

MULTIFREQUENCY VLBA OBSERVATIONS OF TWO COMPACT SYMMETRIC OBJECTS:

0108+388

and

J1944+5448

Sara Rastello

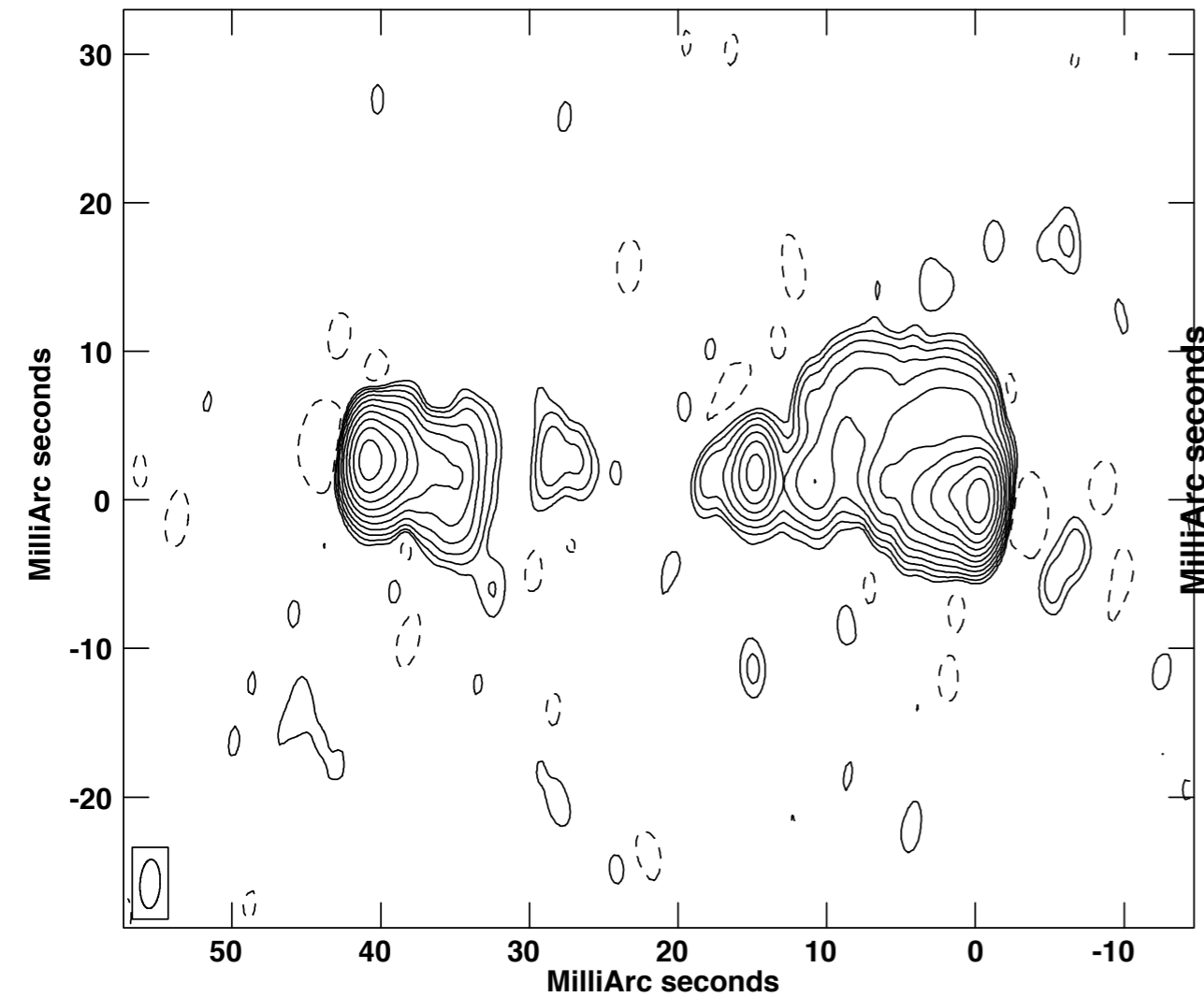
Daniele Dallacasa, Monica Orienti

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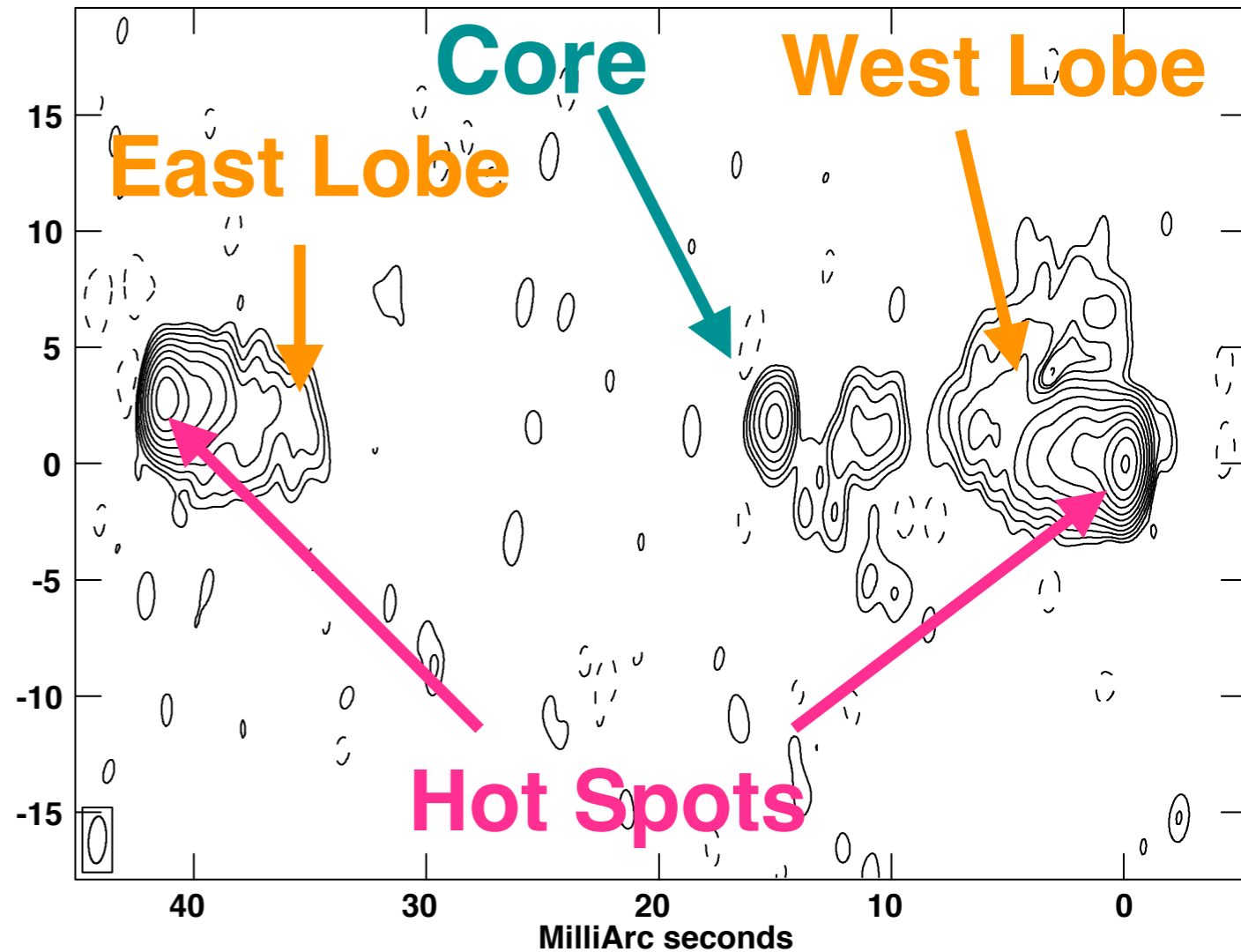
OUTLINE

- **Morphological properties of 1944+5448 and 0108+388**
- **Some useful physical parameters**
- **Hot spots advance speed**
- **Age**
- **Spectrum of 0108+388**

J1944+5448

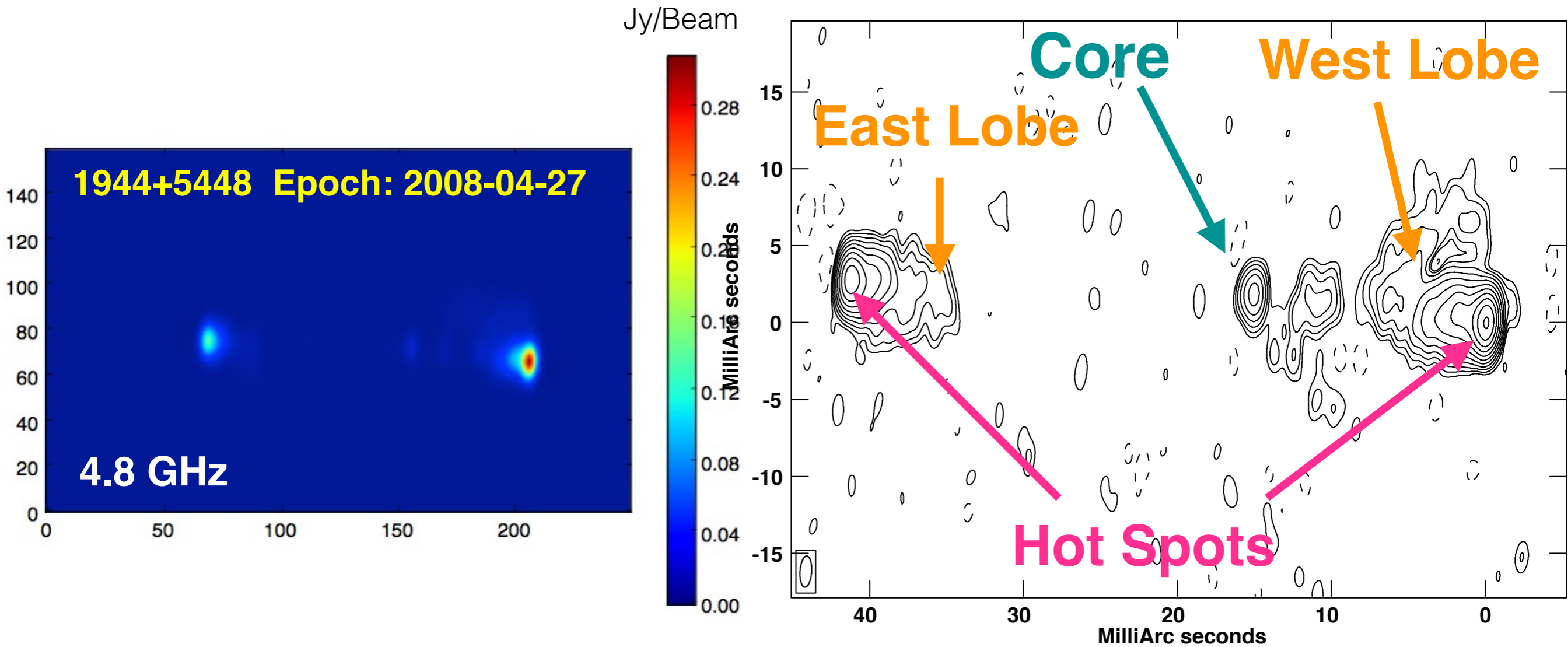


C Band 4.88 GHz



X Band 8.48 GHz

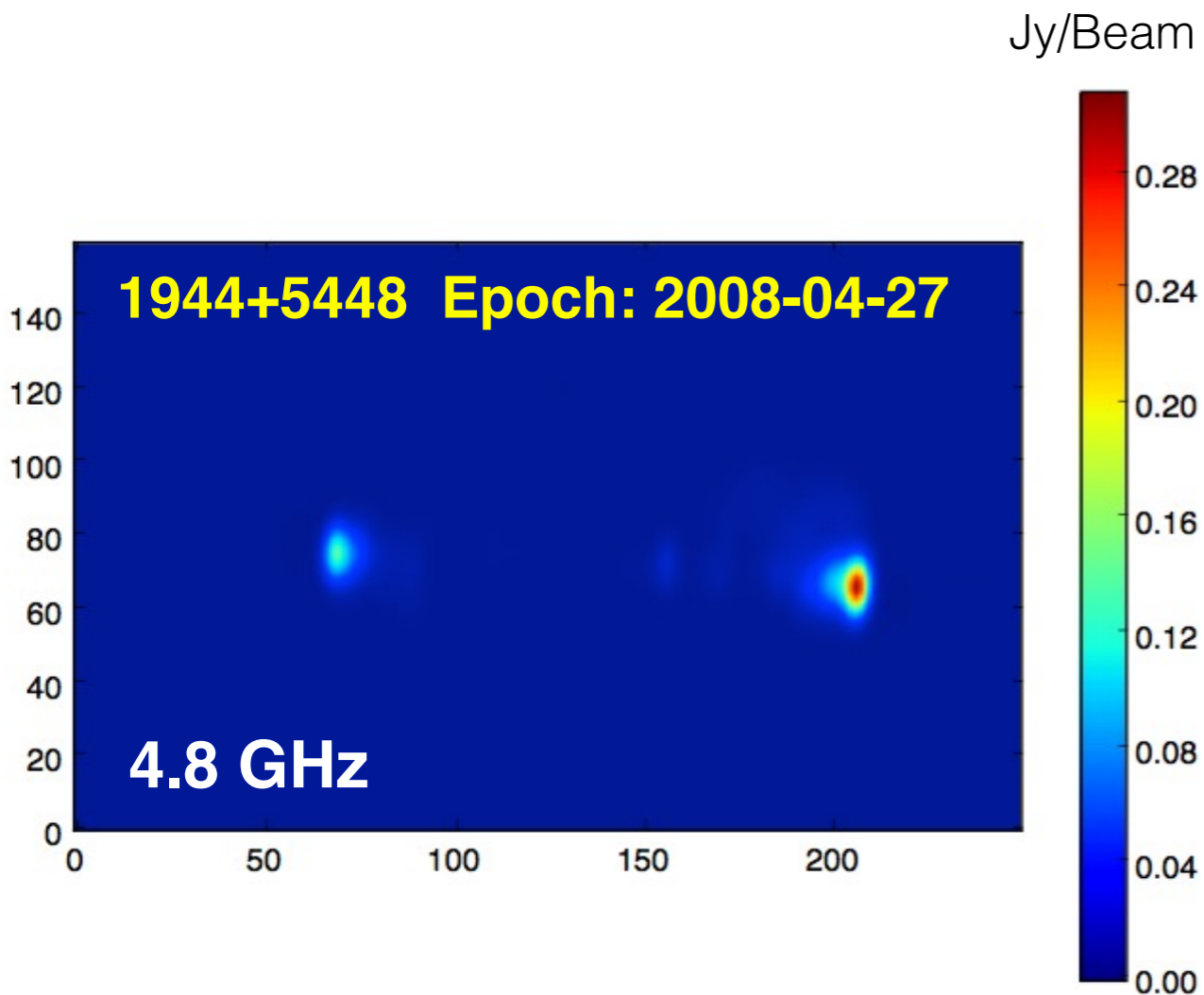
J1944+5448



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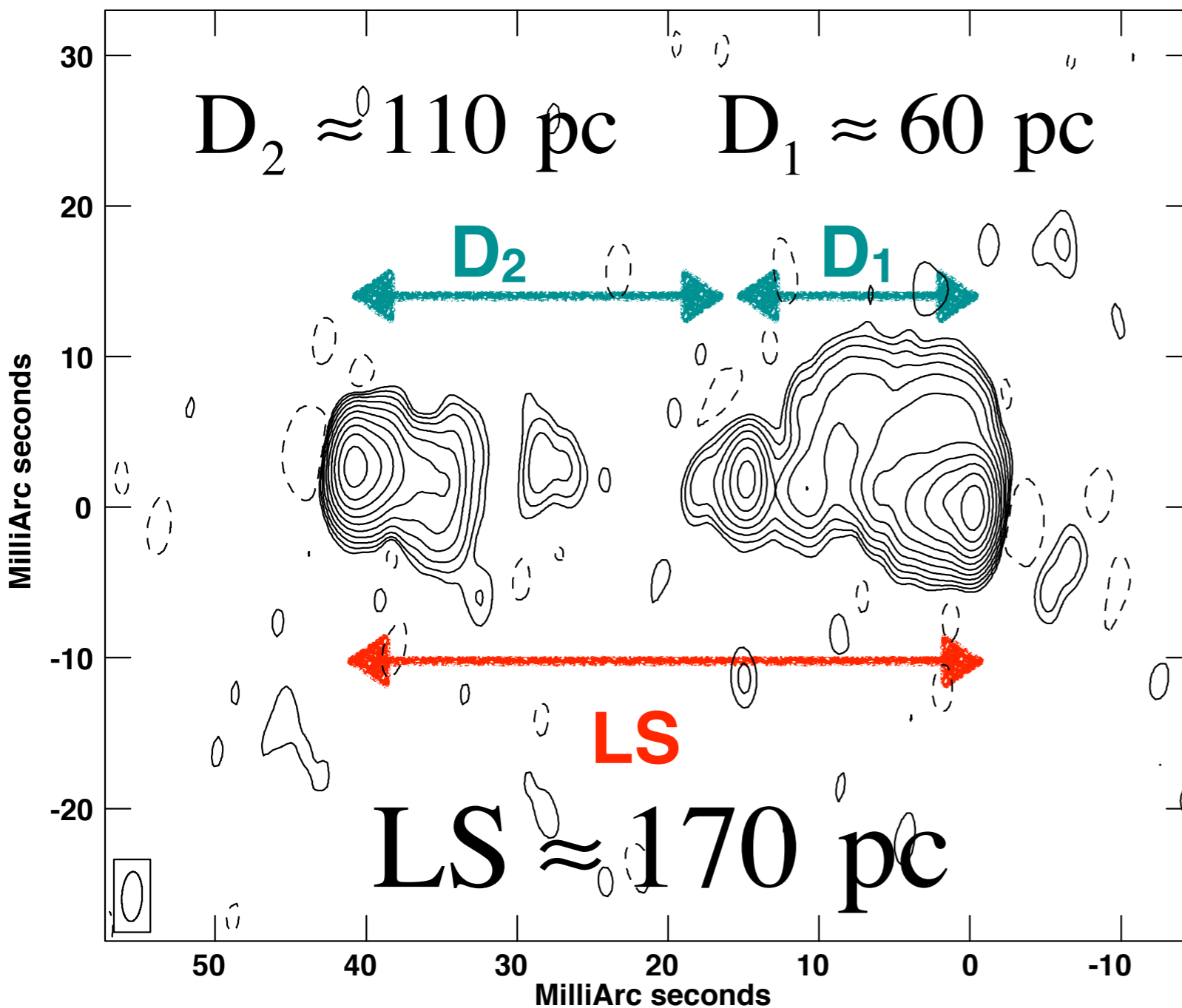
Fluxes Ratio:

$$\text{Hot Spot } \frac{F_W}{F_E} \geq 3$$

$$\text{Hot Spot /Lobe } \frac{F_W}{F_W} \approx 3$$

$$\text{Hot Spot /Core } \frac{F_W}{F_C} \approx 22$$

J1944+5448



$$\text{arm length ratio} = \frac{D_2}{D_1} \approx 2$$

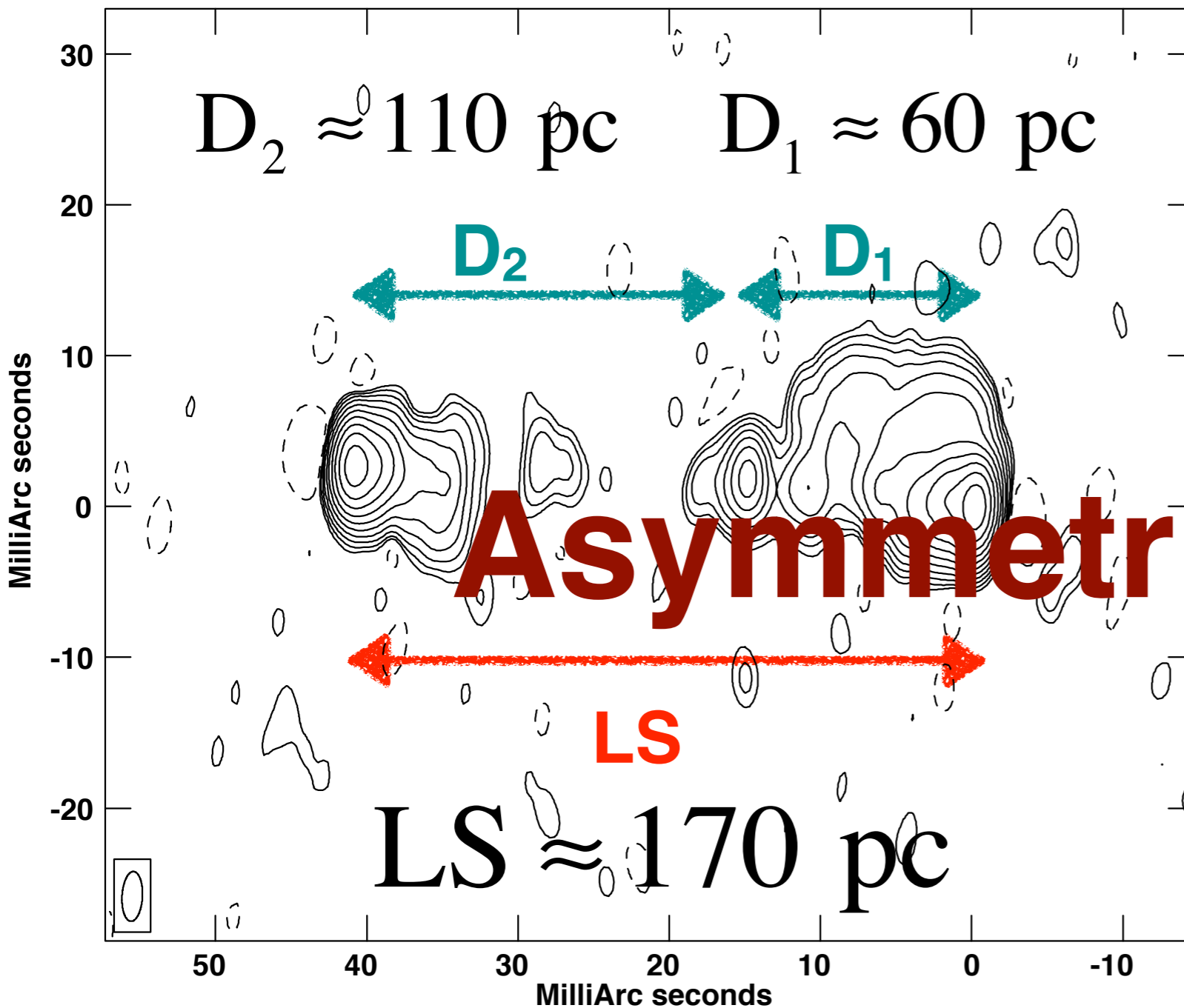
Fluxes Ratio:

$$\text{Hot Spot } \frac{F_W}{F_E} \geq 3$$

$$\text{Hot Spot /Lobe } \frac{F_W}{F_L} \approx 3$$

$$\text{Hot Spot /Core } \frac{F_W}{F_C} \approx 22$$

J1944+5448



$D_2 \approx 110$ pc

$D_1 \approx 60$ pc

Fluxes Ratio:

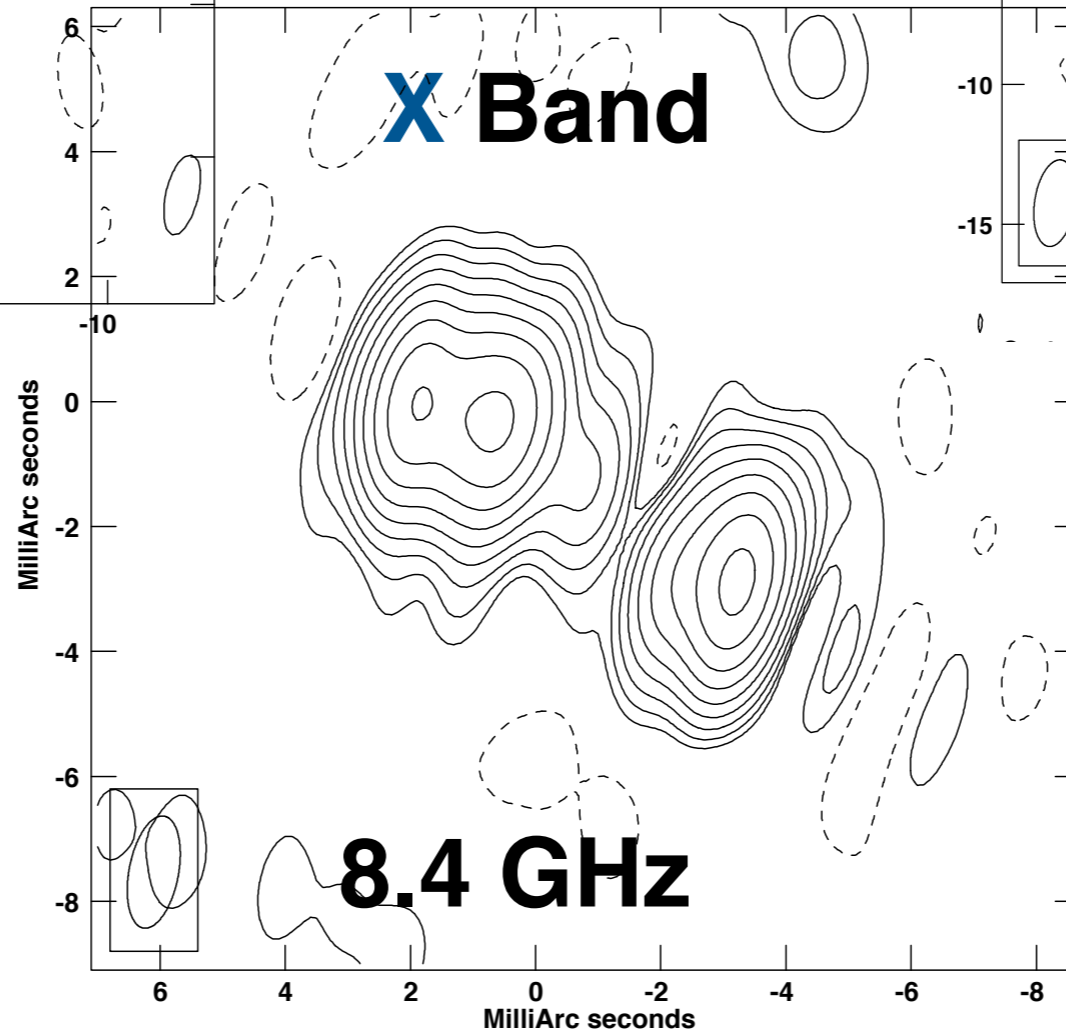
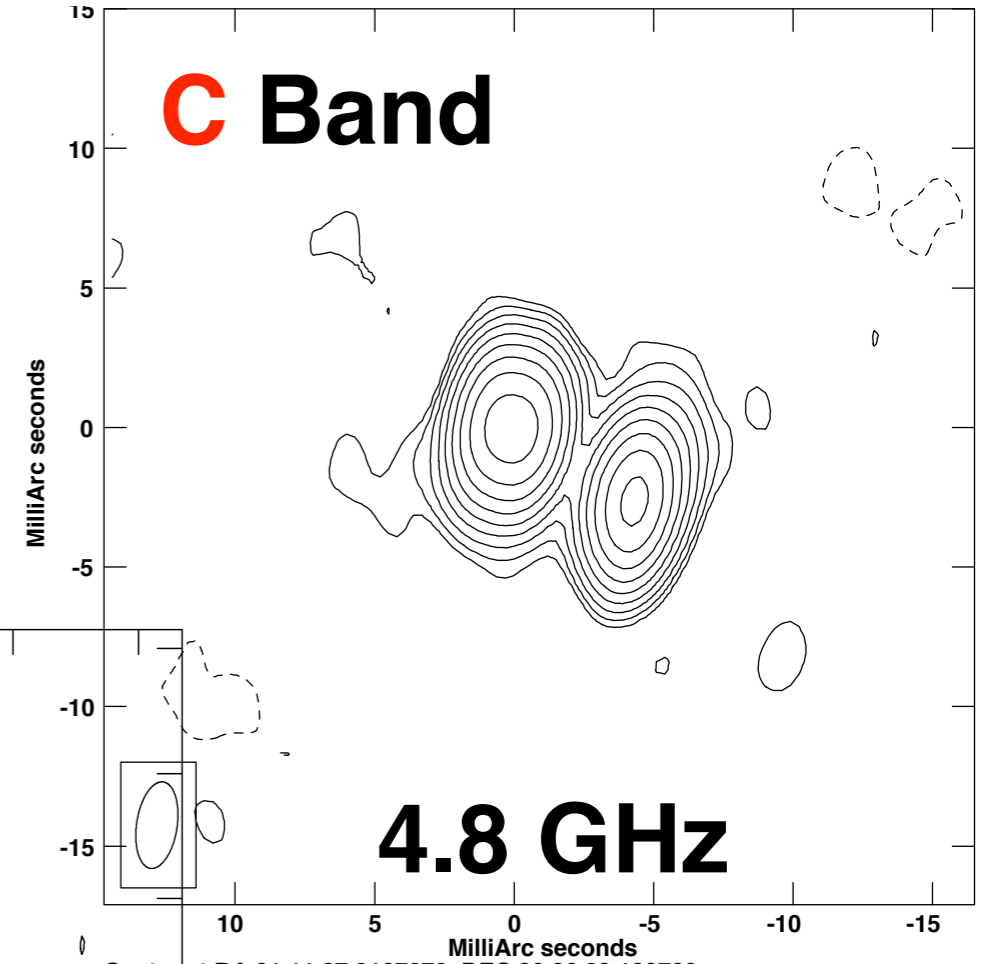
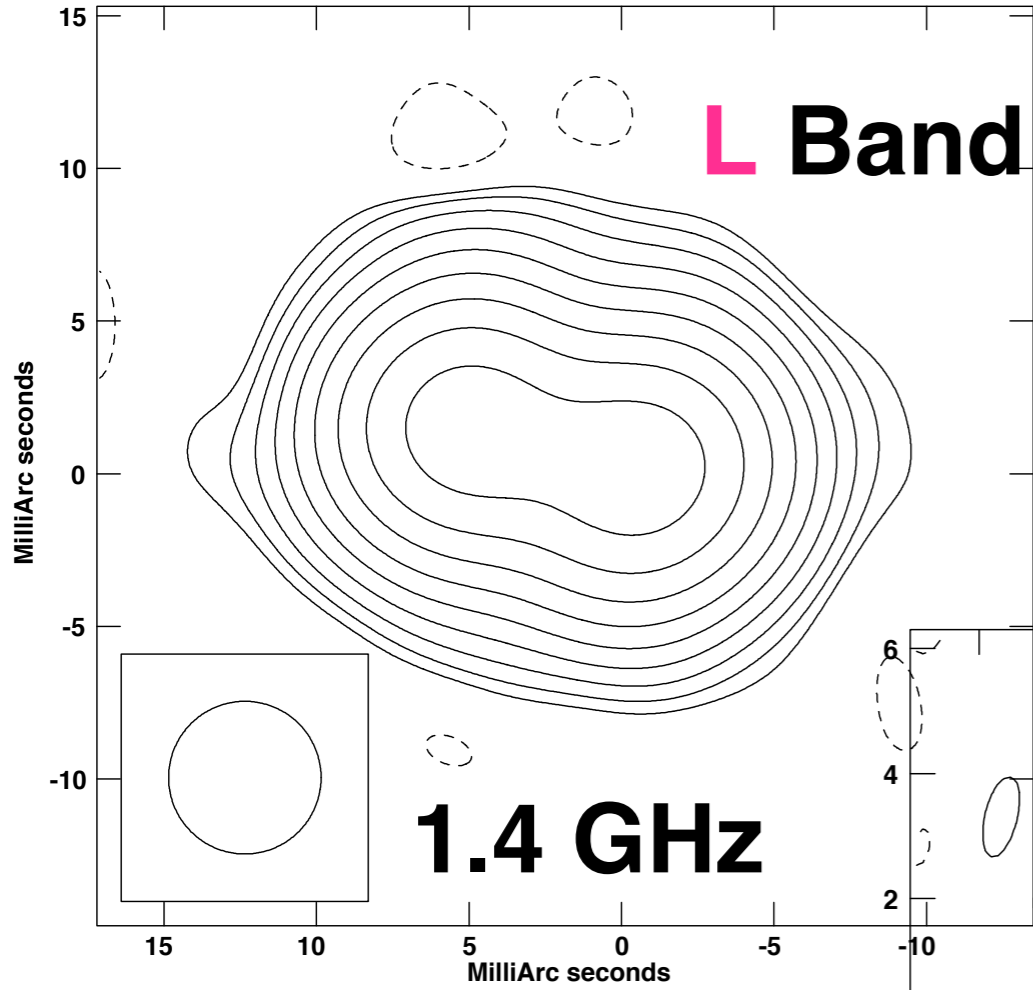
Asymmetric source!!

LS

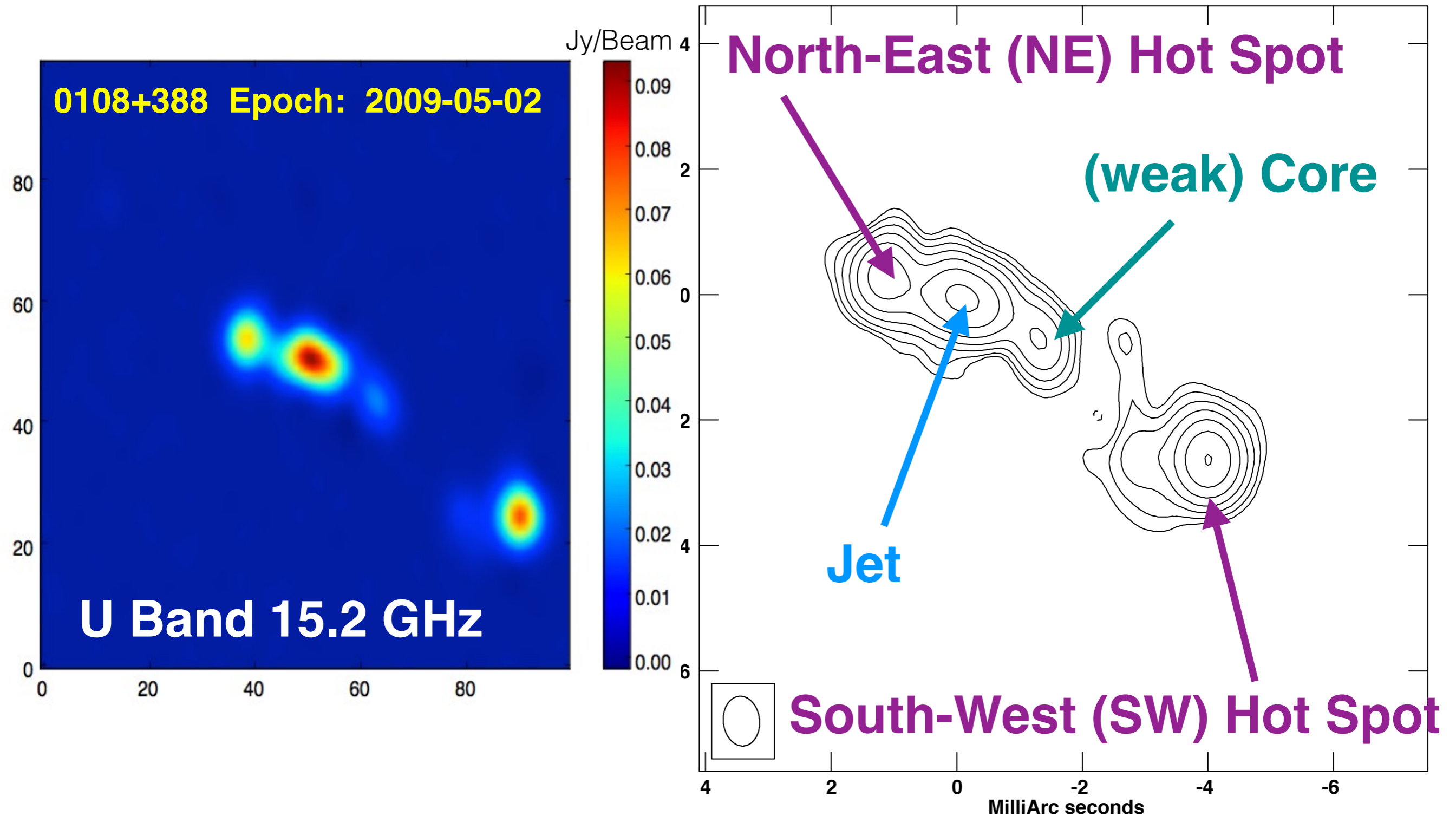
$LS \approx 170$ pc

$$\text{arm length ratio} = \frac{D_2}{D_1} \approx 2$$

0108+388



0108+388



This research has made use of data from the MOJAVE database that is maintained by the MOJAVE team ([Lister et al., 2009, AJ, 137, 3718](#)).

0108+388

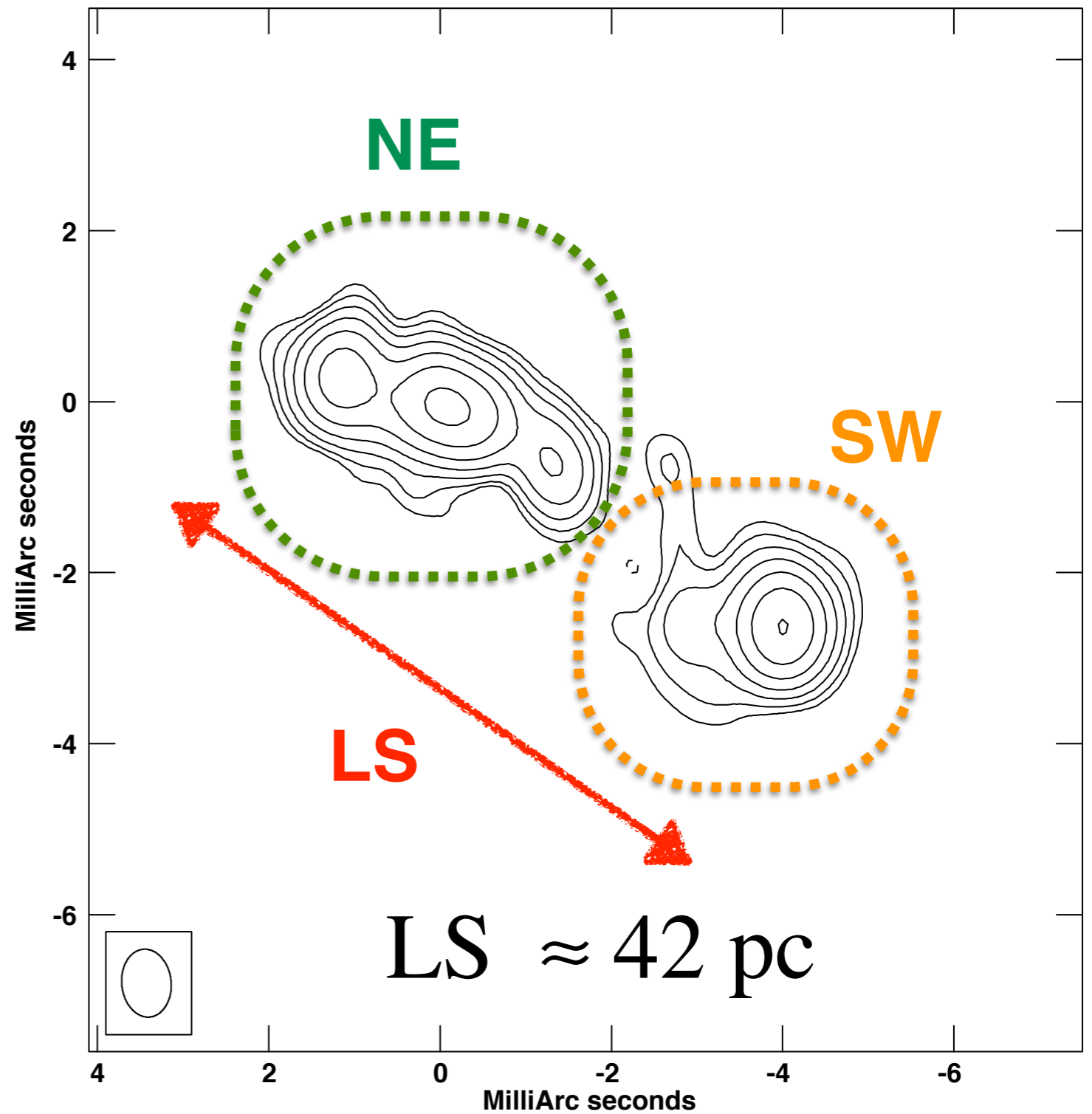
Fluxes Ratio:

$$L \longrightarrow \frac{F_{NE}}{F_{SW}} \approx 1$$

$$C \longrightarrow \frac{F_{NE}}{F_{SW}} = 1.5$$

$$X \longrightarrow \frac{F_{NE}}{F_{SW}} \approx 2$$

$$U \longrightarrow \frac{F_{NE}}{F_{SW}} \geq 2$$

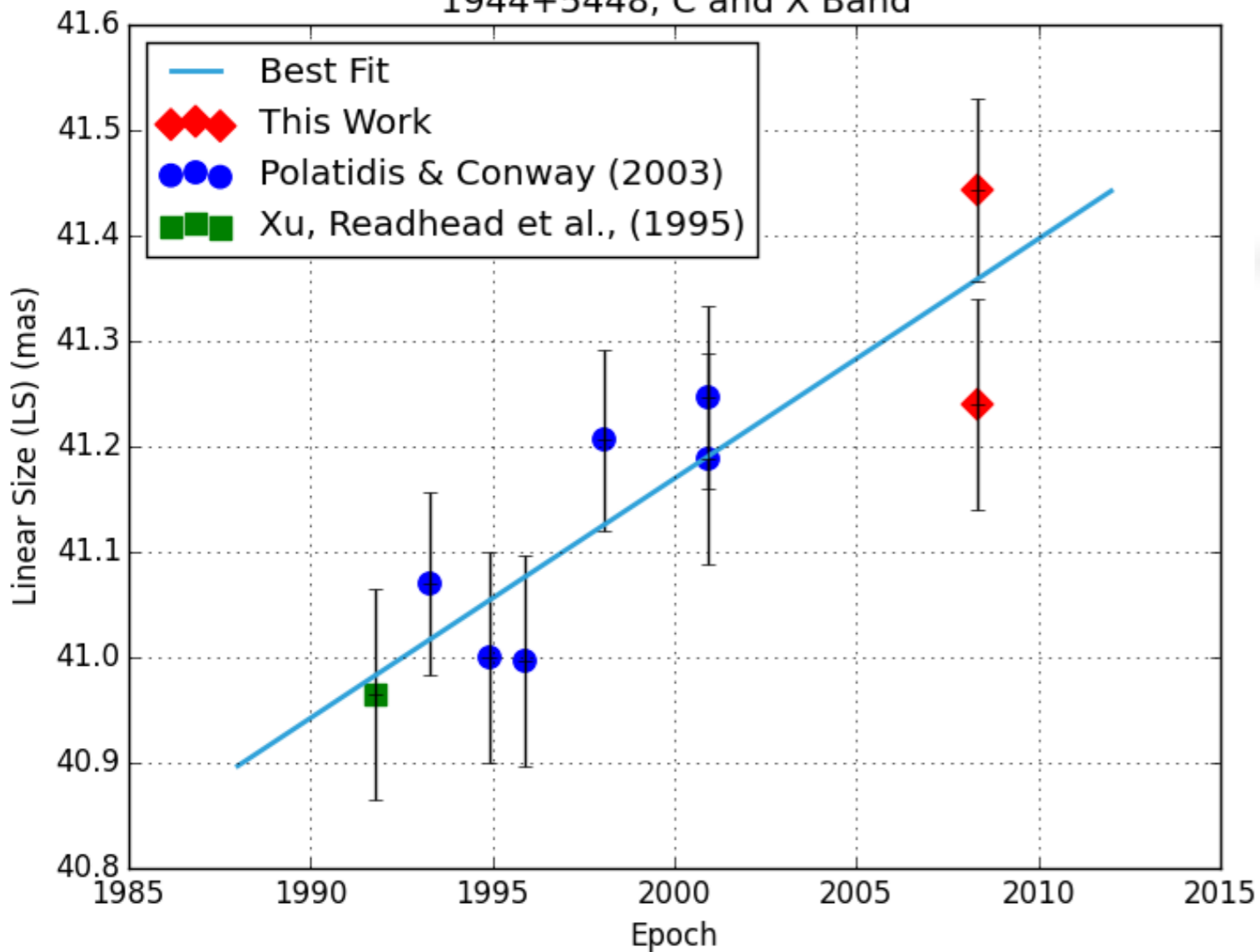


Physical parameters:

<u>1944+543</u>	Volume pc ³	L _(5GHz) erg/s	H _(Equipartition) mG	P _(Equipartition) dyne/cm ²
East lobe	2 10 ⁴	2.6 10 ⁴²	2.7	1.6 10 ⁻⁶
East HS	94.7	1.7 10 ⁴³	22.5	1.7 10 ⁻⁴
West Lobe	9 10 ⁴	1.2 10 ⁴³	2.6	1.7 10 ⁻⁶
West HS	85.6	3.6 10 ⁴³	28.6	1.1 10 ⁻⁴
Core	67.2	1.1 10 ⁴²	12.6	3.3 10 ⁻⁵

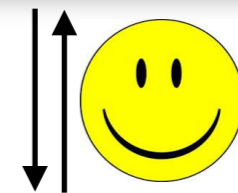
1944+543

1944+5448, C and X Band



Hot Spots separation speed

$$v_{sep} = 0.24 \pm 0.06c$$



$$v_{sep} = 0.26 \pm 0.04c$$

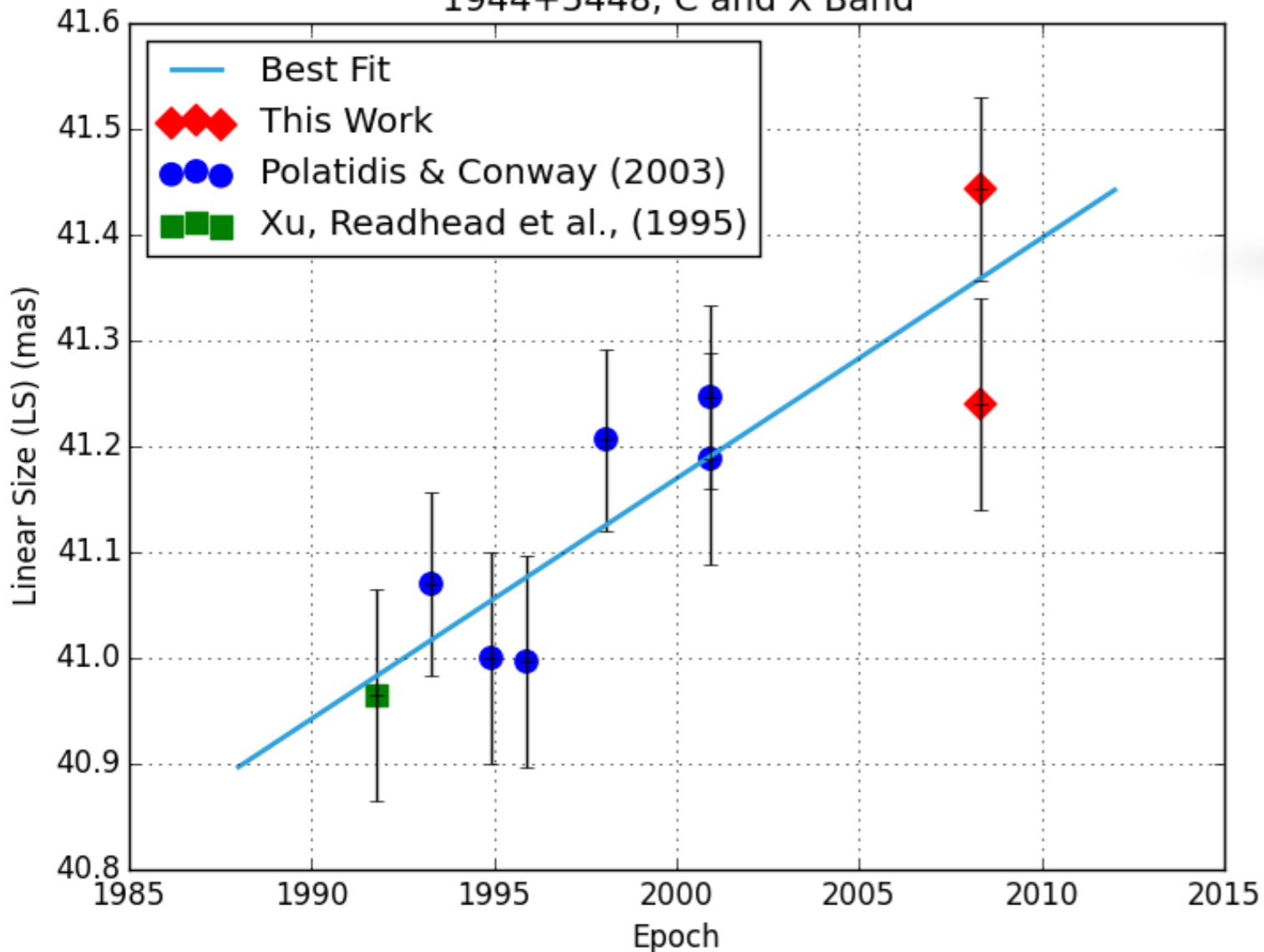
(Polatidis & Conway 2003)

Kinematic Age

$$t_{kin} = 1800 \pm 400 y$$

1944+543

1944+5448, C and X Band



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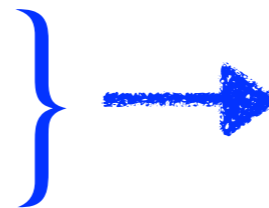
$$t_{kin} = 1800 \pm 400 y$$

Radiative Age

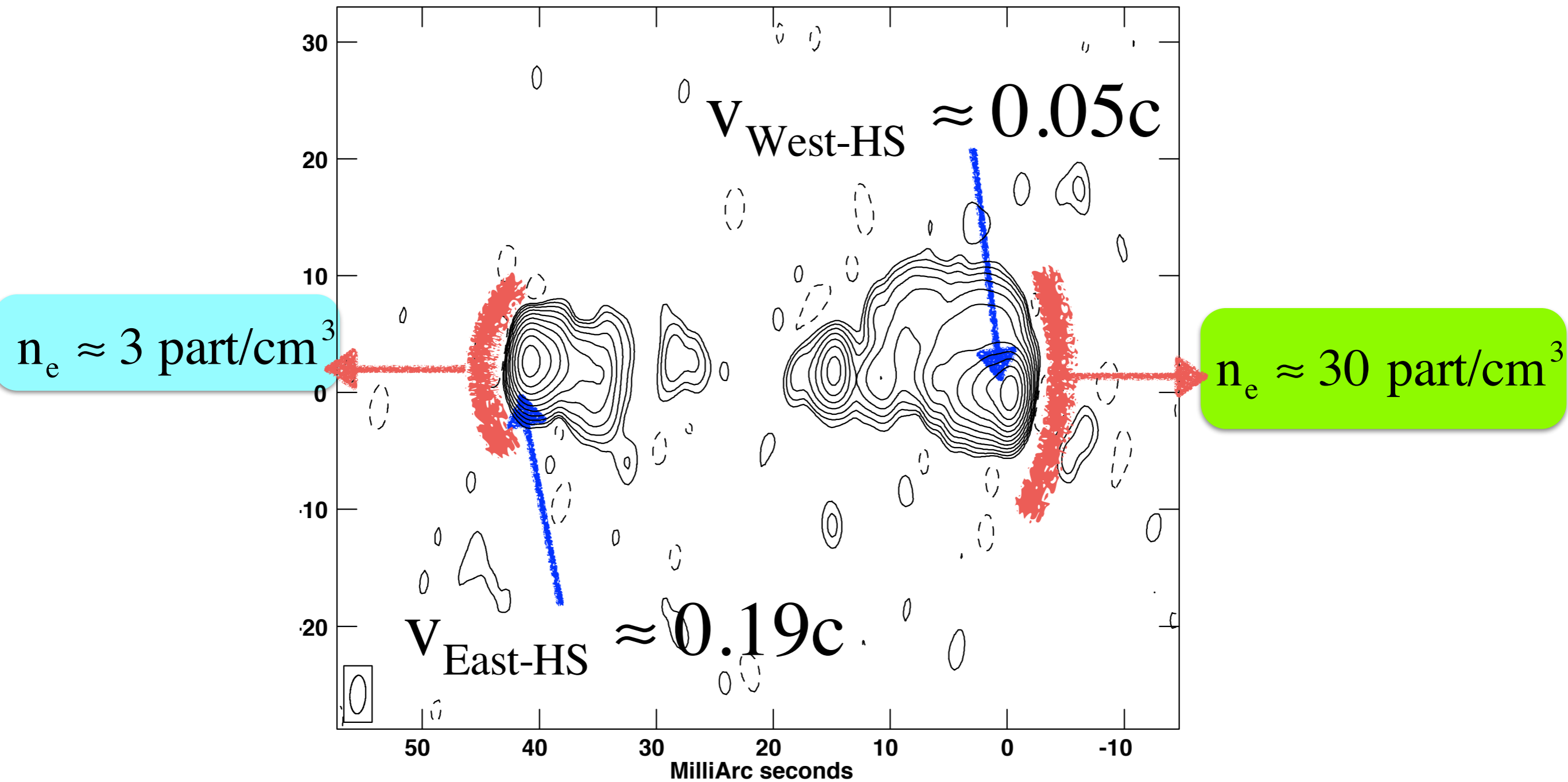
$$t_{rad} = 1900 \pm 500 y$$

$$t_{rad} = 5.03 \times 10^4 \cdot H_{eq}^{-1.5} [(1+z) v_{break}]^{-0.5} \text{ (yr)}$$

Murgia (2003)



J1944+5448



Equipartition Pressure



Ram Pressure

0108+388

Hot Spots separation speed

$$V_{sep} = 0.07 \pm 0.02c$$

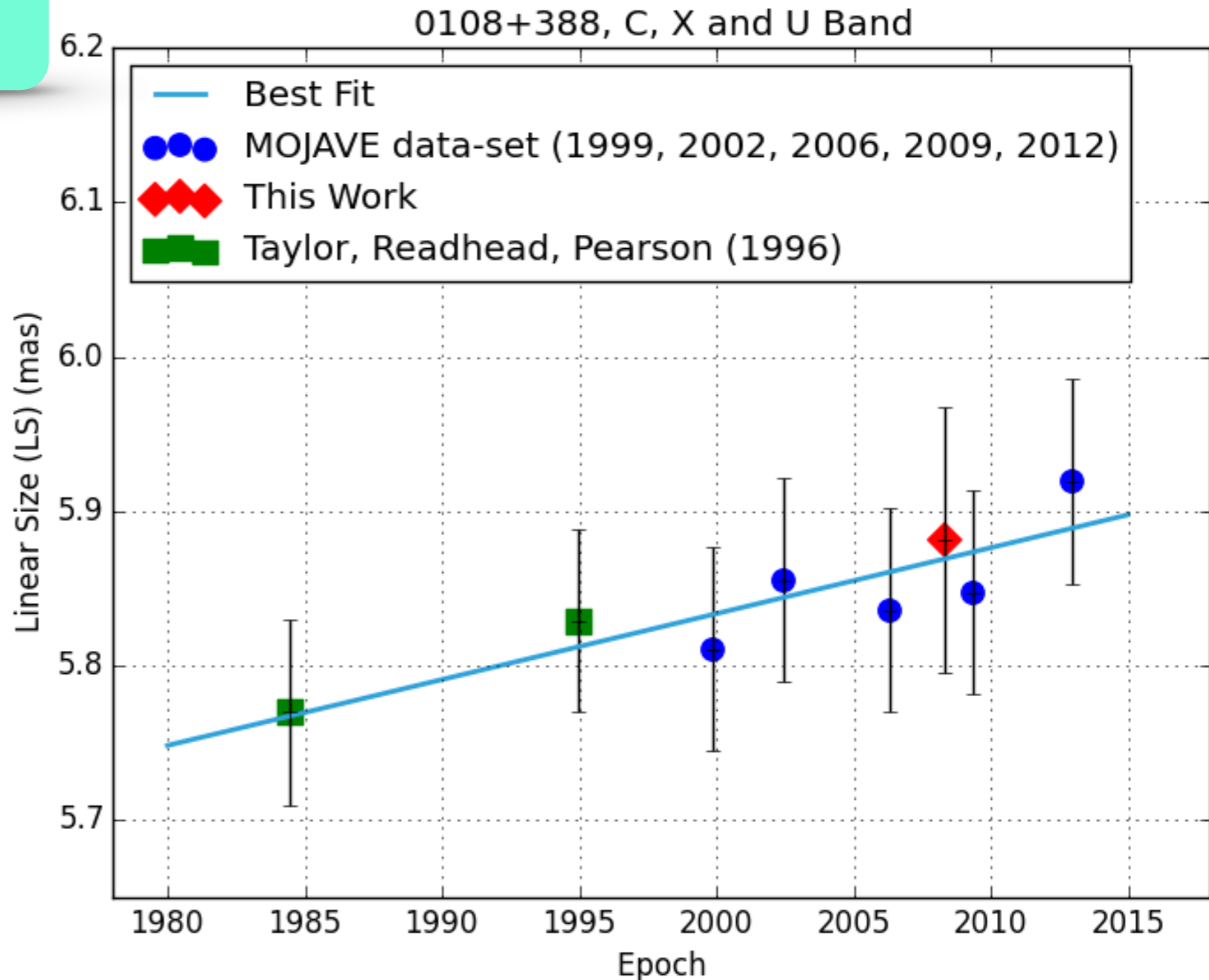


$$V_{sep} = 0.02 \pm 0.01c$$

(Taylor, Marr et al., 2000)

Kinematic Age

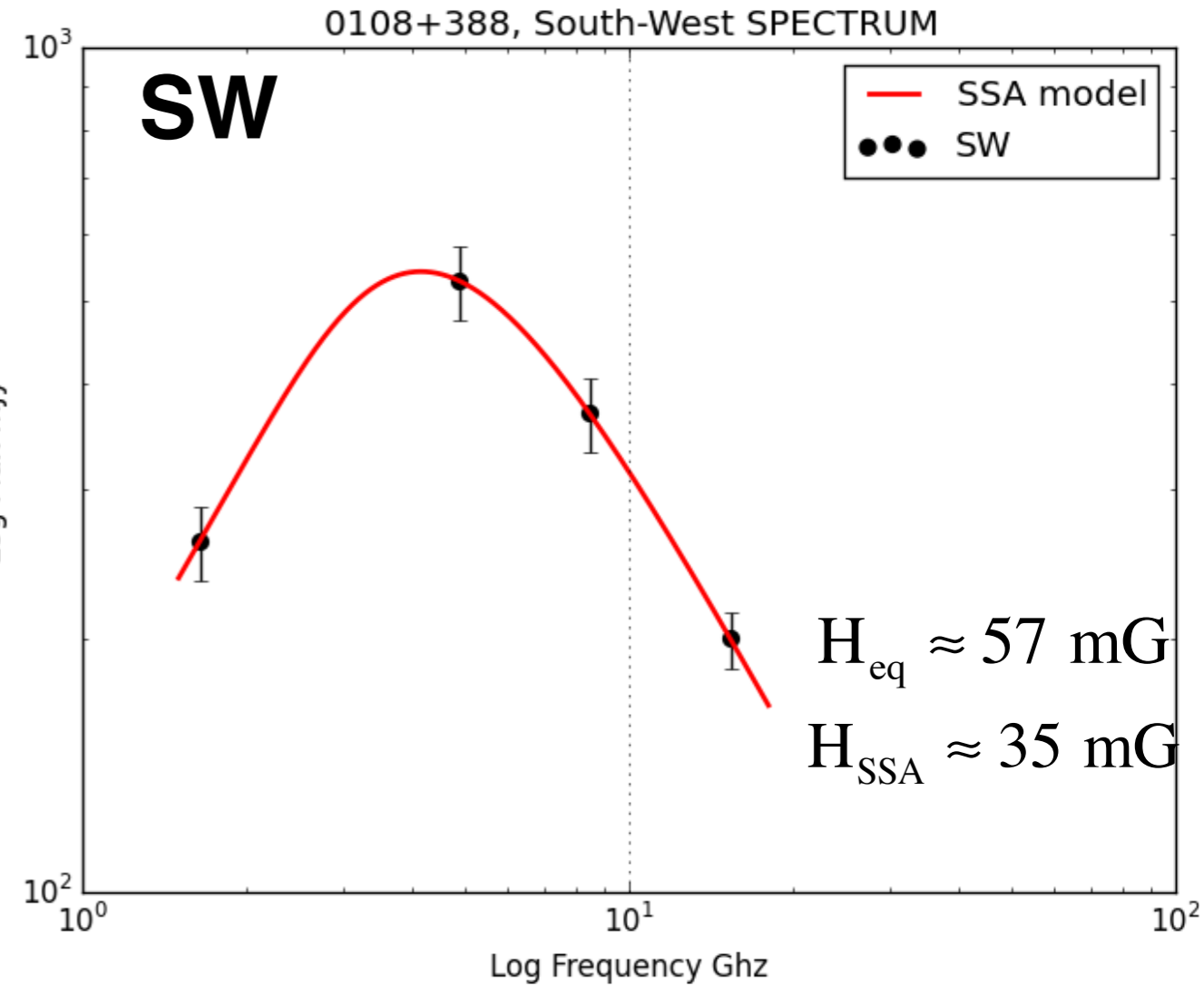
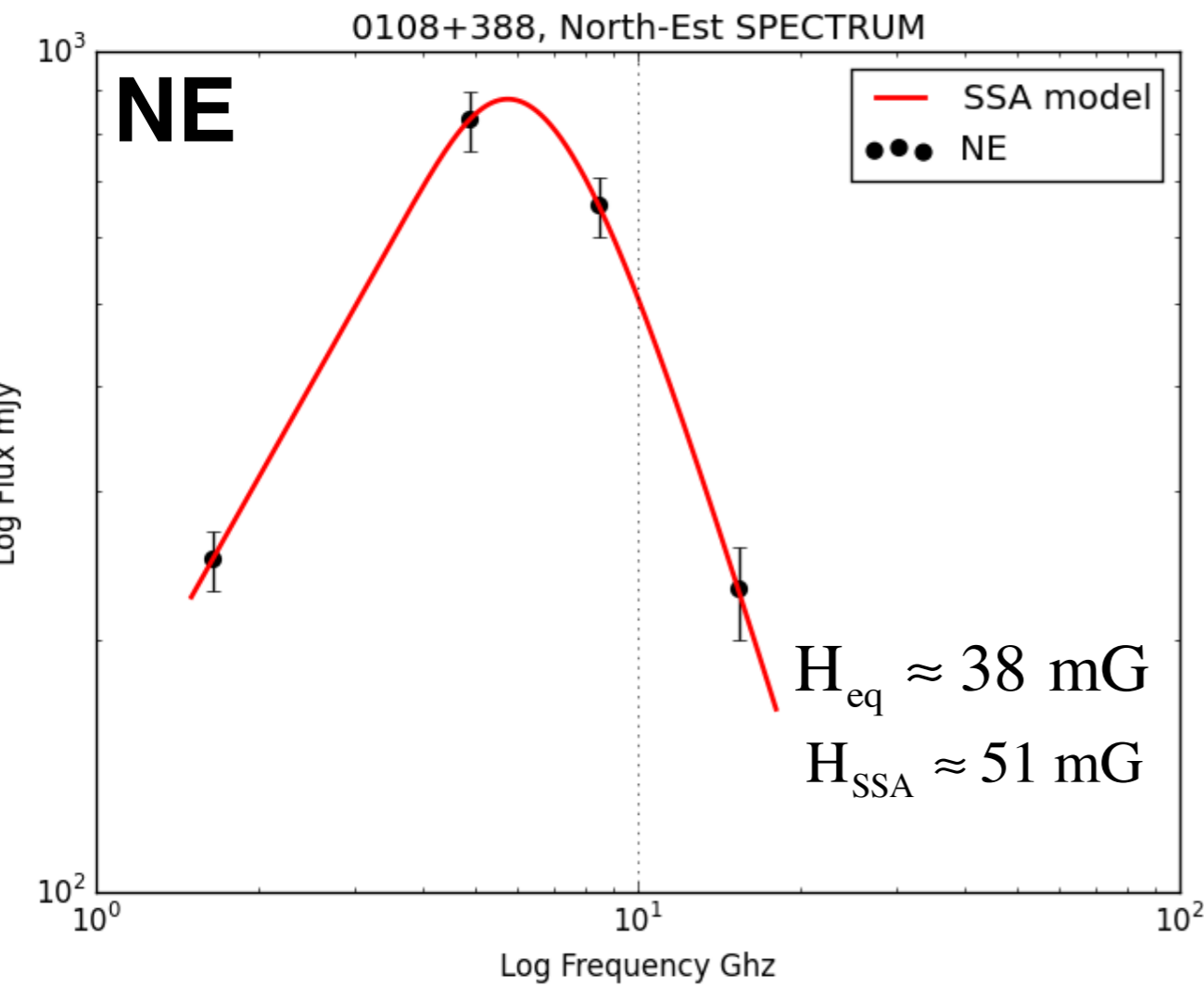
$$t_{kin} = 1100 \pm 500y$$



Physical parameters:

<u>0108+388</u>	Volume pc ³	L(5GHz) erg/s	H(Equipartition) mG	P(Equipartition) dyne/cm ²
North-East	402.9	5.3 10 ⁴⁴	38.1	3.3 10 ⁻⁴
South-West	214.5	3.4 10 ⁴⁴	57.2	7.2 10 ⁻⁴

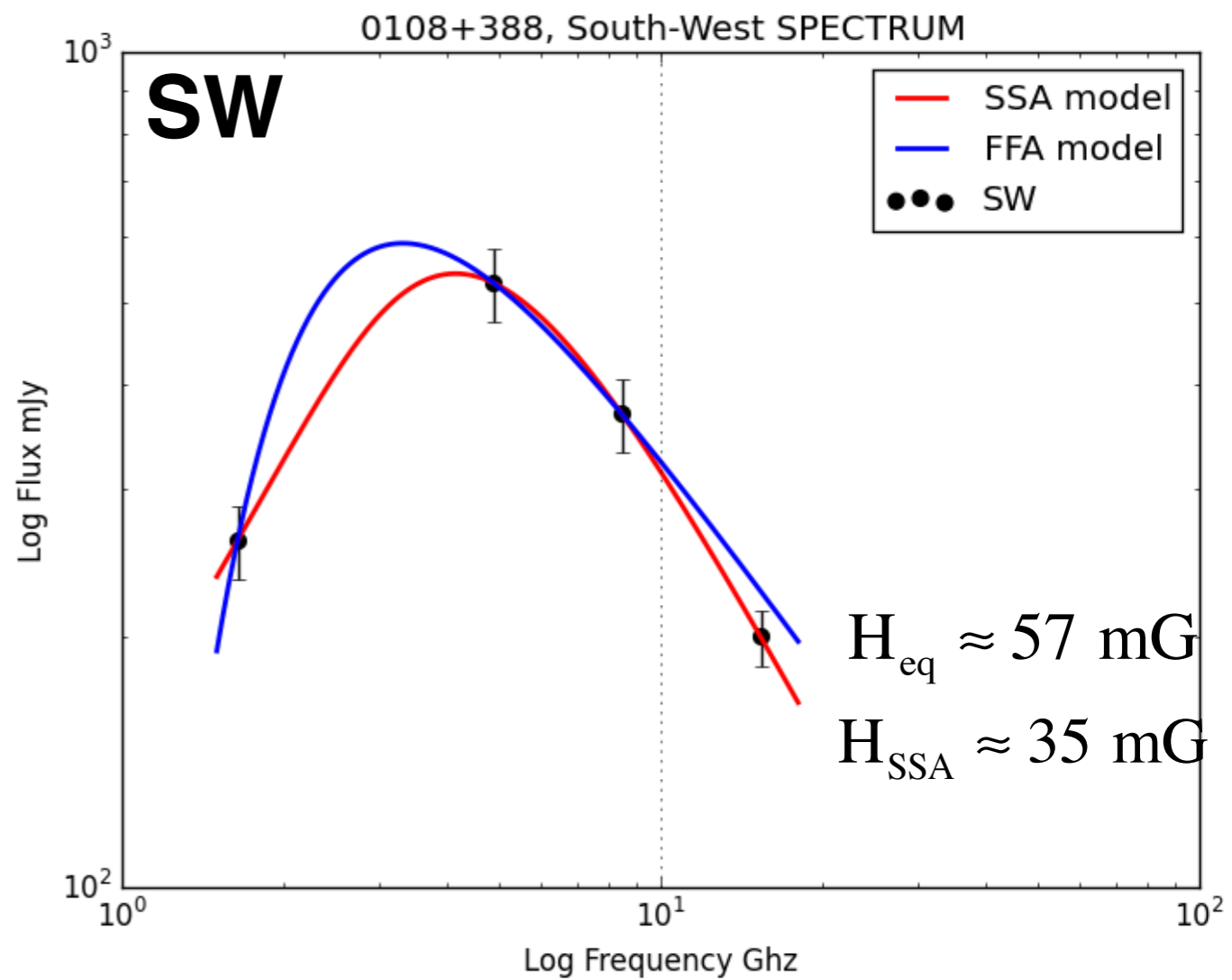
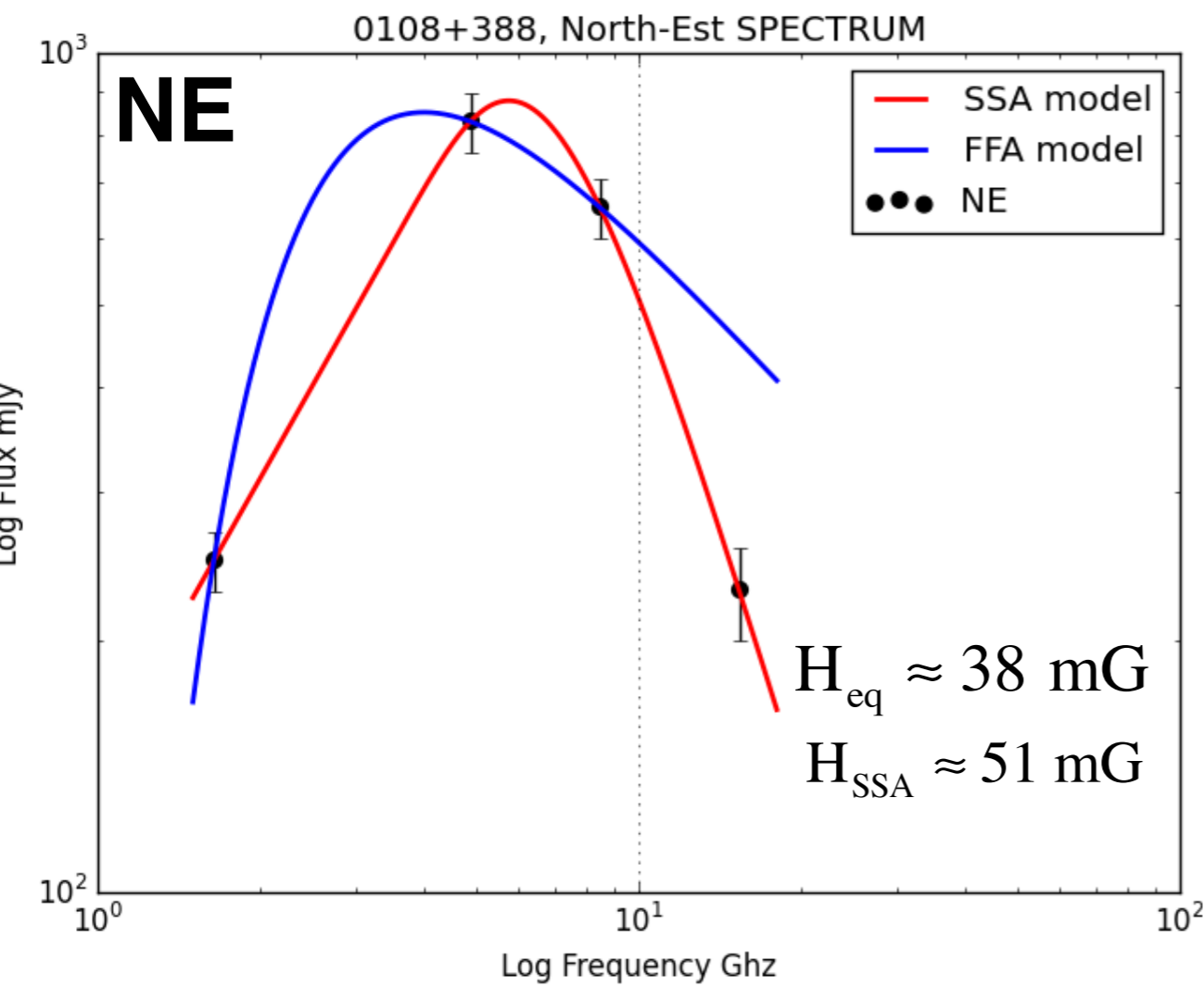
Spectral analysis 0108+388



SSA $\nu_{peak} \approx 6.9 \text{ GHz}$ $S_{peak} \approx 816 \text{ mJy}$

$\nu_{peak} \approx 4.6 \text{ GHz}$ $S_{peak} \approx 530 \text{ mJy}$

Spectral analysis 0108+388



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FFA $v_{peak} \approx 4 \text{ GHz}$ $S_{peak} \approx 849 \text{ mJy}$

$v_{peak} \approx 3.3 \text{ GHz}$ $S_{peak} \approx 590 \text{ mJy}$



Conclusions:

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- **Different magnetic fields**
 - 0108+388**  $H \approx 60\text{mG}$
 - 1944+543**  $H \approx 20\text{mG}$

Conclusions:

- **Young radio sources!** \longrightarrow $t \approx 2 \cdot 10^3$ years
- **1944+5448 inhomogeneous medium** \longrightarrow **Asymmetric source!**
- **Different magnetic fields**
 - 0108+388 \longrightarrow $H \approx 60\text{mG}$
 - 1944+543 \longrightarrow $H \approx 20\text{mG}$
- **0108+388 SSA or FFA?** \longrightarrow



*Thank
for
you
attention!*

SSA Model

In this case: $S(\nu) \propto \nu^\alpha$

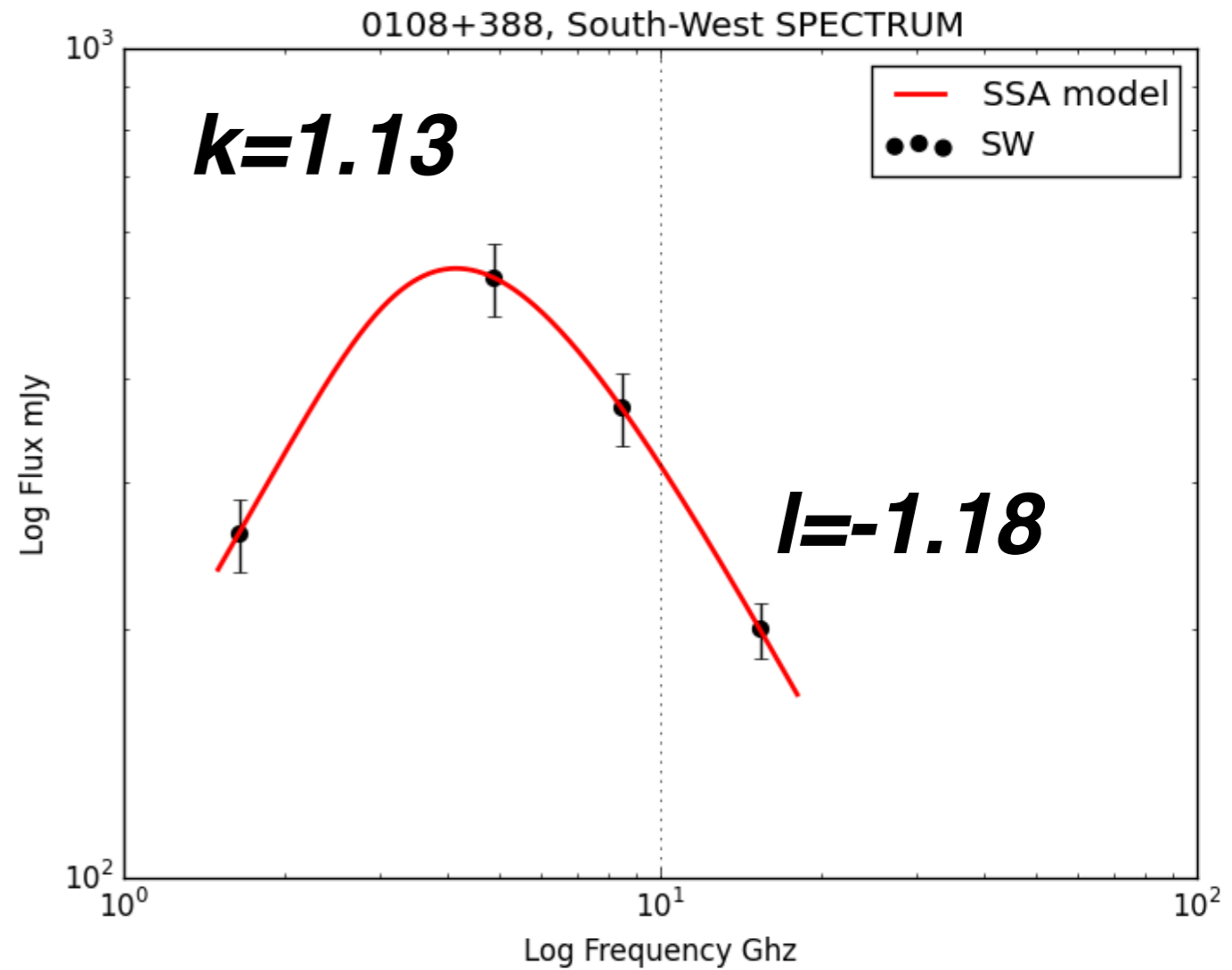
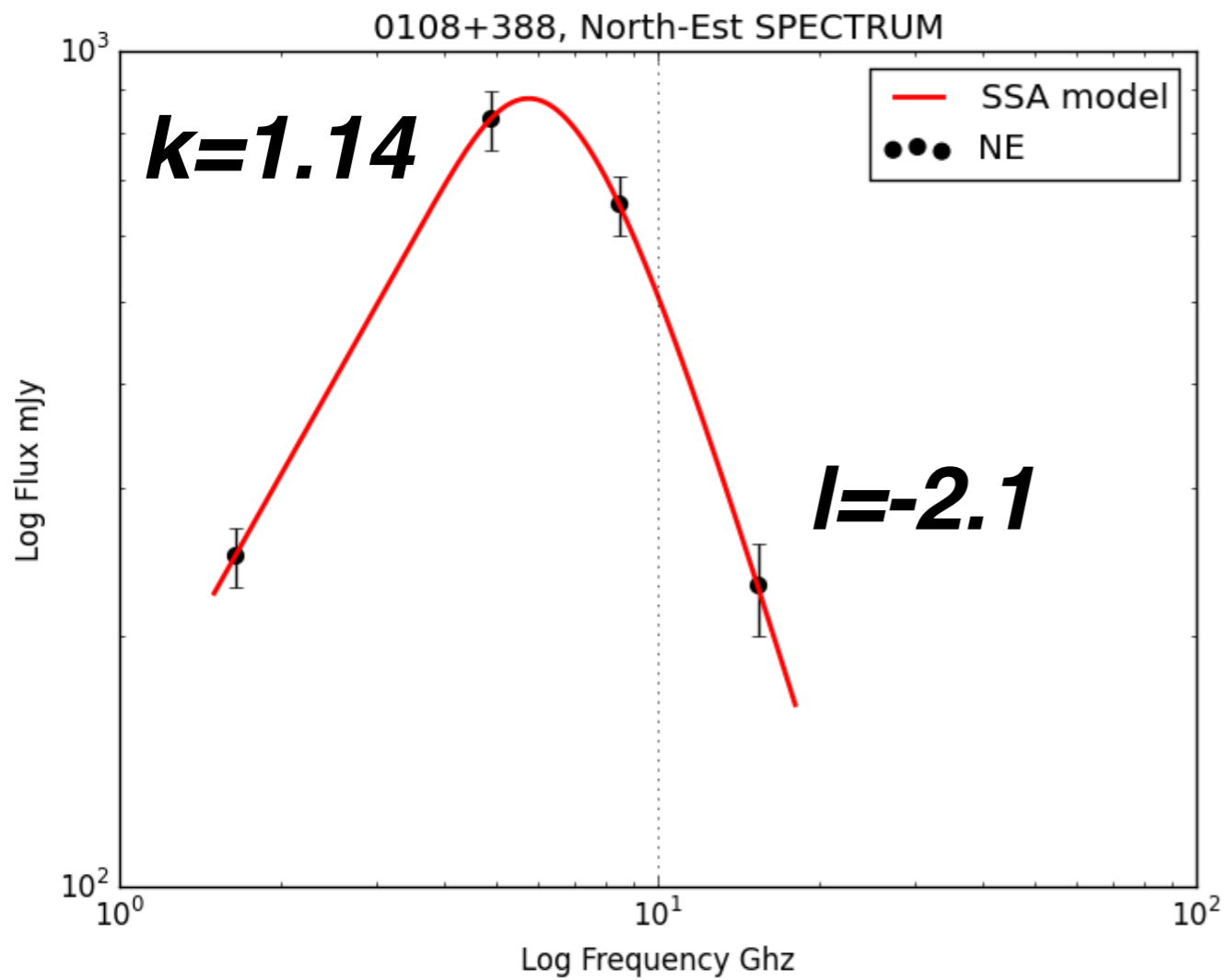
$$S(\nu) = \frac{S_{\max}}{(1 - e^{-1})} \cdot \left(\frac{\nu}{\nu_{\max}} \right)^k \cdot \left(1 - e^{-\left(\frac{\nu}{\nu_{\max}} \right)^{(l-k)}} \right)$$

k → optically thick

l → optically thin

“A new sample of Faint GigaHertz Peaked Spectrum Radio Sources” Snellen, Schilizzi, de Bruyn, Miley, Rengelink, Rottgering, Bremer 2008

SSA Model



FFA Model

Optical depth

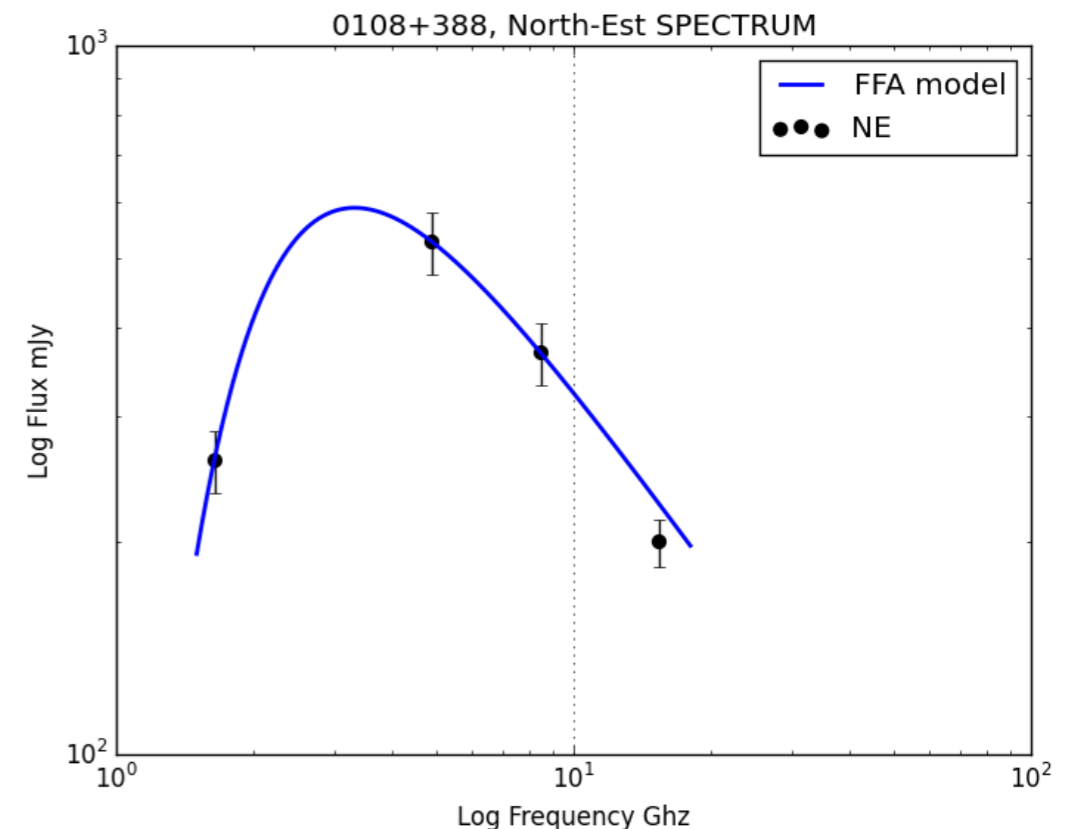
$$\tau_\nu = 0.08235 \nu^{-2.1} \int T_e^{-1.35} N_e^2 dl$$

Emission measure

$$\int N_e^2 dl$$



$$F_\nu = F_0 \left[\frac{\nu}{(1.5359 \text{ GHz})} \right]^\alpha e^{-\left(\frac{\text{EMT}^{-1.35}}{\nu^{2.1}} \right)}$$



“Nonuniform Free Free absorption in the GPS 0108+388”

Marr, Taylor, Crawford 2001, ApJ 550:160-171

FFA Model

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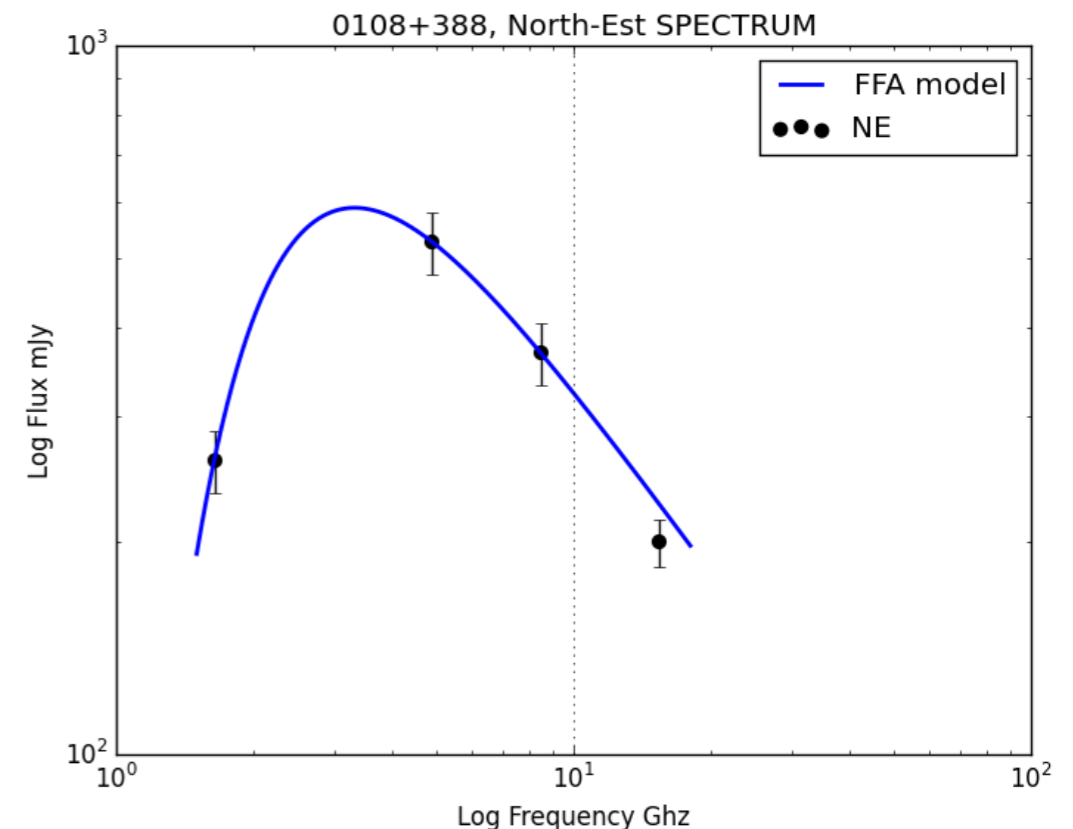
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Magnetic Field

$$H_{\text{eq}} = 4.5^{2/7} (1+k)^{2/7} c_{el}^{2/7} L^{2/7} V^{-2/7}$$

mG

$$H_{\text{SSA}} = f^{-5} \theta_{\text{max}}^2 \theta_{\text{min}}^2 v^5 S^{-2}$$

mG

