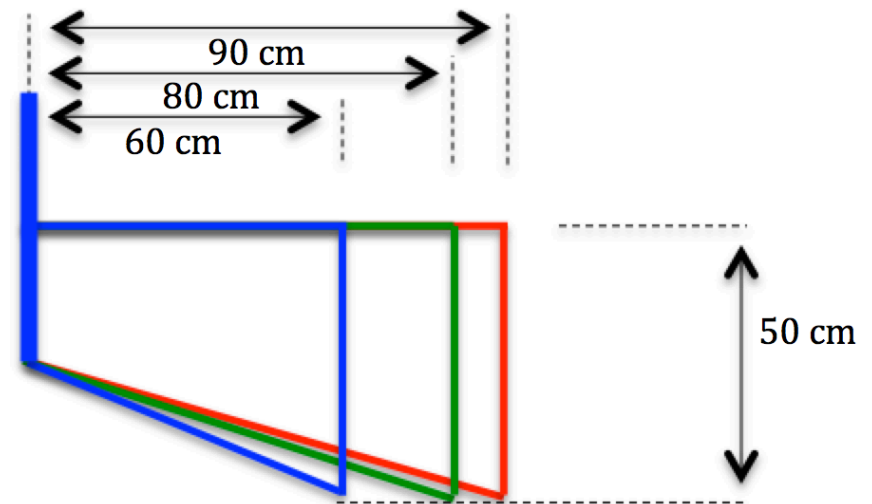


Initial optimization work

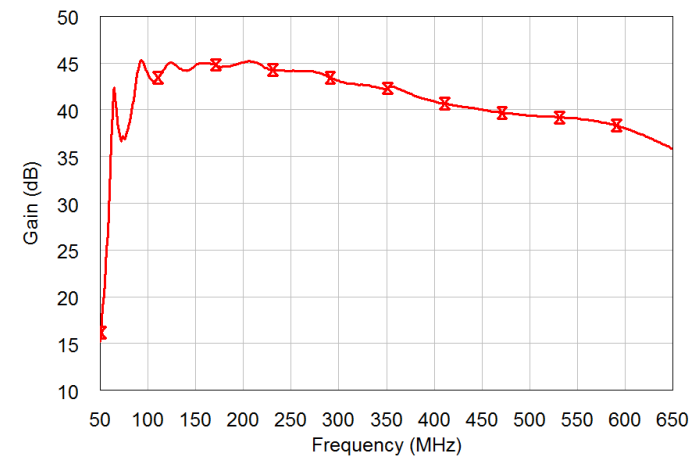
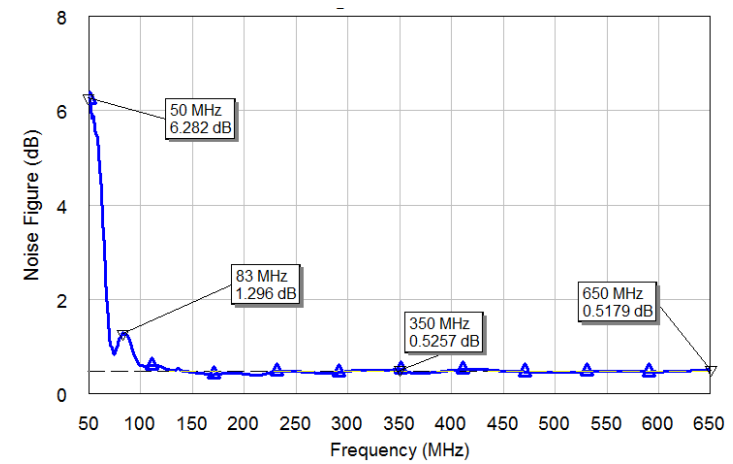
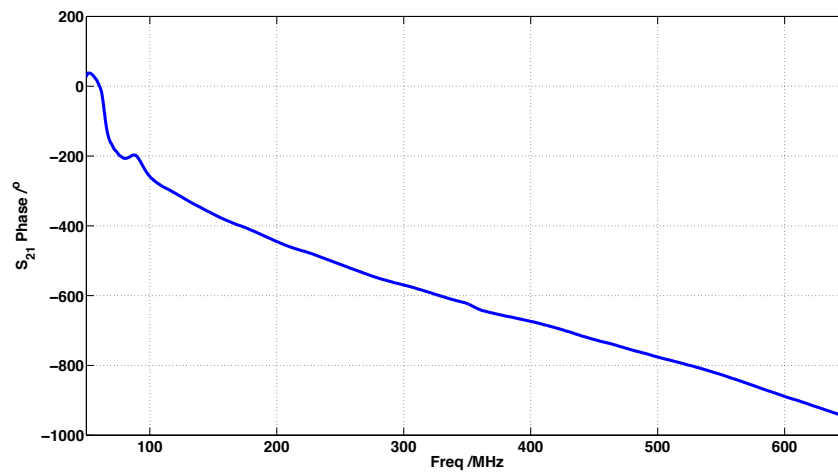
LNA



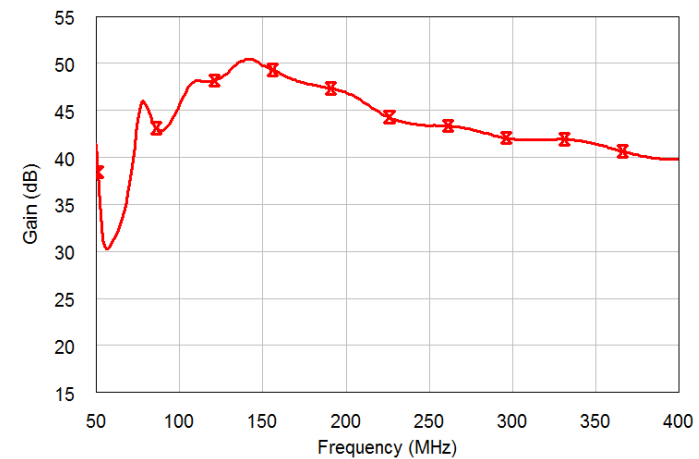
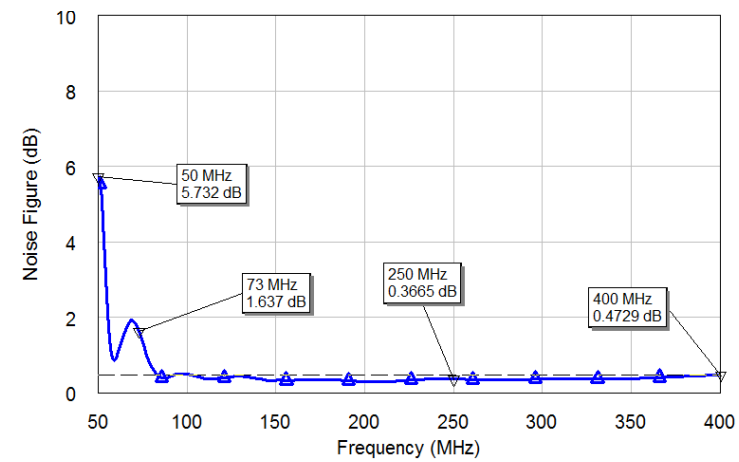
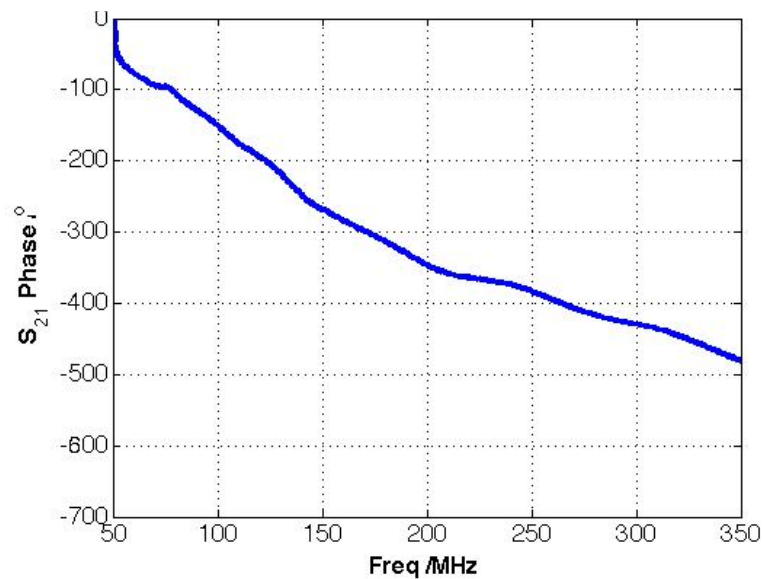
Antenna



Current



Optimized



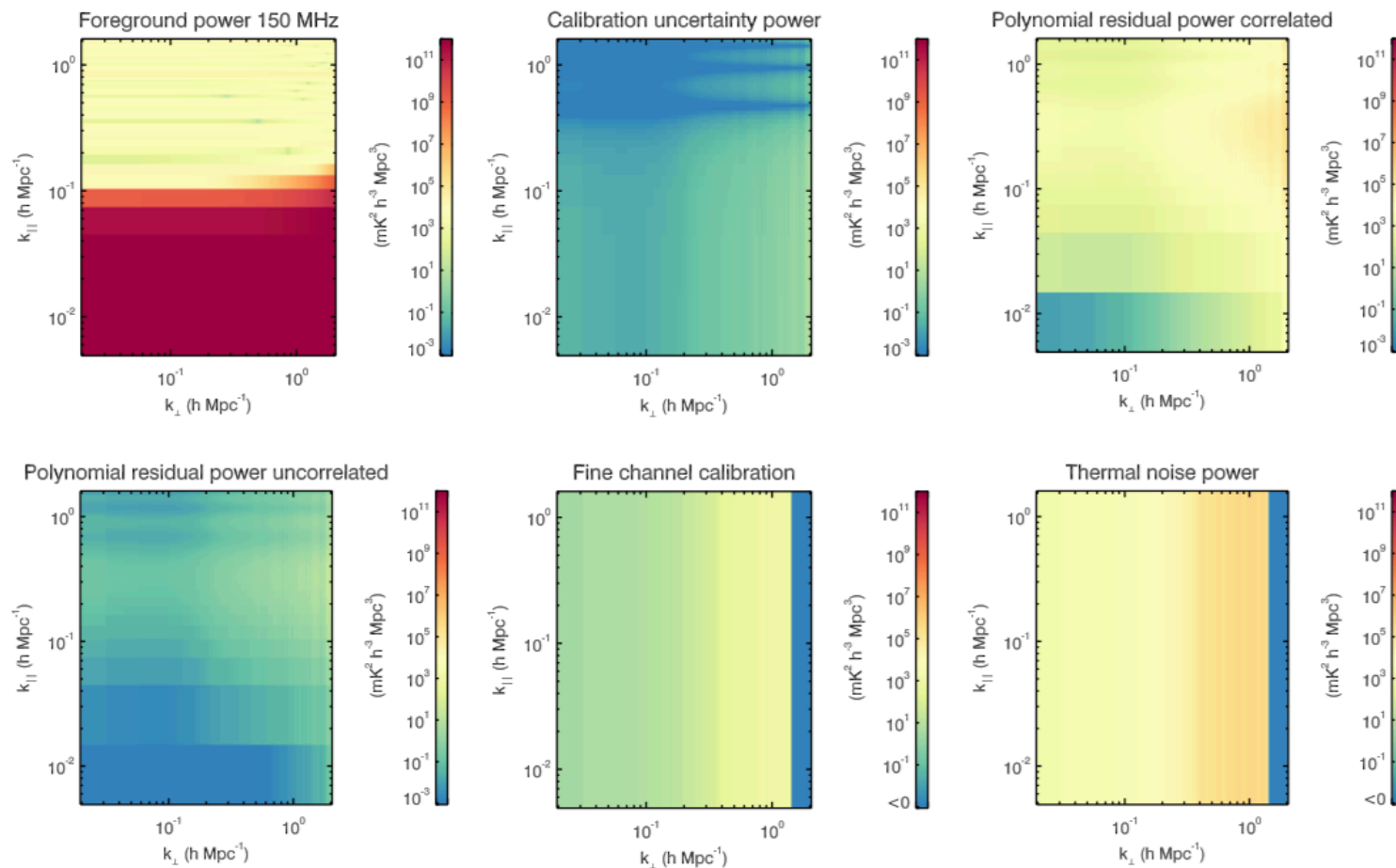
Foundation of the requirements

- No L1 requirements. Based on discussions with the SKAO, the SKA-LOW Science groups and LOFAR and MWA experts.
- Based on the “Spectral calibration requirements of radio interferometers for Epoch of Reionisation science with the SKA” (*Trott, Wayth, PASA 2016*).
- The target is that the “power introduced by deviations from the station beam calibration model will remain below the thermal (statistical) noise power.”
- Passband smoothness requirement based on fitting low order polynomials locally for the calibration (order 3 over 3 or so coarse channels).

Frequency /MHz	δ
50	0.018
100	0.01
150	0.0054
200	0.0056

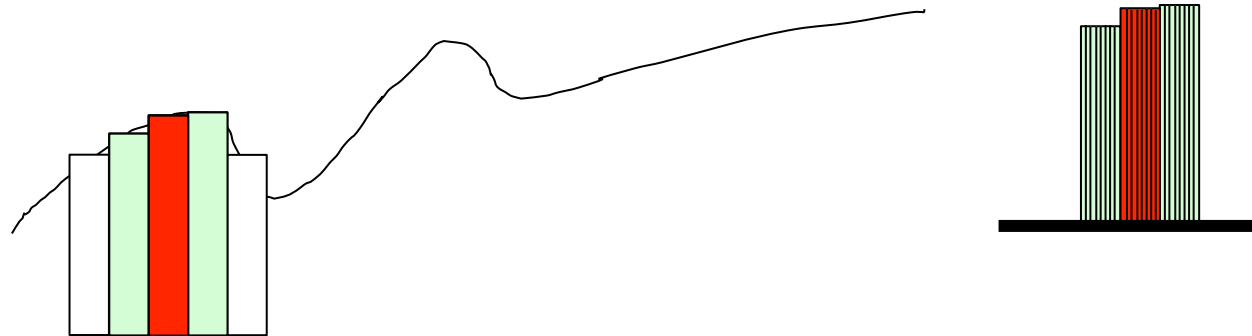
Foundation of the requirements

- *Trott, Wayth, PASA 2016.

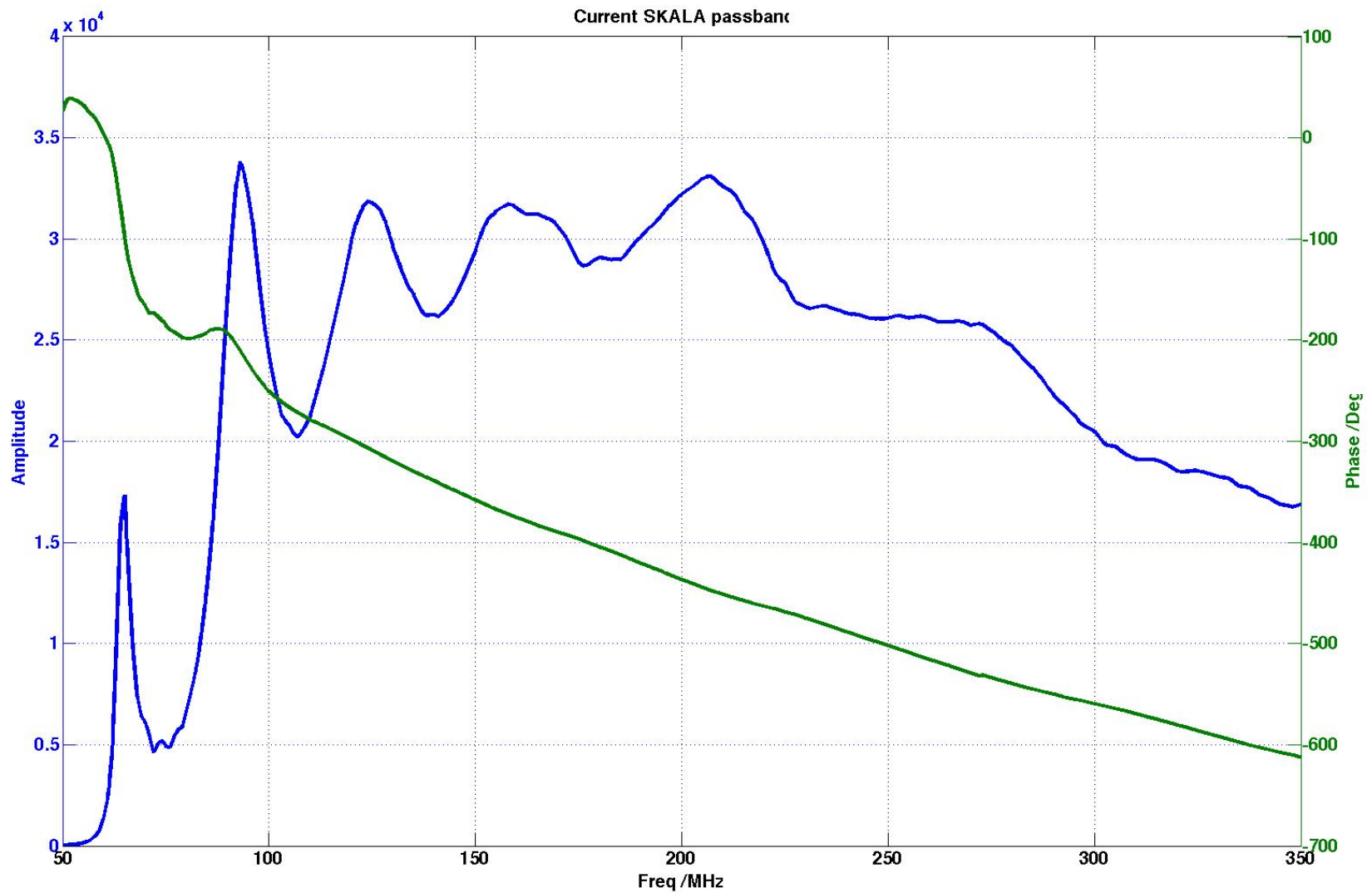


Pass-band calibration

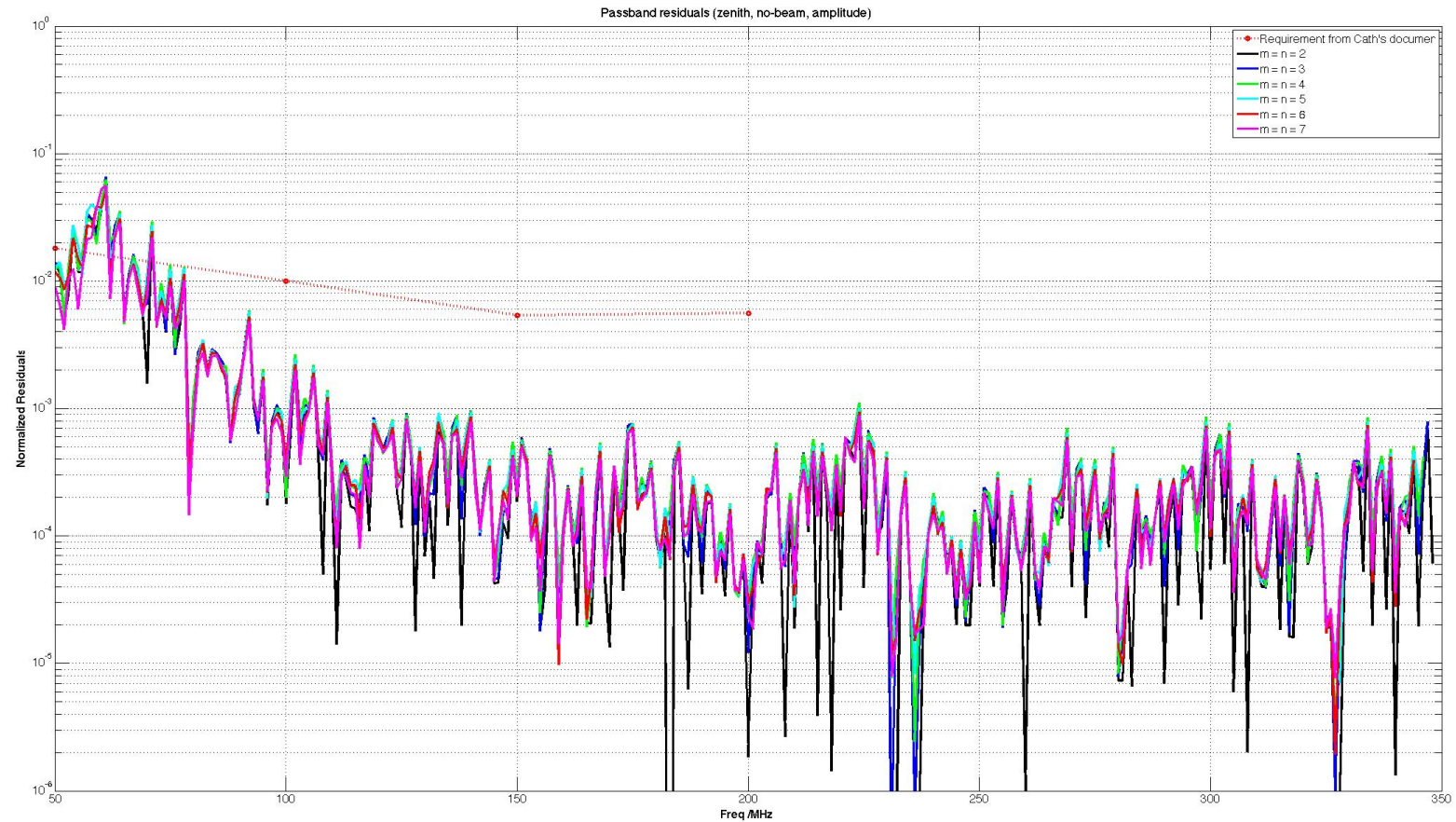
- Based on fitting a low order polynomial to a few coarse channels.
- In this presentation (similar to Comoretto 2016 and Troth 2015) the default is:
 - 3rd order polynomial
 - Using 3 coarse channels (1 channel = 750 KHz) for the fitting of the middle one.
 - 196 fine channels per coarse channel.



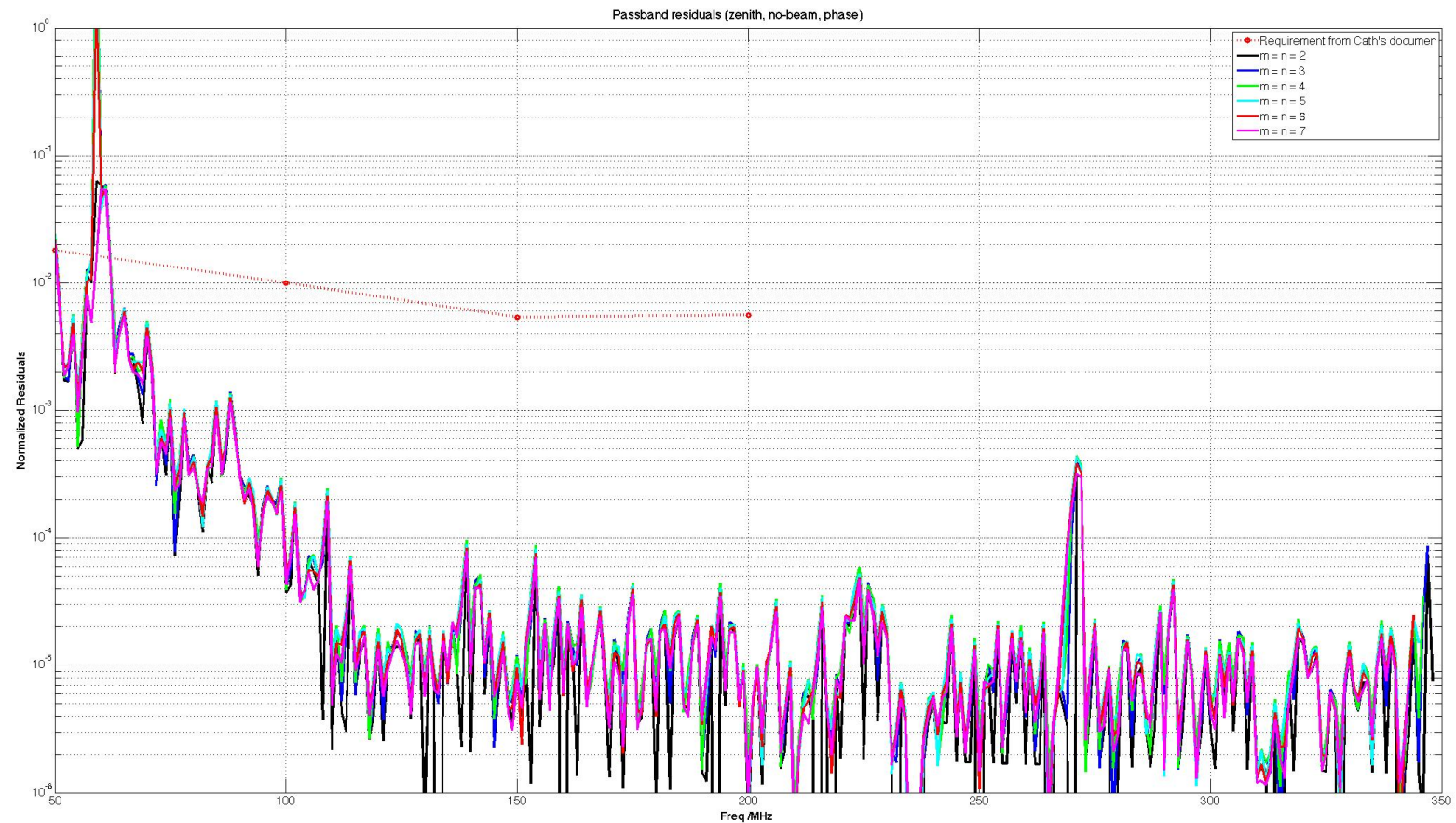
Current analysis of the residuals



Current analysis of the residuals (amplitude)

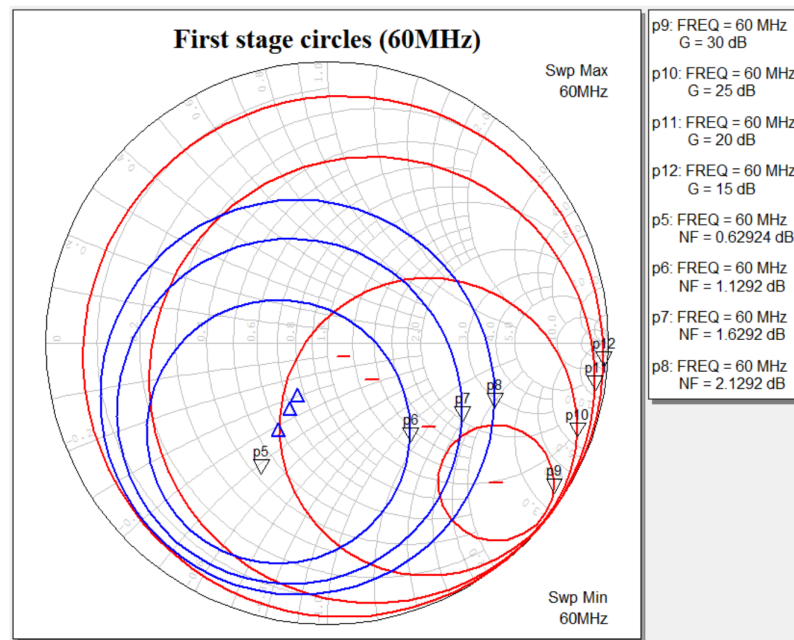


Current analysis of the residuals (phase)

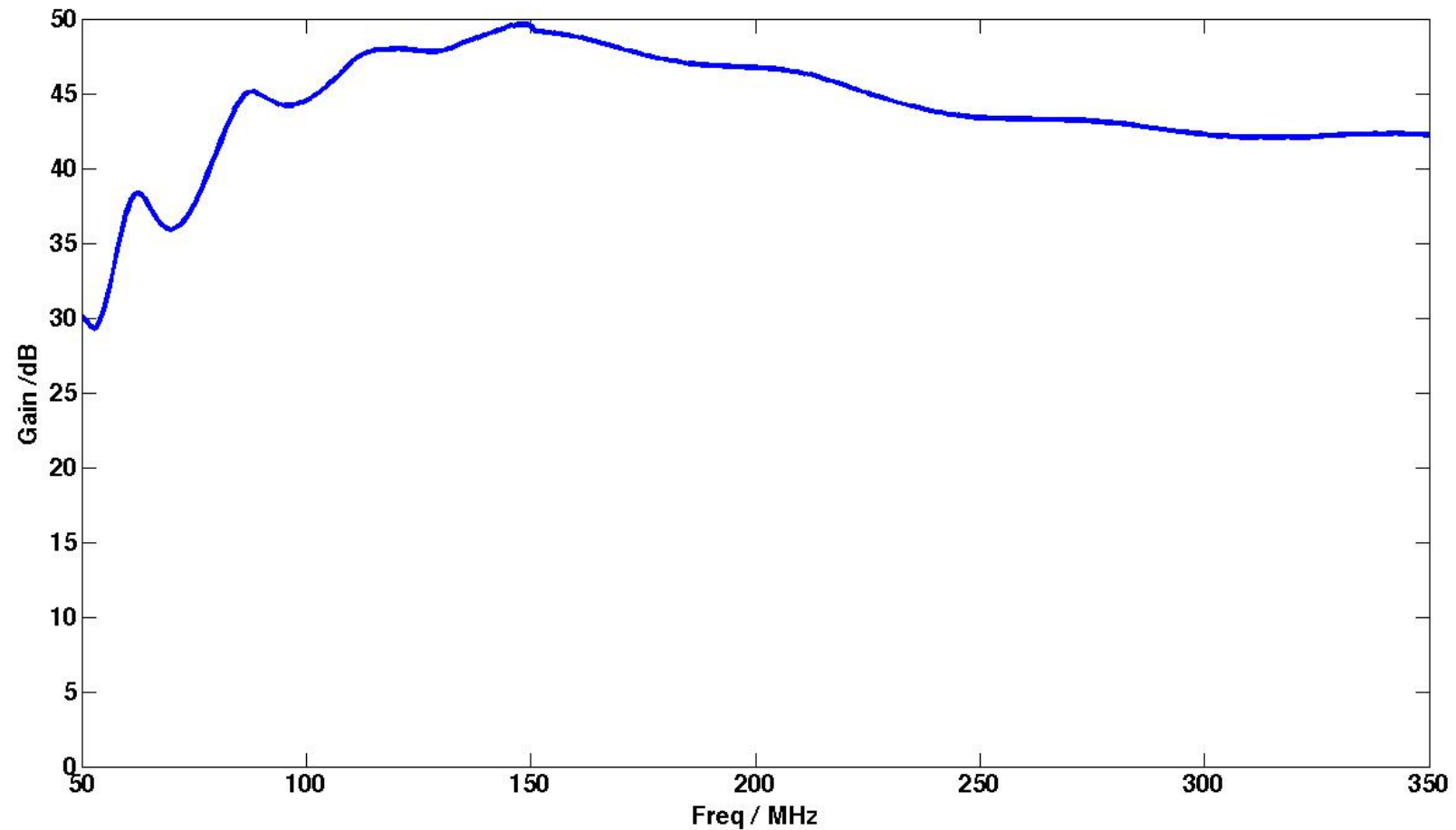


Update on the solution (SKALA3)

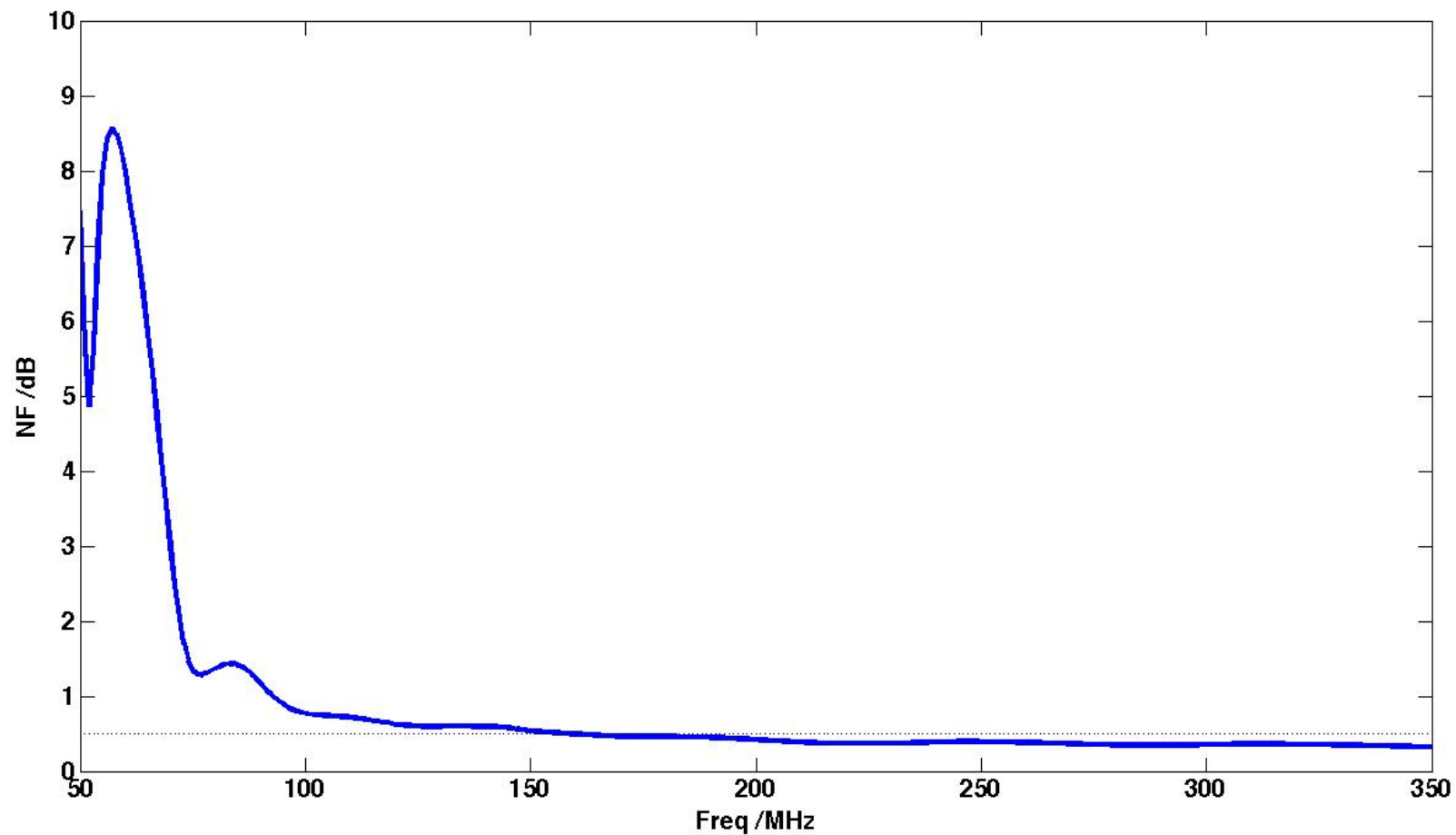
- Increased impedance of antenna (larger bottom arm)
 - 1.6x1.6 m footprint
 - Incidentally: better polarization response, better slow-ripple across band.
- Tuned input network of LNA (reduced series capacitance – 24 pF)



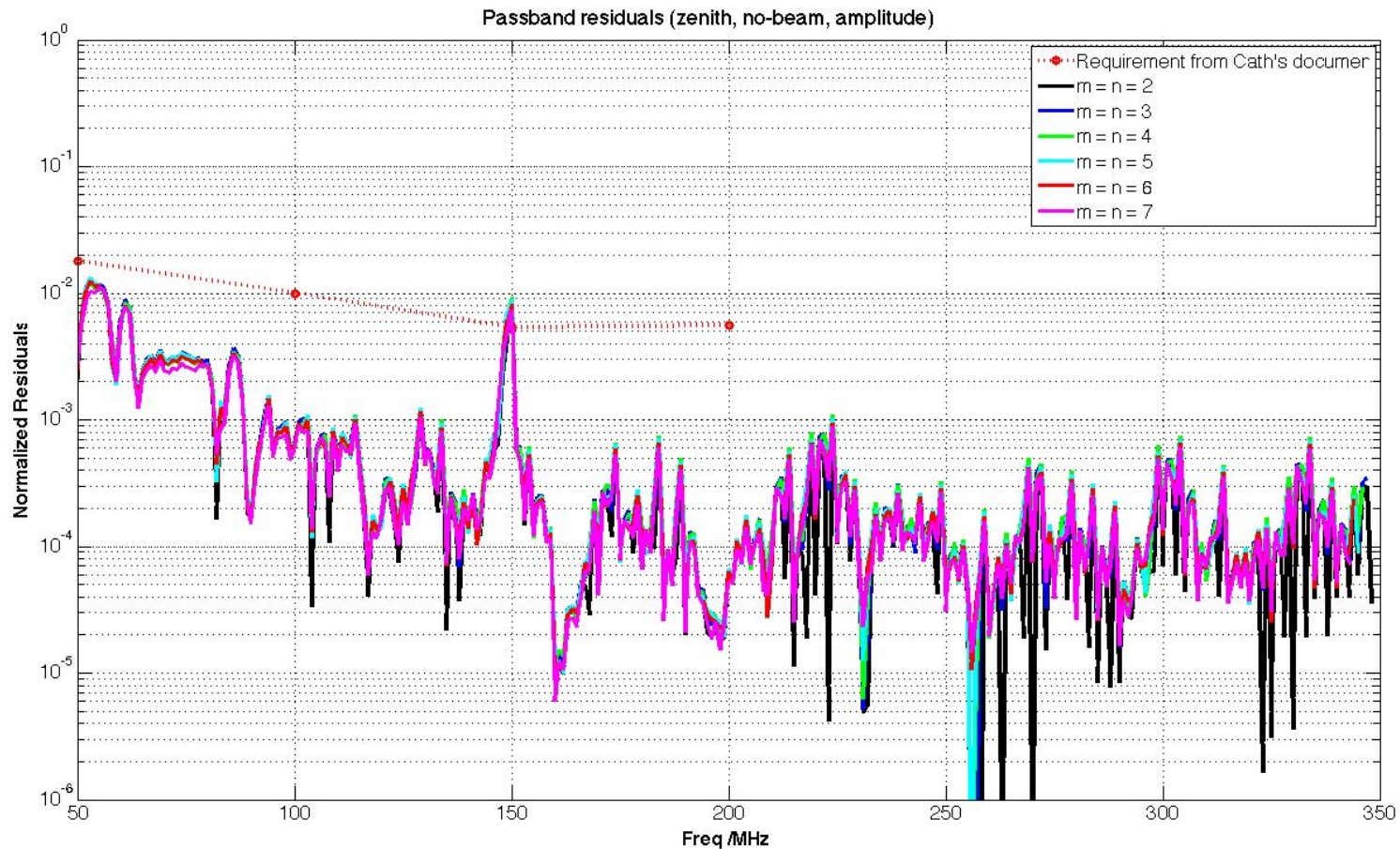
Update on the solution (SKALA3)



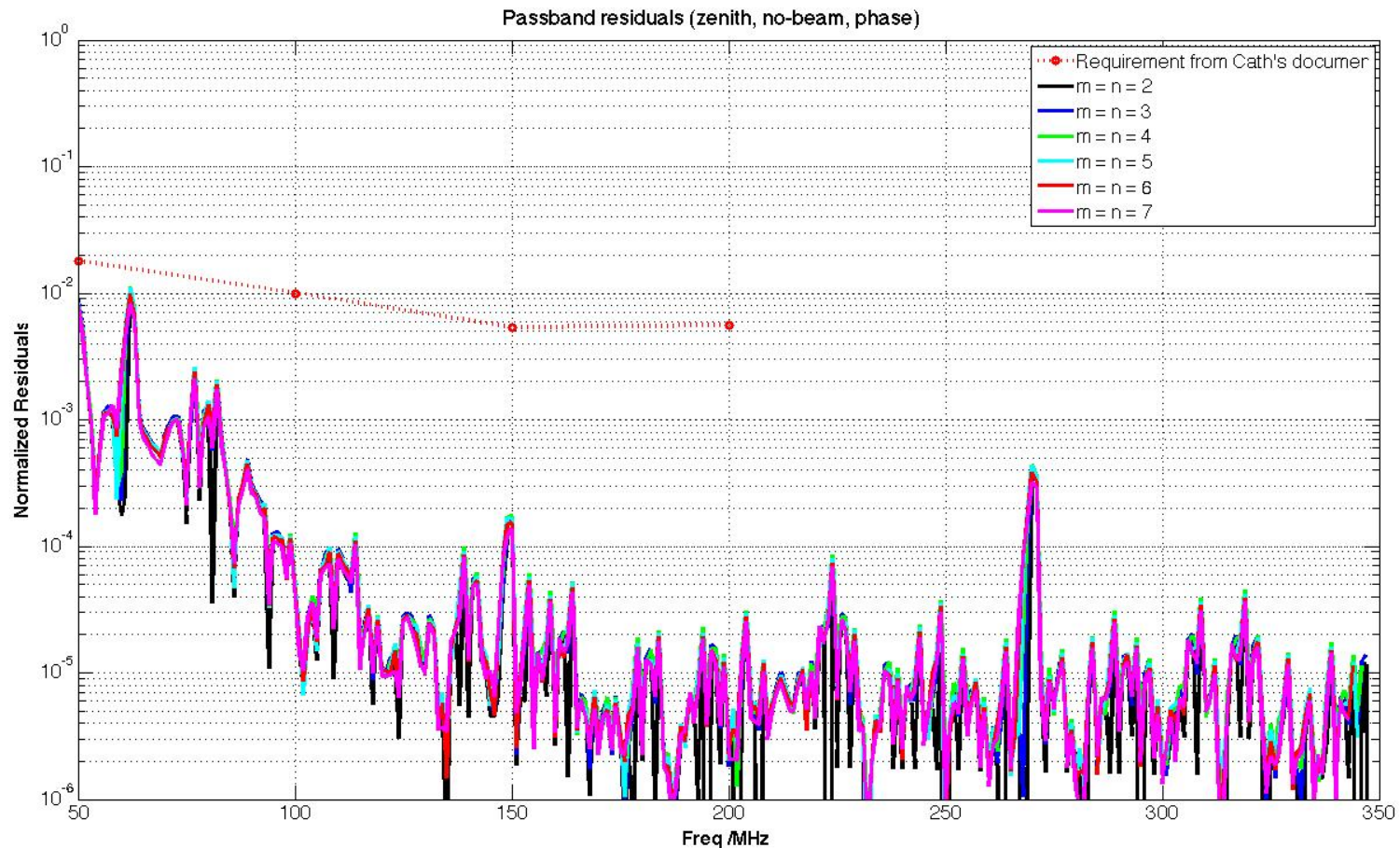
Update on the solution (SKALA3)



Update on the solution (SKALA3)



Update on the solution (SKALA3)



Next steps

- On the analysis:
 - Inclusion of polarization effects and DDEs.
 - OSKAR simulations.
- On the antenna design:
 - Converge into a satisfactory solution (summer 2016). Can we improve further by:
 - Increasing the number of dipoles and their frequency spacing?
 - Tuning the spine angles and dipole lengths?
 - Tuning the antenna length?
 - Further tuning on LNA input network.
 - Build of prototypes to test in isolation and with Pre-AAVS1 and AAVS1 (Q4 2016).
 - The new design may require a slight increase in the station diameter (up to ~40-45 m to maintain the antenna density and accessibility).

Conclusions

- AAVS1 antennas + LNAs are ready
- Optimization of SKALA2 -> SKALA3 under control and on its way

Thank you

