





New morphologies of 6.7 GHz methanol masers

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Introduction:

Methanol maser emission at 6668.518 MHz:

- 1) traces the early stages of massive star formation,
- 2) probes the environment at small scales (~5 AU)
- very close to the protostar (~1000 AU),
- 3) shows different morphologies.

Our aim: To image the homogeneous sample of 31 masers discovered in Torun blind survey and positioned with MERLIN (Cm-Mk) and find out where and when the emission arises (increased sensitivity at EVN). **Observations:** phase-referenced, bandwidth 2MHz, 33 maser sites (+2 - Walsh et al. 1998, Beuther et al. 2002)

Run	Date	Durations	Telescopes
1	Jun 2003	12 hrs	4
2	Nov 2004	12	8
3a	Feb 2006	10	
3 b	Feb 2006	10	7
4 a	Jun 2007	10	9
4 b	Jun 2007	10	9
4 c	Jun 2007	10	8

Results: rms=3-12 mJy/beam, beam ~5mas x16mas

- 1) imaged 31 out of 33 masers,
- 2) known absolute positions towards 29 sources,
- 3) 1936 maser spots registered,
- 4) 318 maser clusters: 265 (83%) with Gaussian profile: mean FWHM=0.41 (+/-0.01) km/s.

Morphology ->

3 (9.5%) masers: simple



3 (9.5%) masers: linear





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1 (3%) maser: pair



12 (39%) masers: elliptical



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Model: rotating & expanding ring (Uscanga et al. 2008, Torstensson *soon*)



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Model & data: G33.980-00.019



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NAME	Vrot	Vexp	Vsys	i(deg)
G23.207-00.377	1.4	5.8	79.5	-69.1
G23.389+00.185	-1.0	2.0	75.1	54.3
G23.657+00.127	7.4	4.2	81.6	-15.8
G24.634-00.324	6.2	1.0	39.5	-70.5
G25.411+00.105	0.0	1.6	95.8	47.2
G26.598-00.024	0.8	1.2	24.6	54.6
G28.817+00.365	-2.6	4.2	89.8	-70.0
G30.400-00.296	-2.8	0.6	101.3	-69.6
G31.047+00.356	-0.4	3.8	80.3	-60.5
G31.581+00.077	-1.0	1.6	98.5	-44.7
G33.980-00.019	0.2	4.2	62.0	61.6
G34.751-00.093	-0.4	2.8	51.3	-52.9



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Conclusions:

better sensitivity → new morphologies,
elliptical morphology - common,
ring (disc/torus) with rotation and outflow.