



The Decaparsec Scale Radio Structures of BL Lac Objects

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Outline

◉ Introduction

- › Objects
- › Observations
 - Two epoch, multi-frequency
 - Decaparsec scales
 - Spine – sheath Polarization
 - Farady rotation

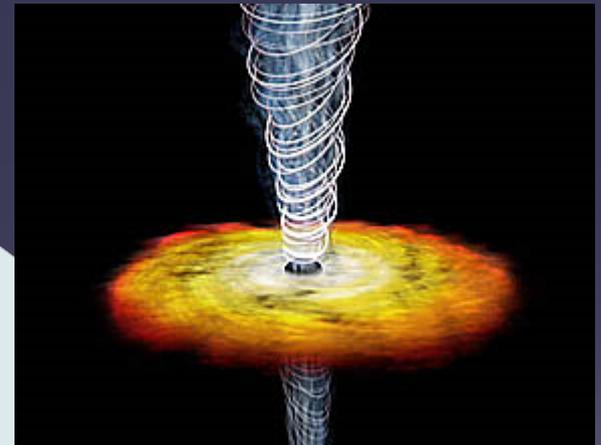
◉ Results

- Polarization structure
- Rotation measure
- Fractional Polarization

BL Lacs

Properties

- › Jet appears close to our line of sight
- › Beaming with high Doppler factors
- › Variable
- › Faint optical line emission
- › Synchrotron emission
 - Linear polarization

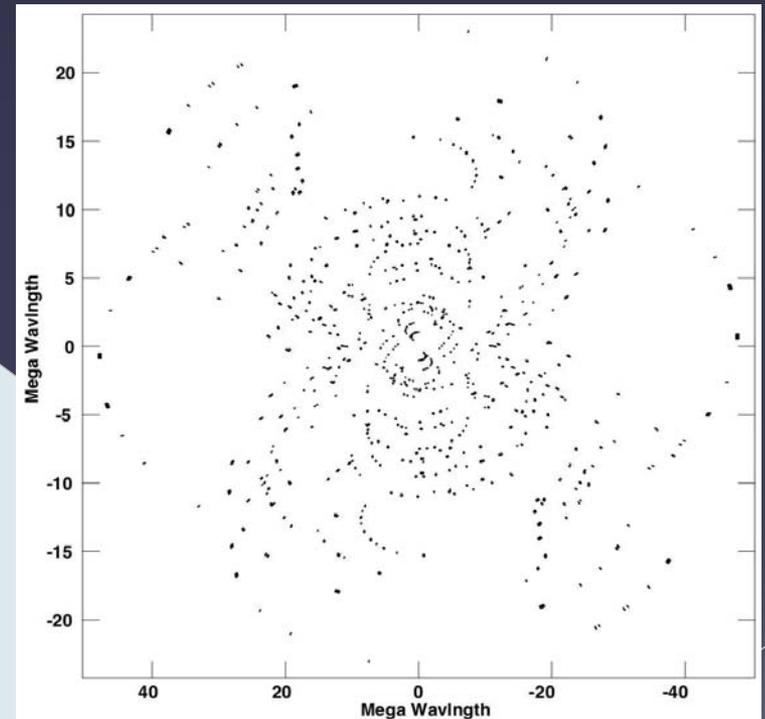


The Sample

- ◉ The sample of 34 northern BL Lac objects selected from the 1JY Catalogue by Kühr and Schmidt (1990).
- ◉ Flux $> 1\text{Jy}$ at 5 GHz.
- ◉ Lack of line emission in the optical band.
- ◉ Sample includes BL-Lac and many sources are part of MOJAVE.

Observations

- 16-17 January 2004
 - V.L.B.A. Data
 - 4 wavelengths between 18 and 22 cm
- 13 December 1999
 - V.L.B.A. Data
 - 18 and 13 cm
 - Good UV coverage
 - Data calibrated and imaged using A.I.P.S.

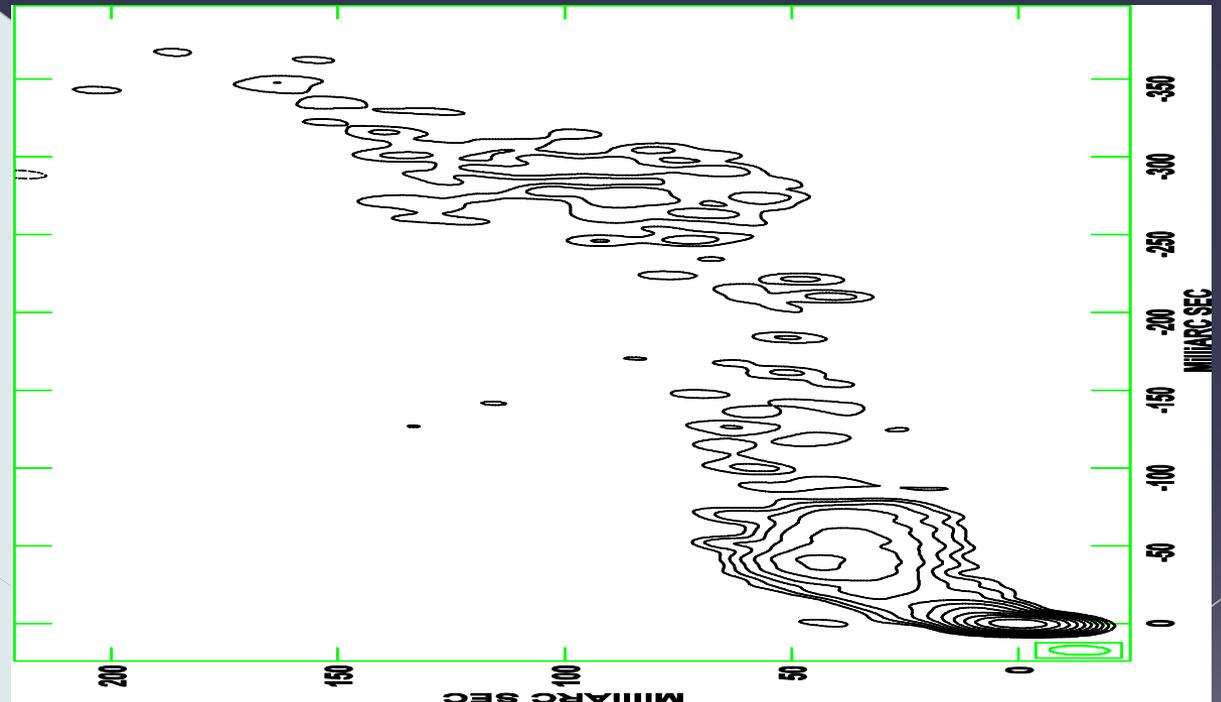


General Trends

- ◎ Morphology
 - › Jet extends from tens to hundreds of mas
 - › Jet emission out to 0.7 kpc
- ◎ Polarization Structure
 - › Interested primarily in the jet polarization
 - › Core lacks resolution
- ◎ Rotation Measure distribution
 - › Electron density and B field along line of sight
- ◎ Evidence for helical magnetic fields

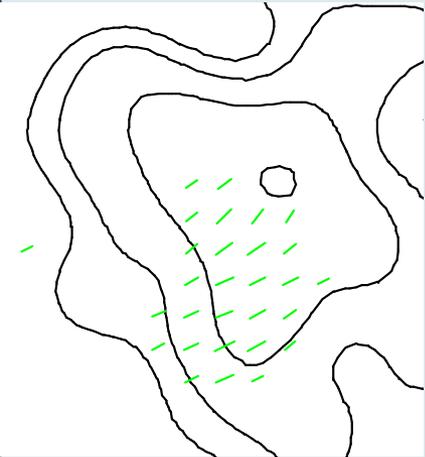
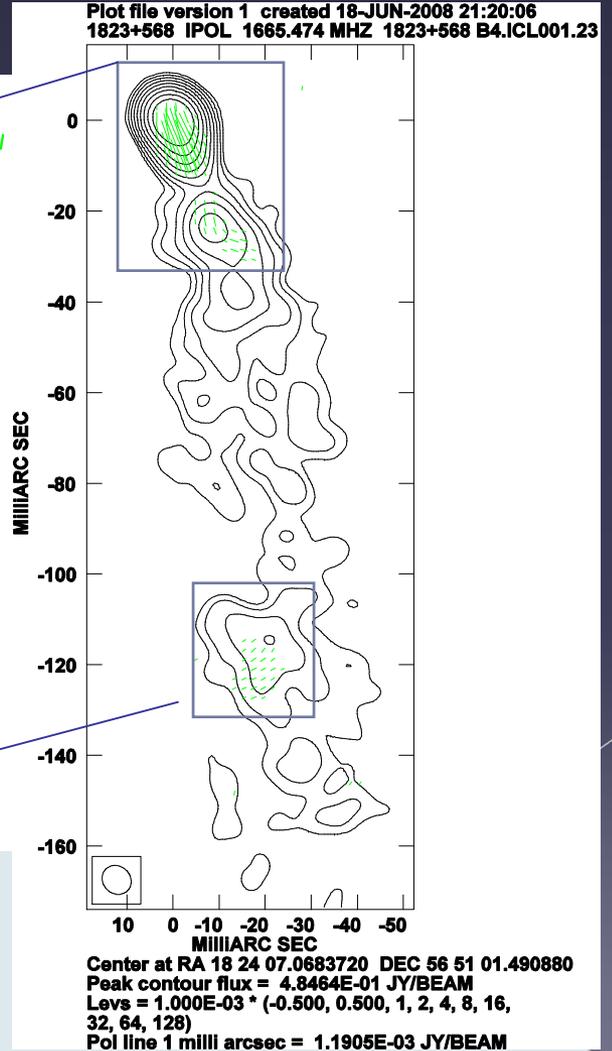
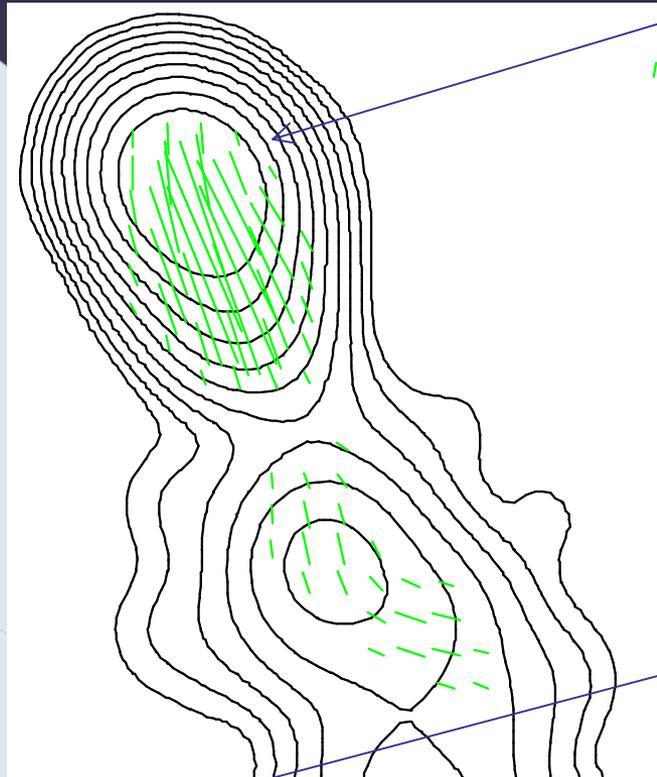
Large scale structure

- ◎ 1538+149
 - › Extended
 - › Signs of Jet instability



Large scale structure

- 1823+568
 - > Extended
 - > Polarized knot in Jet

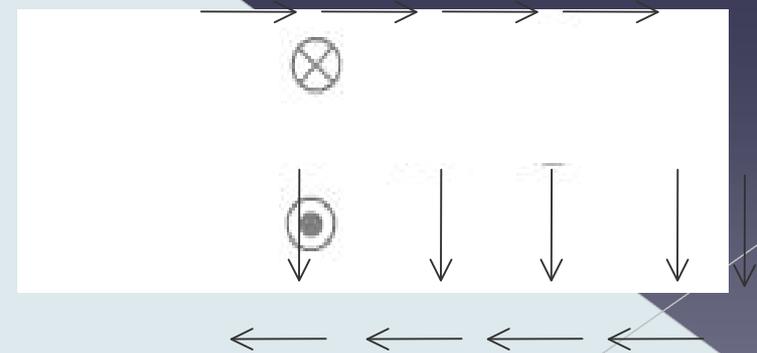


Evidence for Spine-Sheath structure

- What is it?
 - Observed tendency for E field to align with Jet direction (Parallel and Perpendicularly)
- Optical Depth
 - E relative to B
- Implications
 - Helical B Field

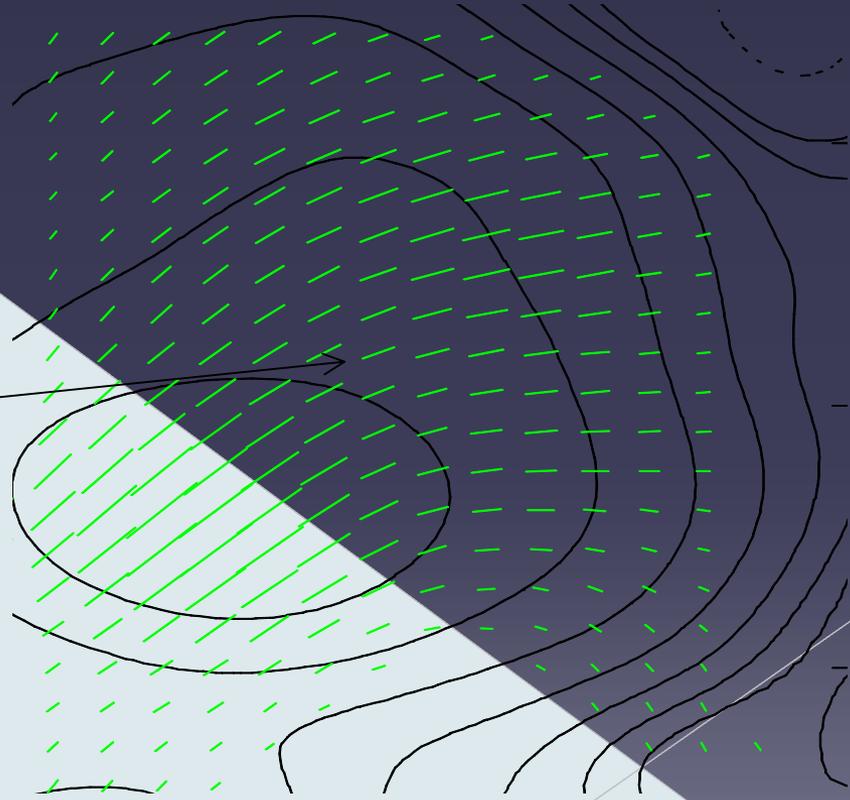
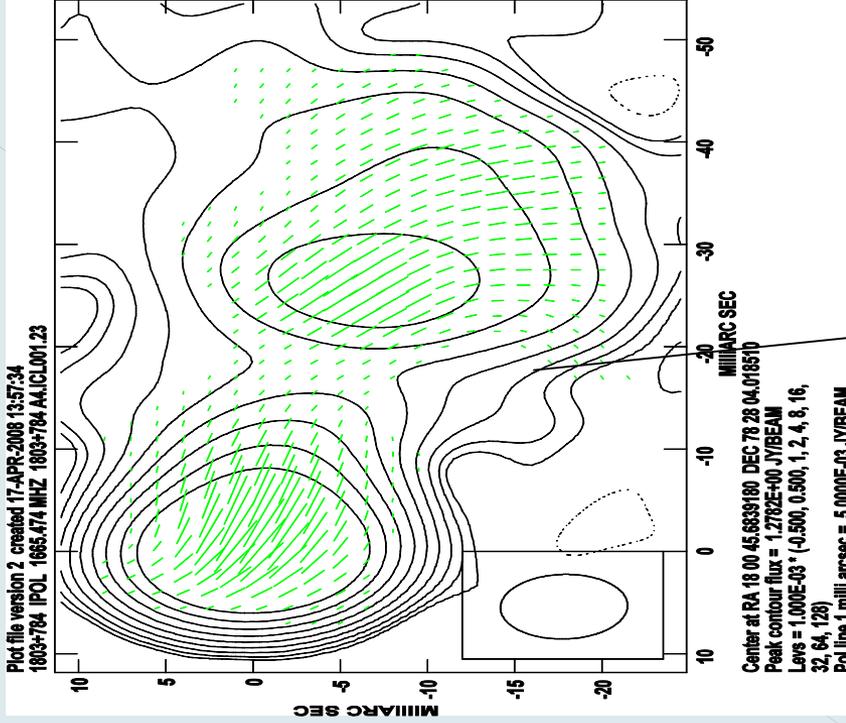


Below: Magnetic Field



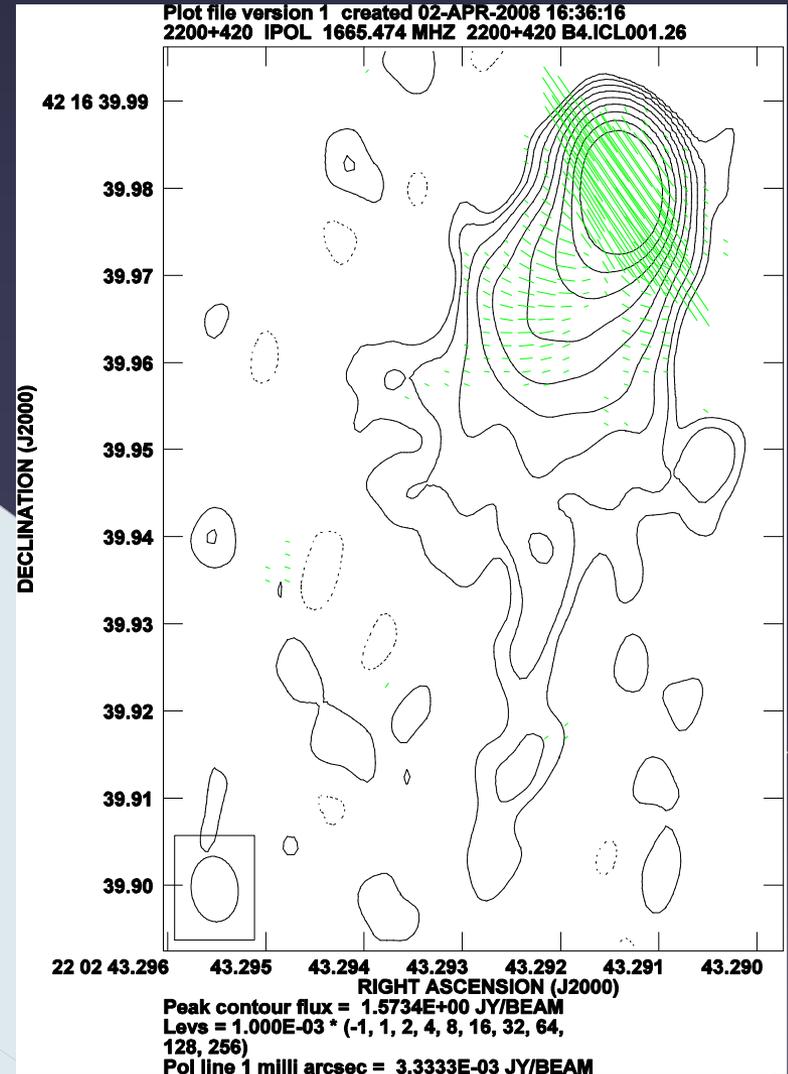
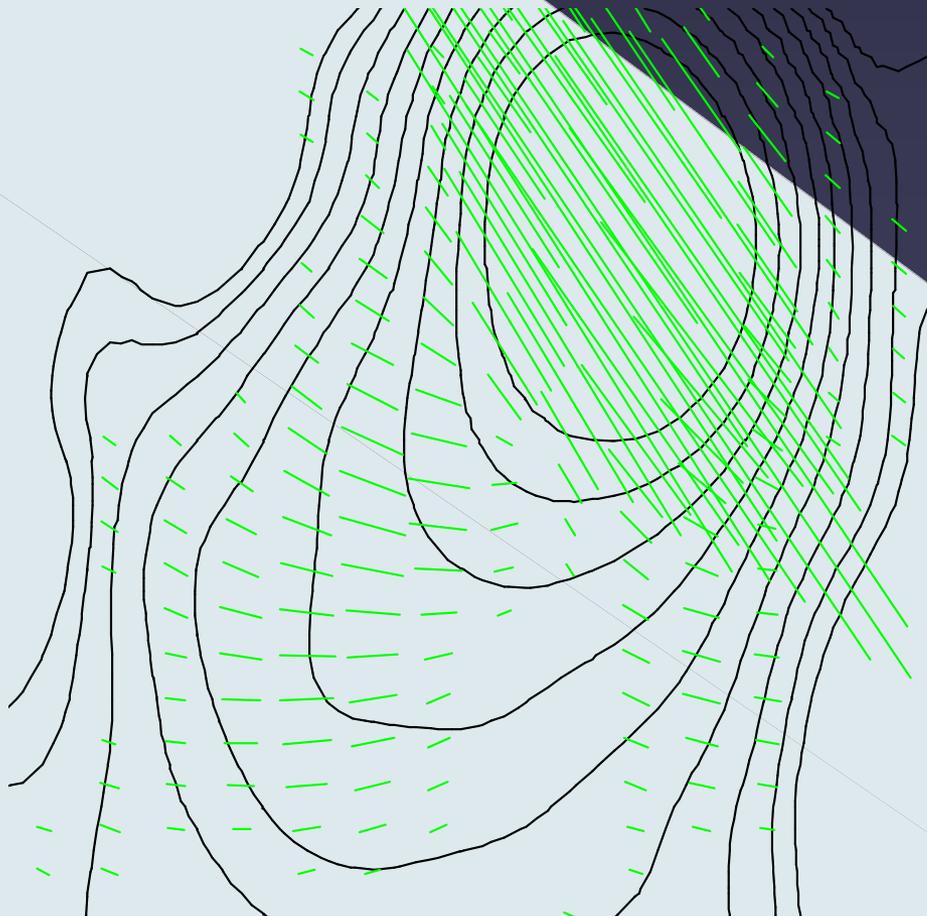
Polarisation structure

⊙ 1803+784



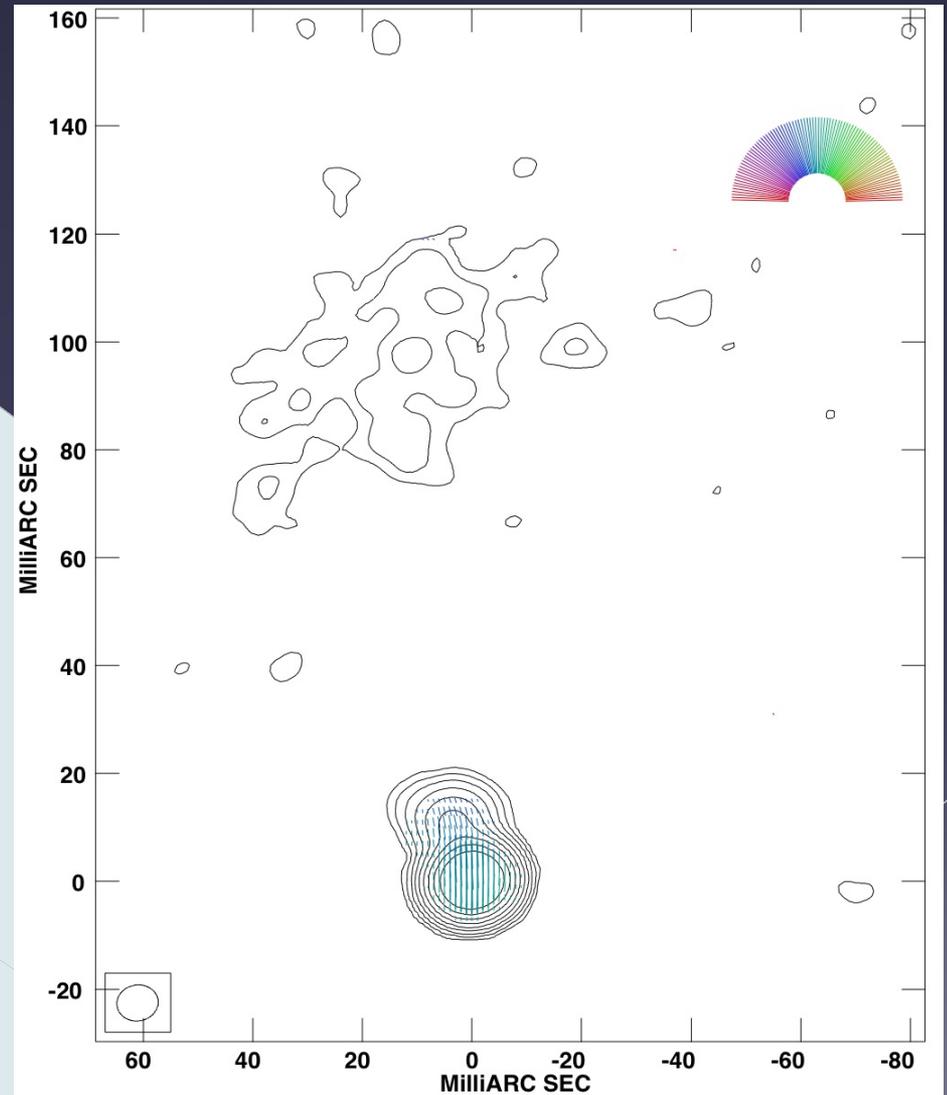
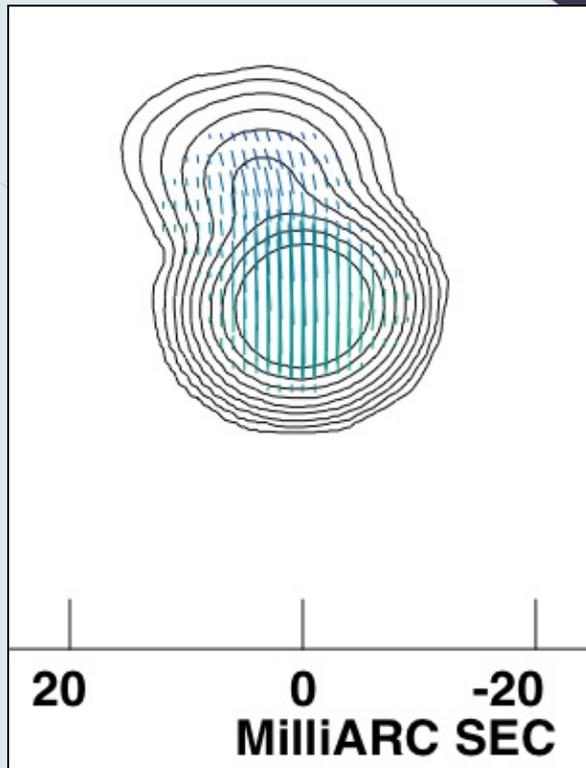
Polarisation structure

■ BL-Lac



Polarisation structure

⊙ 0716 + 714
> 2004 18cm

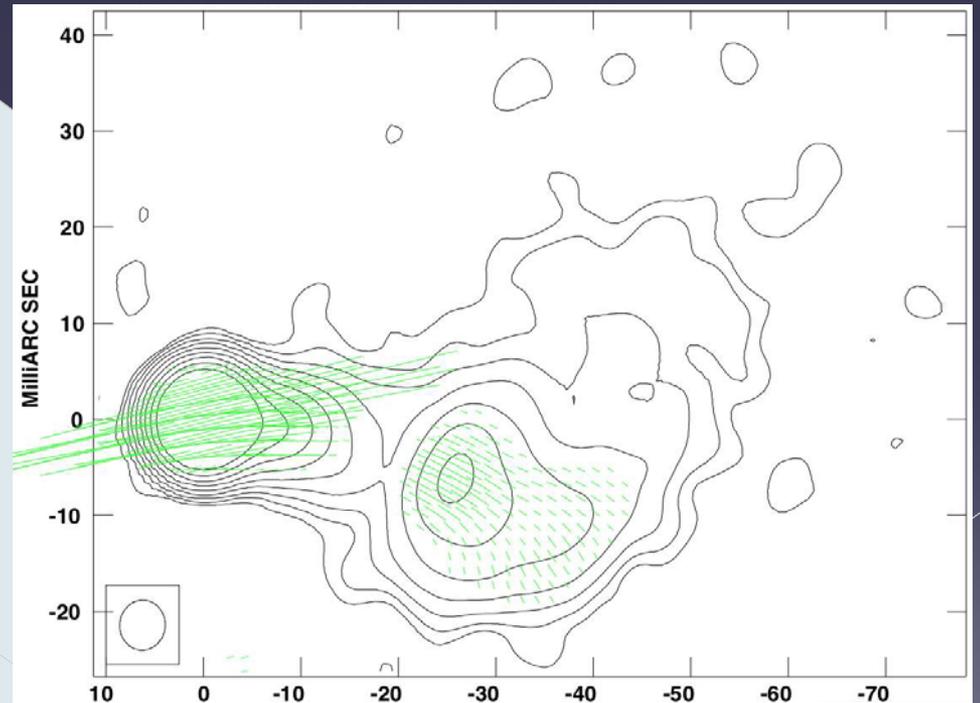
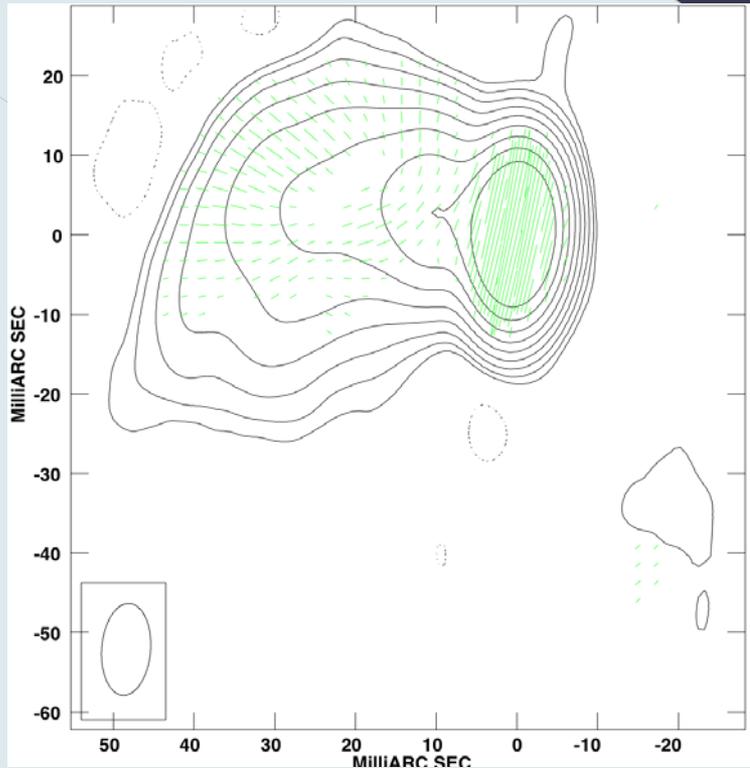


Multi-Epoch Observations

⊙ 0735+178

⊙ 13 cm, 1999

⊙ 1803+784



Faraday Rotation

- Rotation of the plane of linear polarization.
- RCP and LCP have different velocity in the plasma.

$$\chi_{obs} = \chi_0 + RM\lambda^2$$

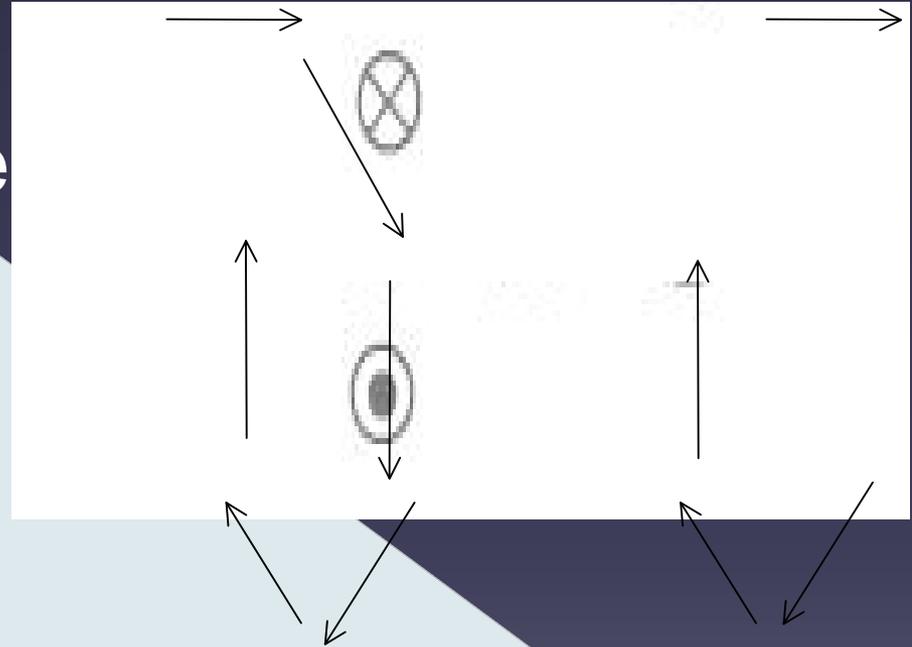
- Lambda squared dependence
- Dependant on magnetic field and free electron population.

Faraday Rotation effects

- Due to ISM in our own galaxy
- Pushkarev (2001)
- VLA observations (18 – 22 cm)
- Lambda squared dependence
- An RM of just 20 can cause rotations of up to 50 °

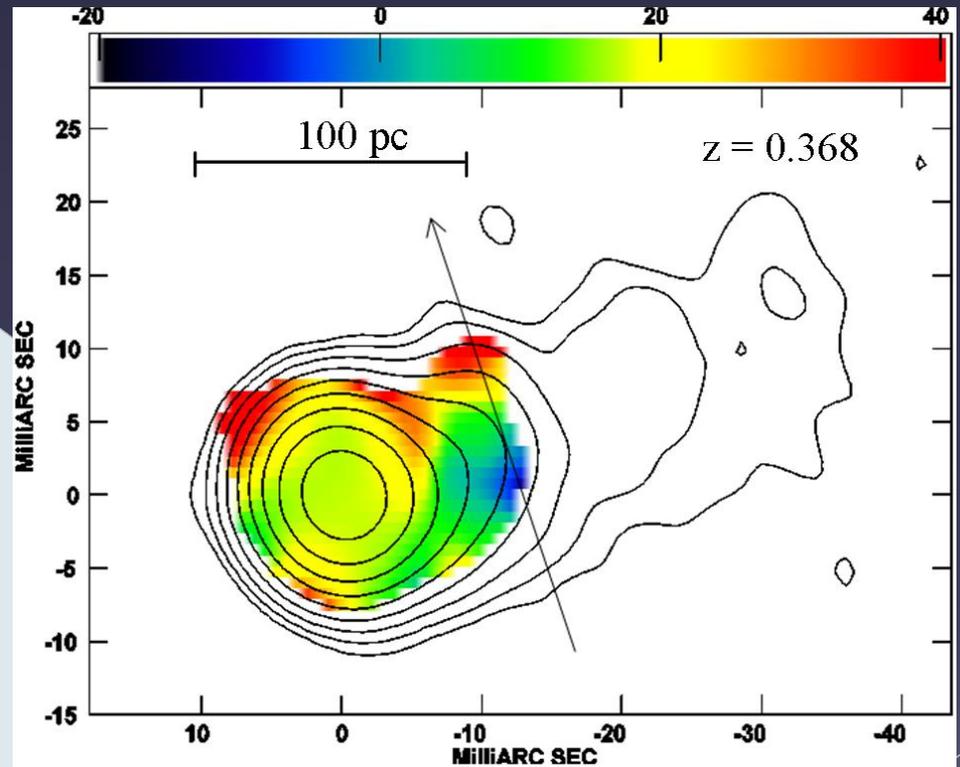
Rotation Measure Gradients

- Gradients due to the changing line of sight B field

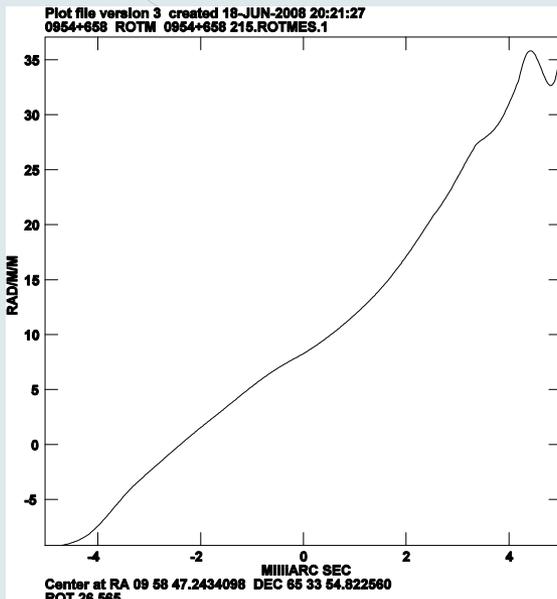


RM Intrinsic to the source

- ⊙ 0954+658
 - > Observed at other VLBA frequencies too.

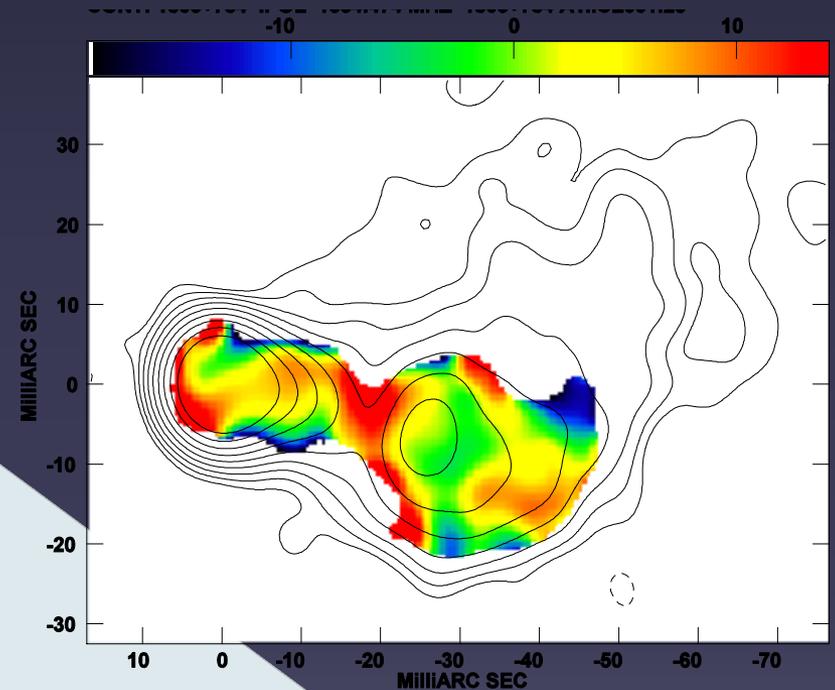


Above: 2004, 18 - 22cm RM
Left: Slice as shown by arrow
RM from -9 to +35 Rad / m²



RM Intrinsic to the source

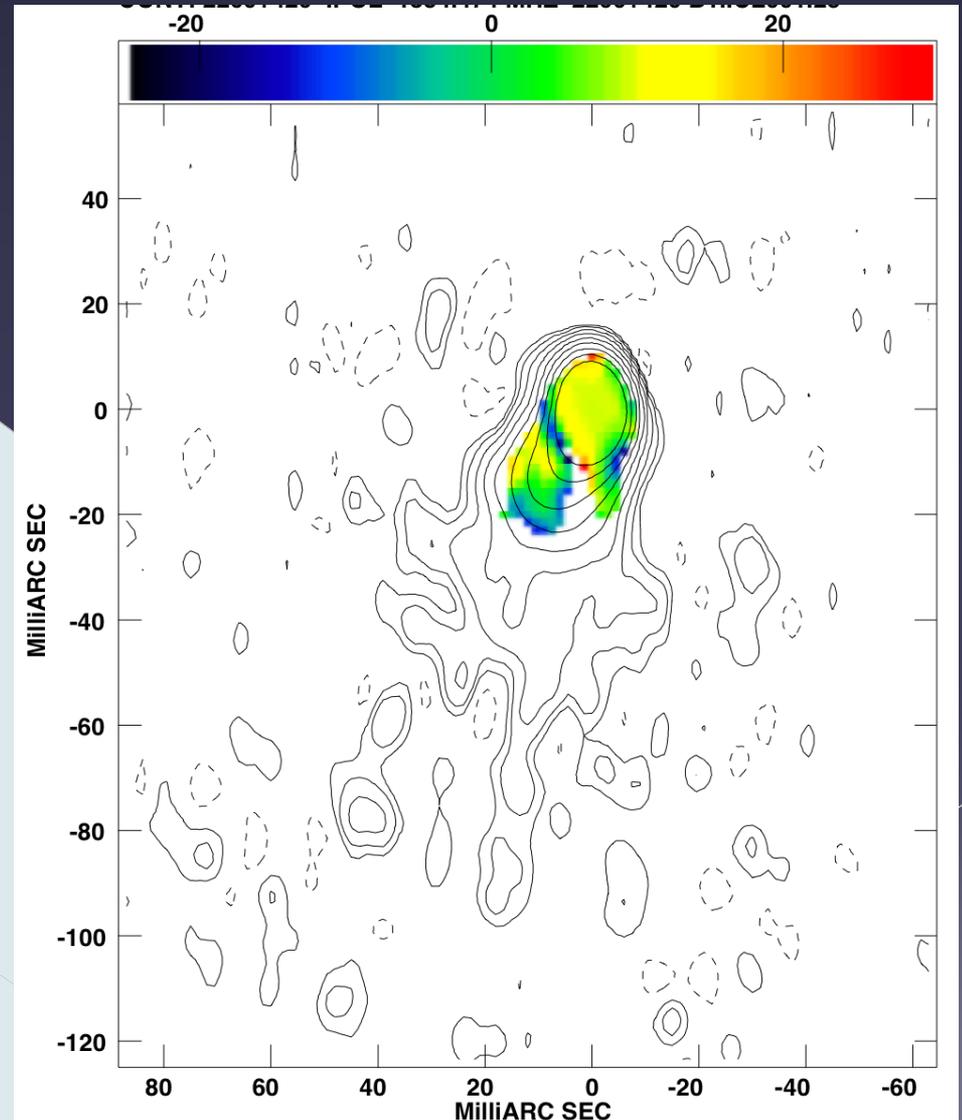
- 1803+784
 - Also observed at other VLBA frequencies.
 - Plenty of structure
 - Possible gradients
 - Needs further investigation and error analysis



RM Intrinsic to the source

⊙ BI Lac

- › Weak gradient across inner jet.
- › Modest faraday rotation from ISM of our galaxy, but weak RM at the source

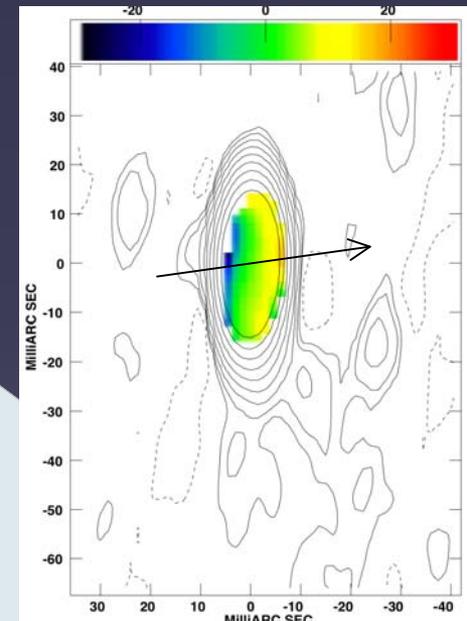
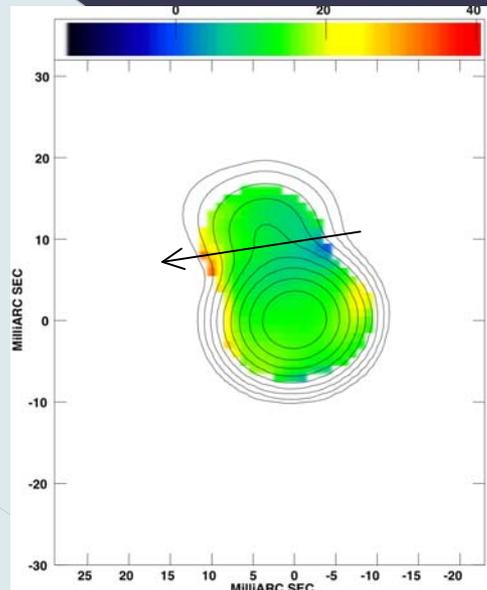
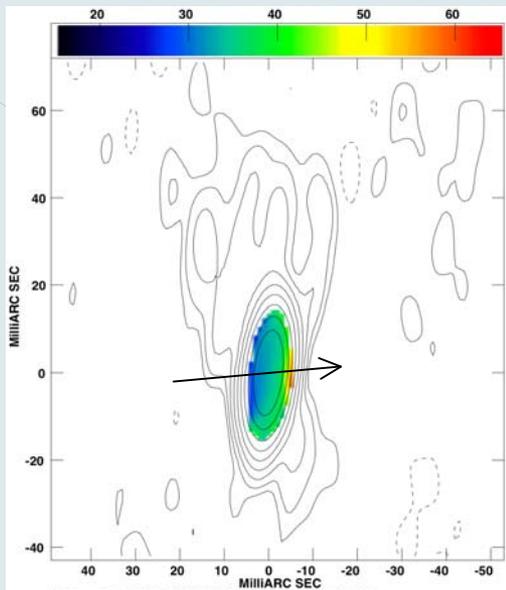


RM Intrinsic to the source

■ 0048-097

■ 0716+714

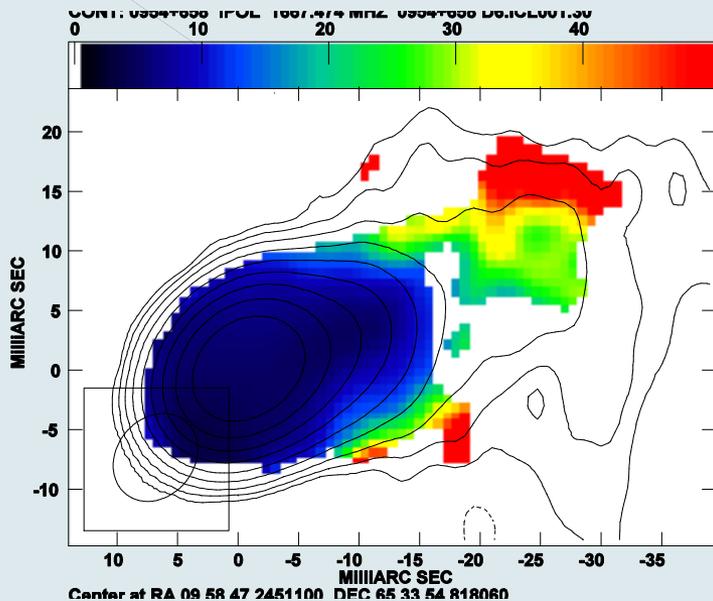
■ 2155-152



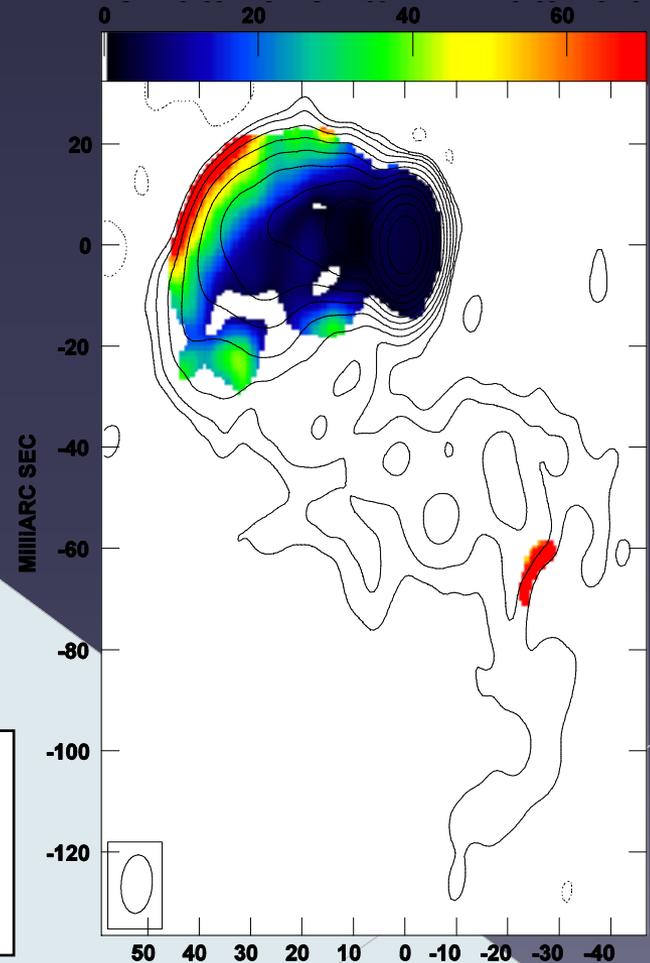
■ Both 0716+714 and 2155-152 gradients are in same direction as higher frequency observations of Mahmud

Fractional Polarisation Maps

- Further evidence for helical fields, cf Lyutikov (2005).



Left: 2004, 18 cm
0954+658
Right: 1999, 18 cm
0735+178



Project Status

- Imaging just completed for all data.
- Rotation measure maps needed for all sources from 2004 data.
- Rotation measure maps between 13 and 18cm from 1999 possible
- Categorise sources due to morphologies or polarisation structure.

Summary

- ◉ Decaparsec scale observations
- ◉ Rich structure
- ◉ Spine - sheath polarisation structures
- ◉ RM gradients
- ◉ Evidence for Helical Magnetic fields
 - › Ordered polarization structures



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polarisation observations of
34 BL-Lac Objects

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